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A. Lisé-Pronovost, F. Salomon, J.-P. Goiran, G. St-Onge, A.I.R. Herries, J.-C. Montero-Serrano, D. Heslop, A.P. Roberts, V. Levchenko, A. Zawadzki, H. Heijnis



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# Dredging and Canal Gate Technologies in Portus, the Ancient Harbour of Rome, Reconstructed from Event Stratigraphy and Multi-Proxy Sediment Analysis

Lisé-Pronovost, A.<sup>1</sup>, Salomon, F.<sup>2,3</sup>, Goiran, J.-P.<sup>4</sup>, St-Onge, G.<sup>5</sup>, Herries, A. I. R.<sup>1</sup>, Montero-Serrano, J.-C.<sup>5</sup>, Heslop, D.<sup>6</sup>, Roberts, A. P.<sup>6</sup>, Levchenko, V.<sup>7</sup>, Zawadzki, A.<sup>7</sup>, Heijnis, H.<sup>7</sup>

<sup>1</sup>The Australian Archaeomagnetism Laboratory, Palaeoscience Labs, Department of Archaeology and History, La Trobe University, Melbourne Campus, Bundoora, 3086, VIC, Australia

<sup>2</sup>Laboratoire Image Ville Environnement UMR 7362, Université de Strasbourg, 3 rue de l'Argonne 67083 Strasbourg cedex, France

<sup>3</sup>Department of Archaeology, University of Southampton, Southampton SO17 1BF, United Kingdom

<sup>4</sup>Centre national de la recherche scientifique, Maison de l'Orient et de la Méditerranée, Lyon 69365, France

<sup>5</sup>Institut des sciences de la mer de Rimouski, Canada Research Chair in Marine Geology, Université du Québec à Rimouski and GEOTOP, Rimouski G5L 2Z9, Québec, Canada

<sup>6</sup>Research School of Earth Sciences, Australian National University, Canberra, 2601, ACT, Australia

<sup>7</sup>Australian Nuclear Science and Technology Organisation (ANSTO), Lucas Heights, 2234, NSW, Australia

## ABSTRACT

Ancient harbours are rich archives of human-environment interaction. However, dating harbour deposits and correlating their stratigraphy is a major challenge because of typically high sedimentation rates over short periods and possible curative dredging events. Portus, the maritime harbour of Rome at the height of the Roman Empire, was a port complex composed of basins and canals connecting the commercial harbour to Rome via the Tiber River. Sediment core CPS1 in the narrowest of these canals, Canale Traverso, is located centrally in what was the capital city's commercial hub and contains a continuous harbor depositional record with average sedimentation rates greater than 1 cm per year. Here we use piston coring, high-resolution core scanning and a multi-proxy sediment analysis including for the first time paleo- and rock-magnetism, and bulk and clay mineralogy in order to overcome the problems of dating harbour deposits and correlating their stratigraphy. The method allowed precise identification of major reworked events, including a dredged deposit and a hyperpycnal deposit, which improve the chronostratigraphy and water depth reconstruction, and sheds light on harbour technologies at the height of the Roman Empire. A debris layer with abundant ceramic fragments and rocks marks the decommissioning of Canale Traverso and provides a new chronostratigraphic marker at Portus. Multi-proxy riverine input signatures point to the possible use of canal gate technology for water flow management.

1 1. INTRODUCTION  
2

3 Ancient harbour deposits around the Mediterranean Sea have high accumulation rates of up to one  
4 to four centimeters per year (Goiran and Morhange, 2001; Marriner and Morhange, 2006; 2007;  
5 Morhange and Marriner, 2010; Goiran et al., 2010; Salomon et al., 2012), which is orders of  
6 magnitude higher than in most natural environments such as lakes and oceans (Sadler, 1981). As  
7 such, ancient harbour sediment archives can potentially provide high-resolution sedimentary  
8 records. With time resolution comparable to varved lake sediment and ice core palaeoclimate  
9 records, ancient harbour archives also have the advantage of abundant historical reports. The high  
10 time resolution of ancient harbour archives is of interest across disciplines for documenting human-  
11 environment interaction, palaeoclimate and history, as well as geomagnetic field change. Such  
12 records could, for example, allow further investigation of hypothesised and debated links between  
13 climate stability and Roman Empire expansion and contraction (e.g., Büntgen et al., 2011;  
14 McCormick et al., 2012; Mensing et al., 2015).

15 Dating is a major challenge that impedes most interpretations of ancient harbour archives. It is  
16 challenging because of the large uncertainty of radiocarbon dating and calibration relative to the  
17 short time periods covered by ancient deposits, which typically do not exceed several hundred years  
18 (e.g., Goiran and Morhange, 2001; Salomon et al., 2012). Dating also depends on the available  
19 artefacts, if any, found within the stratigraphy, such as ceramic fragments that can be associated  
20 with a specific period. Most importantly, possible dredging of ancient harbour infrastructure  
21 significantly complicates attempts to build precise chronologies (Salomon et al., 2016). Historical  
22 reports and geoarchaeological evidence indicate that mechanical mud removal was performed to  
23 maintain sufficient draught for boats to navigate waterways since the 4<sup>th</sup> century BC, and more  
24 extensively during the Roman Empire (Morhange and Marriner, 2010). Dredging boats unearthed in  
25 Marseille constitute direct archaeological evidence for that technology (Pomey, 1995). Further  
26 stratigraphic evidence also exists such as a cut-and-fill talus exposed during excavation in

27 Marseilles (Morhange et al., 2003), and wedged scars in volcanic tufa bedrock in the ancient  
28 harbour of Naples (Morhange and Marriner, 2010). In sediment cores, however, dredging is at best  
29 tentatively identified from radiocarbon date inversions and chronological hiatuses (e.g., Marriner  
30 and Morhange, 2006; Marriner et al., 2006; Salomon et al., 2012; Delile et al., 2014; Goiran et al.,  
31 2014; Salomon et al., 2016; Stock et al., 2016). Extreme storm, tsunami, and flood events can also  
32 cause chronological hiatuses and ‘instantaneous’ deposition (St-Onge et al., 2004; Goiran et al.,  
33 2014; Vött et al., 2014; Hadler et al., 2015; Rübke and Vött 2017).

34 Canal gates represent another technology that may have affected sedimentation, but for which  
35 evidence is elusive. Historical reports indicate that canal gates may have been used for flood  
36 control, navigation or protection (Plinius (nat. hist. III 9,53); Bockius, 2014), yet direct physical  
37 evidence is limited to rocky ancient Mediterranean harbours, where rock walls have man-made  
38 cuttings (Allen, 1853; McCann, 1979; Erol and Pirazzoli, 1992). Without precise identification of  
39 dredging events, extreme weather events, canal gate use, and associated chronological hiatuses  
40 and/or reworked deposits within a sedimentary sequence, a precise age-depth model cannot be  
41 obtained. Precise identification of event deposits in lake and marine sediment archives is commonly  
42 achieved using piston coring to recover the most undisturbed stratigraphy, and high-resolution  
43 sediment analysis such as core scanning and magnetic properties (e.g., Duchesne et al., 2006;  
44 Storen et al., 2010; Gilli et al., 2013). There are no universal criteria to identify what caused event  
45 deposits, and consideration of the local setting and a multi-proxy approach are important. Expected  
46 sedimentological evidence for dredging includes sharp contacts, sedimentary discontinuities, and  
47 age hiatuses from removal of dredged sediment, and possibly disturbed stratigraphy and mixing of  
48 sediment layers from shoveling. In contrast, hydrodynamic high-energy storm, tsunami, flood, and  
49 canal gate control events would likely produce graded beds from changes in flow intensity and  
50 sediment transport capacity. Event deposits must be taken into account to build robust age-depth  
51 models.

52 Portus, the ancient harbour of Rome, is one of the best-studied ancient harbour sites (Keay et al.,  
53 2005; Keay and Paroli, 2011). The Emperor Claudius initiated Portus' construction in AD 42 near  
54 the Tiber River mouth on the Tyrrhenian Sea, with a second phase of development under the  
55 Emperor Trajan (AD 112-117) (Keay et al., 2005), which included construction of the emblematic  
56 hexagonal basin that is visible today (Fig. 1). Portus was a large harbour complex that served as the  
57 commercial hub and store for Rome at the height of the Roman Empire, receiving goods from all  
58 around the Mediterranean Sea. Today, remains of Portus are landlocked three kilometers from the  
59 coastline, with part of the port complex lying under the Leonardo da Vinci-Fiumicino International  
60 Airport. With exceptionally high sedimentation rates in its basins and canals (Giraudi et al., 2009;  
61 Goiran et al., 2010; Salomon et al., 2012; 2014, Delile et al., 2014), the sedimentary archive of  
62 Portus has the potential to provide high-resolution geoarchaeological, palaeoclimatic and  
63 paleomagnetic records. No previous study has provided precise chronologies because of: 1) the  
64 disturbed stratigraphy of cores that were drilled using rotating core barrels and extraction of  
65 sediment into plastic trays; 2) possibly unidentified hiatuses and/or event deposits; and 3) large  
66 radiocarbon dating and calibration uncertainties relative to the short period covered by the harbour  
67 deposit. As a result, age-depth models from Portus are mostly vertical with several meters of mud  
68 having the same age, and the stratigraphy of cores is difficult to correlate to each other. At Portus,  
69 sedimentary radiocarbon date inversions provide hints of dredging (Delile et al., 2014; Salomon et  
70 al., 2012; Goiran et al., 2010) and dredging in Portus basins is depicted in an epigraph at ca AD 400  
71 (Coccia, 1993; Giraudi et al., 2009). Flood gates or canal locks were also hypothesised on the canals  
72 of Portus, including the Canale Traverso (Testaguzza, 1970), the Fossa Traiana, and the Canale  
73 Romano (Salomon et al., 2014); however, there has been no conclusive evidence for these  
74 technologies. Portus was clearly influenced by fluvial input however, no flood layer has been  
75 identified (Salomon et al., 2012; 2014). Flood and/or storm/tsunami deposits have been reported in  
76 the harbour of Ostia, also located on the Tiber delta, in the centuries preceding the construction of  
77 Portus (Goiran et al, 2014; Hadler et al, 2015; Marriner et al, 2017). Here we apply methods

78 commonly used in marine geology, including piston coring, high-resolution core scanning, and  
79 multi-proxy sediment analyses (grain size, paleo- and rock-magnetism, and bulk and clay  
80 mineralogy), together with radiocarbon dating and Bayesian age modelling to precisely identify  
81 event deposits and to better constrain the Portus chronostratigraphy, and to document Roman  
82 harbour technologies.

## 83 2. MATERIAL AND METHODS

### 84 2.1 Field Work

85 Coring operations were completed in September 2011 using a stationary hydraulic piston system  
86 operated by the *Centre d'Étude Techniques de l'Équipement (CETE) Méditerranée*. This method  
87 allows recovery of relatively undisturbed and continuous sediment sequences with core sections  
88 preserved in polyvinyl chloride (PVC) tube liners. The cores were sealed and shipped to the *Institut*  
89 *des sciences de la mer de Rimouski (ISMER)* of the *Université du Québec à Rimouski (UQAR)* for  
90 analysis. Based on the high sediment accumulation rates (Salomon et al., 2012) and the results of a  
91 pilot rock-magnetic study (cf. section 2.2) it was decided to core Canale Traverso. The coring site in  
92 Canale Traverso was selected near the access channel (Fig. 1B) based on the hypothesis that the  
93 hydraulic energy would dissipate and, thus, sediment would accumulate most rapidly where canal  
94 widens. Two cores, CPS1 and CPS4, were recovered from that location (Figs. 1B, D). CPS1 was  
95 selected for high-resolution multi-proxy study because of its thicker apparent harbour deposit.

### 96 2.2 Pilot Study

97 Samples from previous rotary drilling at Portus (Goiran et al., 2010; Salomon et al., 2012) were  
98 analysed at ISMER prior to field work to evaluate the potential for paleo- and environmental  
99 magnetism studies at Portus. Eight sediment samples were chosen from core TR-20 in the access  
100 channel and seven samples from core CT1 in Canale Traverso (see core locations in Fig. 1B). The  
101 performed analyses include temperature dependence of magnetic susceptibility (50 to 700 °C),  
102 alternating field (AF) demagnetisation of the natural remanent magnetisation (NRM), hysteresis  
103 properties, and isothermal remanent magnetisation (IRM) acquisition (methods summarized in Lisé-

104 Pronovost et al., 2013). Results of the pilot study (supplementary material 1) indicate that Canale  
105 Traverso was the most promising site with a strong magnetic signal dominated by low coercivity  
106 magnetite with single domain (SD) to vortex magnetic state, and a stable NRM, which are suitable  
107 for paleomagnetic recording (Tauxe, 1993). By comparison, sediment from the access channel (TR-  
108 20) has a weak magnetic signal likely associated with dissolution of detrital ferrimagnetic particles  
109 in less oxygenated and more organic-rich sediment (e.g., Roberts, 2015), because the site is located  
110 further away from fluvial input to the harbour.

### 111 2.3 High-Resolution Analyses

112 Magnetic, physical, and mineralogical analyses of core CPS1 were undertaken at high-resolution on  
113 whole core sections, on half-core split surfaces, on continuous u-channels (2 cm x 2 cm x core  
114 length), and on a series of discrete samples (Table 1). The discrete samples are a series of 50  
115 standard 8 cm<sup>3</sup> cubes (2 cm x 2 cm x 2 cm) that were used for magnetic analysis and ca 5 g of  
116 sediment that was used for mineralogical analysis, taken at ca 10 cm intervals down-core (see  
117 sampling strategy log in Fig. 2). This discrete sampling interval is comparable to the highest  
118 sampling resolution previously achieved in Portus.

#### 119 2.3.1 Physical Properties

120 CT-scan images and gamma ray attenuation provide indications of sediment density. They are  
121 measured rapidly at high-resolution and are non-destructive (St-Onge et al., 2007; Fortin et al.,  
122 2013). CT-scan images of whole cores was obtained at the *Institut national de la recherche*  
123 *scientifique Eau-Terre-Environnements* (INRS-ETE) using a Siemens SOMATOM Definition sliding  
124 gantry CT-Scanner. Tomograms were acquired continuously with a pixel size of 0.06 cm and were  
125 then transferred into digital format using a standard Hounsfield scale (Hounsfield, 1973). Higher  
126 density and higher atomic numbers generate greater X-ray attenuation, which is represented by  
127 higher CT number and lighter tones on a CT scan image. Gamma ray attenuation was measured on  
128 whole cores at 1-cm intervals using a Cesium-137 gamma ray source and detector mounted on a

129 multi-sensor core logger (MSCL; Geotek Ltd.) at ISMER, UQAR. Imaging of half-core split  
130 surfaces was done using a linescan imaging Geoscan IV (Geotek Ltd) instrument.

### 131 2.3.2 Magnetic Properties

132 Magnetic mineral assemblages in the sediment were characterised using room-temperature  
133 magnetic properties, including the magnetic susceptibility, NRM and laboratory-induced  
134 magnetisations, and coercivity-dependent parameters. Magnetic susceptibility primarily reflects  
135 ferrimagnetic mineral concentration and is influenced by sediment composition, magnetic  
136 mineralogy, and grain size. The volume magnetic susceptibility ( $\kappa$ ) was measured on the whole  
137 cores using a Bartington Instruments MS2C loop sensor at 1-cm stratigraphic intervals, and then on  
138 the half-core split surface with a MS2E high-sensitivity point sensor at 0.5-cm stratigraphic  
139 intervals at ISMER. Frequency-dependent magnetic susceptibility ( $\chi_{FD\%}$ ) of cube samples was  
140 measured using a MS2B dual frequency sensor at The Australian Archaeomagnetism Laboratory  
141 (TAAL), La Trobe University.  $\chi_{FD\%}$  is defined as  $\chi_{FD\%} = (\chi_{LF} - \chi_{HF}) / \chi_{LF} \times 100$ , with low and high  
142 frequencies (LF and HF) of 0.46 and 4.6 kHz, respectively (Dearing, 1999).  $\chi_{FD\%}$  is controlled by  
143 the grain size distribution of superparamagnetic (SP) and stable SD particles (Eyre, 1997).  
144 Measurements were repeated at least six times; average values and standard errors are reported.

145 The NRM and ARM were measured and stepwise demagnetised on u-channels from 0 to 100 mT in  
146 13 and 5 steps, respectively, using a 2-G Enterprises cryogenic magnetometer for u-channels at  
147 ISMER. ARM was imparted in a peak AF of 100 mT with a direct bias field of 0.05 mT.  $\kappa_{ARM}$  is the  
148 ARM normalised by the direct bias field. Principal component analysis (PCA) of NRM  
149 demagnetisation data was performed using the Puffin Plot software (Lurcock and Wilson, 2012).  
150 Four data points were masked at section ends to account for the width of the response function of  
151 the cryogenic magnetometer pick-up coils (Weeks et al., 1993) and data from intervals with  
152 incompletely filled u-channels or gaps were removed. The median destructive field (MDF) is the  
153 AF required to demagnetise half of the initial remanence. MDF is a coercivity indicator and is a  
154 useful grain size indicator when the magnetic mineral assemblage is uniform (Dankers, 1981;

155 Dunlop and Ozdemir, 1997). Like  $\kappa$ , ARM is controlled mainly by the concentration of  
156 ferrimagnetic particles. In addition, SD grains more easily acquire ARM than coarser multi-domain  
157 (MD) grains (Maher, 1988; Evans and Heller, 2003), so ARM is also a useful magnetic grain size  
158 indicator (i.e.,  $\kappa_{ARM}/\kappa$ ,  $MDF_{ARM}$ ).

159 Hysteresis loops, IRM acquisition curves, and backfield demagnetisation curves were measured for  
160 each sample in a maximum field of 1 T using a Princeton Measurement Corp. vibrating sample  
161 magnetometer at the Australian National University (ANU). The bulk coercive force ( $H_c$ ), the  
162 remanent coercive force ( $H_{cr}$ ), the saturation magnetisation ( $M_s$ ), and the saturation remanence ( $M_r$ )  
163 were obtained and the ratios  $M_r/M_s$  and  $H_{cr}/H_c$  are used as magnetic grain size indicators (Day et al.,  
164 1977). The hard isothermal remanent magnetisation ( $HIRM = (IRM_{1T} + IRM_{0.3T})/2$ ; King and  
165 Channell, 1991) and the S-ratio ( $S\text{-ratio} = -IRM_{-300mT}/SIRM$ ; Stober and Thompson, 1979) were  
166 calculated; they reflect the absolute concentration of high coercivity magnetic minerals and the  
167 relative concentration of low to high coercivity minerals, respectively (Liu et al., 2012).  $B_{1/2}$  is half  
168 the applied field required for a sample to reach saturation and reflects the ease with which a sample  
169 is magnetised; it is a coercivity indicator. The 50 measured IRM acquisition curves were unmixed  
170 using an unsupervised approach (Heslop and Dillon, 2007) to identify and characterise the end-  
171 members that represent magnetic grain populations. Finally, first-order reversal curve (FORC;  
172 Roberts et al., 2000) diagrams for seven samples (CPS1-3-7, 4-5, 4-11, 5-5, 5-8, 6-8, and 7-3) were  
173 obtained using the same instrument to further explore the magnetic coercivity spectrum, domain  
174 structure, and magnetic interactions of different sediment types.

### 175 2.3.3 Granulometry

176 After completion of magnetic analyses, sediments in cube samples were air-dried at room  
177 temperature. To prepare samples for grain size analysis, dried sediments were sieved through a 2-  
178 mm mesh and were gently homogenised using an agate mortar and pestle. Subsamples of about 2 g  
179 were treated repeatedly with hydrogen peroxide ( $H_2O_2$ ) and gentle heating to remove organic  
180 matter, until peroxide addition to the sediment did not react further. Finally, sediment dispersion of

181 the treated and rinsed samples was achieved by adding 1 mL of 0.5M sodium hexametaphosphate  
182 ( $\text{NaPO}_3$ ) and ultrasonicated for 15 minutes. Grain size analyses were performed using a Malvern  
183 Mastersizer 2000 laser diffraction spectrophotometer at the Australian Nuclear Science and  
184 Technology Organisation (ANSTO). Each prepared sample underwent three successive 10-second  
185 measurement runs using continuous sonication to disperse aggregated particles. All statistical grain-  
186 size parameters were calculated with the GRADISTAT software (Blott and Pye, 2001) using the  
187 Folk and Ward graphical method (Folk and Ward, 1957). Logarithmic biplots of the median (D50)  
188 and 90<sup>th</sup> percentile diameter (D90) or the Passega diagram (Passega, 1964; Bravard and Peiry,  
189 1999) were used to estimate the hydrodynamic energy associated with sediment deposition from  
190 decantation to rolling.

#### 191 2.3.4 Bulk and Clay Mineralogy

192 Before mineralogical analysis, the detrital sediment fraction of the samples was isolated using 10  
193 mL of peroxide (30%  $\text{H}_2\text{O}_2$ ) and 10 mL of hydrochloric acid (0.5M HCl) to remove organic matter  
194 and biogenic carbonate, respectively. Deflocculation was achieved by successive washings with  
195 distilled water. Next, for bulk mineralogy, sediment samples were ground with a McCrone  
196 micronizing mill with agate grinding elements for 5-10 min with 10 mL of ethanol to obtain a  
197 consistent grain size of  $<10 \mu\text{m}$ . The slurry was oven dried overnight at  $60^\circ\text{C}$  and was slightly  
198 homogenised with an agate mortar to avoid agglomeration of finer particles during drying. Random  
199 powder samples were analysed by X-ray diffraction (XRD) using a PANalytical X'Pert Powder  
200 diffractometer (copper anode; 45 kV; 40 mA intensity). Samples were scanned from  $5^\circ$  to  $65^\circ$  two-  
201 theta in steps of  $0.020^\circ$  two-theta and a counting time of 2 seconds per step.

202 For semi-quantification of major mineralogical components, bulk sediment XRD scans were  
203 processed using the software package X'Pert High-Score Plus (PANalytical) with Rietveld full-  
204 pattern fitting. This method permits semi-quantification of whole-sediment mineralogy with a  
205 precision of 5–10% for phyllosilicates and 5% for grain minerals. Major mineralogical components  
206 quantified by this technique are quartz, K-feldspar, plagioclase feldspar, pyroxene, magnetite,

207 pyrite, and phyllosilicates. The sum of quartz, phyllosilicates, K-feldspar, and Na-plagioclase  
208 contents is a detritus index that reflects changes in detrital influx or terrigenous supply (e.g., Keller  
209 and Pardo, 2004; Mort et al., 2008; Montero-Serrano et al., 2015). The ratio of fine detrital material  
210 (phyllosilicates) to coarse detrital material (quartz, plagioclases, and K-feldspars) is denoted  
211  $Phy/(Qz+Pl+Feld)$  and indicates hydrodynamic sorting intensity.

212 Clay-mineral associations (smectite, illite, kaolinite, and chlorite) were studied using XRD  
213 following the protocols of Bout-Roumzeilles et al. (1999) and Montero-Serrano et al. (2009). The  
214 clay fraction was separated by settling according to Stokes's law, concentrated by centrifugation,  
215 and oriented by wet smearing on glass slides. Analyses were run from  $2.49^\circ$  to  $32.49^\circ$   $2\theta$  on a  
216 PANalytical X'Pert Powder diffractometer. Three runs were made after air-drying each sample with  
217 ethyleneglycol vapour saturation for 12 hours and heating at  $490^\circ\text{C}$  for 2 hours. Each clay mineral  
218 is characterised by its basal layer plus interlayer interval ( $d$ ) from XRD analysis (Brown and  
219 Brindley, 1980). Semi-quantitative clay mineral abundance estimation (smectite, illite, chlorite, and  
220 kaolinite), based on peak areas, was performed using the MacDiff® 4.2.5 software (Petschick,  
221 2000). The reproducibility of measurements is estimated to be 5% for each clay mineral, as checked  
222 by replicate sample analysis.

223

#### 224 2.3.5 Radiocarbon dating

225 Radiocarbon dating was performed using an Accelerator Mass Spectrometry (AMS) at ANSTO  
226 (Fink et al., 2004). Six samples of short-lived terrestrial material were measured, including three  
227 seeds, one coniferous bud, and two pollen/charcoal samples (Table 2). Measured ages were  
228 calibrated using the  $^{14}\text{C}$  calibration program CALIB 7.0 (<http://calib.qub.ac.uk/calib/calib.html>; 1.0  
229 (Stuiver et al., 2017) and the IntCal13 dataset (Reimer et al., 2013) and are stated using 2 sigma  
230 errors. The pollen extraction by density separation with lithium heteropolytungstate (LST) heavy  
231 liquid revealed abundant charcoal in the two pollen samples, and charcoal could not be separated  
232 from the pollen prior to dating. The dated carbon for samples OZS598 and OZS602 must, thus,

233 come from both pollen and charcoal with possibly different respective ages. This was taken into  
234 account using Bayesian age modelling with the Charcoal Outlier Model in OxCal v4.2.3 (Bronk  
235 Ramsey and Lee, 2013), with an overall agreement of 93%. This model considers sample types and  
236 for the two pollen samples that also incorporated charcoal, outliers to older times are allowed to  
237 account for possible older wood in charcoal. The same radiocarbon calibration and age modelling  
238 method was applied to previously published radiocarbon dates from Portus cores TR14 (Delile et  
239 al., 2014) and CT1 (Salomon et al., 2012).

## 240 3. RESULTS

### 241 3.1 Stratigraphy

242 Six sedimentary units, labelled A to F from base to top, are defined based on lithology from core  
243 descriptions, photos, CT-scanning, density (gamma-ray attenuation and CT-number), magnetic  
244 susceptibility, and physical grain size data (Fig. 2). These units are described as follows.

#### 245 3.1.1 Unit A

246 Unit A occurs at the base of the core (696 cm) to 573 cm and is composed of laminated yellow  
247 muddy sands characterised by relatively high density and magnetic susceptibility values (Fig. 2).  
248 Grain size analysis of four discrete samples indicate unimodal distributions and mean grain size  
249 ranging from 47 to 119  $\mu\text{m}$ . The upper 10 cm has fine laminations of sandy muds and muddy sands.  
250 A sharp oblique discontinuity cuts these laminations and marks the end of unit A. Unit A has  
251 relatively high quartz content (up to 45%), coarse detrital material (Phy/(Qz+Pl+Feld)), and  
252 virtually no pyrite or pyroxene (Fig. 3). The magnetic particle assemblage is dominated by low  
253 coercivity minerals, as indicated by saturation of the IRM below 300 mT, which results in S-ratio  
254 values close to one and low HIRM values (Fig. 3). The low-coercivity magnetic component EM1  
255 has the highest relative contribution in unit A, reaching 86% of the IRM (Fig. 3).

## 256 3.1.2 Unit B

257 Unit B is identified from depths of 573 to 480 cm and is composed of massive sands and muddy  
258 sands with cm-sized mud clasts. Immediately above the sharp oblique discontinuity at 573 cm is a  
259 38-cm-thick massive medium sand layer with unimodal grain size distribution (mode 340  $\mu\text{m}$ ;  
260 sample 7-1; Fig. 2E) that encompasses the coarsest mean grain size in the studied core CPS1 (291  
261  $\mu\text{m}$ ; sample 7-1; Fig. 2E). At 535 cm, a sharp  $\kappa$  drop of one order of magnitude and an increased  
262 mud content marks a change from sand to muddy sand (Figs. 2C-D). The muddy sand has a  
263 bimodal grain size distribution, with a first mode similar to the massive sand (341  $\mu\text{m}$ ) and a much  
264 finer second mode of 39  $\mu\text{m}$  (each mode is calculated from 4 samples). Sample 6-8 was taken from  
265 an individual mud clast within the muddy sands and is a distinctly finer medium silt with mean  
266 grain size of 8  $\mu\text{m}$  (Fig. 2E). Unit B is clearly composed of ungraded reworked sediments. This unit  
267 is capped with fine laminations of very coarse silts (mean grain size of 47  $\mu\text{m}$ ; sample 6-5). The  
268 mineralogical and magnetic properties in the massive sands of unit B are similar to these of unit A.  
269 In contrast, the muddy sands with mud clasts in unit B have relatively lower quartz content (30 to  
270 34%), finer detrital minerals ( $\text{Phy}/(\text{Qz}+\text{Pl}+\text{Feld})$ ), and higher pyrite (0.5 to 1.4%) contents (Fig. 4).  
271 The magnetic component EM2 appears in unit B and reaches maximum values in mud clast sample  
272 6-8 (Fig. 4).

## 273 3.1.3 Unit C

274 Unit C from 480 to 414 cm is composed of graded beds of muddy sands and sandy muds  
275 intercalated with a 30-cm-thick central layer (430-460 cm) of abundant cm-size rocks, ceramic  
276 fragments, wood, and other organic debris. A sharp change in gamma ray attenuation and CT-  
277 number at 480 cm (Fig. 2A-B) reflects the distinct densities of sedimentary units B and C. The grain  
278 size distributions of the six discrete samples from unit C reveal, from base to top, a coarsening  
279 upward sequence and a fining upward sequence. Similarly, the concentration of ferrimagnetic  
280 minerals as indicated by  $\kappa$ ,  $M_s$ , ARM, and NRM increases steeply from the base of unit C to the  
281 debris layer, and decreases gradually to the top of unit C. The three samples within the debris layer

282 have trimodal grain size distributions with the coarser mode reaching 590  $\mu\text{m}$  (Fig. 2E). In contrast,  
283 samples below and above the debris layer have bimodal distributions (224 and 6  $\mu\text{m}$ ). Mineralogical  
284 and magnetic properties have sharp and large amplitude changes in the debris layer of unit C.  
285 Notably HIRM, magnetic component EM1, and the pyroxene content reach peak values (Fig. 3). In  
286 addition, paleomagnetic inclinations depart markedly from the expected value for a geocentric axial  
287 dipole (GAD) field at the latitude of the coring site. Instead of dipping 61° below the horizontal  
288 plane, the magnetic inclination is flattened horizontally and has reversed polarity.

#### 289 3.1.4 Unit D

290 Unit D is from 414 to 165 cm and is the thickest and most homogenous sedimentary unit of core  
291 CPS1. The base of unit D is characterised by a sharp increase in mud content from 44% (sample 5-8  
292 at 423 cm in unit C) to 86% (sample 5-7 at 413 cm in unit D). Overall, unit D has lower amplitude  
293 variations in density, magnetic susceptibility, and mud content (Fig. 2). Two sub-units D1 (414-220  
294 cm) and D2 (220-165 cm) are identified. D1 is composed of medium silts with average grain size of  
295  $13 \pm 3 \mu\text{m}$  (19 samples) and unimodal grain size distributions (Fig. 3E). D1 has virtually no  
296 particles bigger than 2 mm. In contrast, sub-unit D2 (6 samples) has particles over 2 mm and  
297 organic debris, it is composed of coarsening upward coarse to very coarse silts with average grain  
298 size of  $30 \pm 4 \mu\text{m}$ , and grain size distributions are bimodal at the base and trimodal at the top (Fig.  
299 2E). Unit D has distinctively finer detrital minerals (Phy/(Qz+Pl+Feld)), relatively higher pyrite and  
300 gypsum contents, and lower S-ratios than the other units (Fig. 3), and there are three magnetic grain  
301 populations (Fig. 4). The two sub-units have remarkably different mineralogy and magnetic  
302 properties. The detrital minerals are much coarser in sub-unit D2 than D1, and the quartz content  
303 increases from 29% to 36% (Fig. 3). Sub-unit D2 has peak values of HIRM, EM1, and pyroxene;  
304 the same signature occurs in the debris layer of unit C and is found in overlying unit E (Fig. 3).  
305 Moreover, in sub-unit D2 the magnetic grain size indicators  $M_r/M_s$ ,  $H_{cr}/H_c$ ,  $MDF_{ARM}$ ,  $\chi_{FD}$  (Fig. 5)  
306 and the open shape of hysteresis loops (supplementary material 3) all indicate smaller ferrimagnetic  
307 grains.

## 308 3.1.5 Unit E

309 Unit E from 165 to 50 cm is composed of abundant potshards and shell fragments, rocks, and  
310 organic debris associated with large amplitude changes in density, magnetic susceptibility, and mud  
311 content (Fig. 2). The irregular surface of the debris-dominated unit E prevented the use of the multi-  
312 sensor core logger for core section 2. U-channel sampling was also impossible with many debris  
313 pieces being larger than the u-channel width (2 cm). Nevertheless, the four discrete sediment  
314 samples analyzed for granulometry are coarse silts to fine sands with highly variable average grain  
315 size (ranging from 23 to 158  $\mu\text{m}$ ) and uni- to tri-modal grain size distributions. Magnetic  
316 mineralogy and grain size are also variable (Figs. 3 and 5). There is a trend toward finer detrital  
317 minerals and the pyrite content decreases sharply to minimum values (Fig. 3). Magnetic component  
318 EM2 has maximum values (Fig. 4), with  $\chi_{fd}$  values from 3 to 7 being indicative of fine SD and SP  
319 grains.

## 320 3.1.6 Unit F

321 The uppermost unit F from 50 to 0 cm is composed of silts with a color transition from yellow to  
322 brown. The yellow mud has unimodal grain size distribution and average grain size of 9  $\mu\text{m}$   
323 (sample 1-2; Fig. 2E). The brown mud has a trimodal distribution and average grain size of 19  $\mu\text{m}$   
324 (sample 1-1). Magnetic component EM1 dominates with values up to 79%, similar to unit A.

## 325 3.2 Magnetic Mineral Assemblage

326 Magnetic properties provide information about the nature, source, and transportation of magnetic  
327 particle populations in a sediment. The studied bulk sediment is dominated by low coercivity  
328 magnetic minerals, as indicated by IRM saturation in fields below 300 mT (Fig. 4). FORC diagrams  
329 (supplementary material 3) and  $\chi_{FD\%}$  values (Fig. 5) indicate a mixture of SD, MD, and SP domain  
330 states.  $\text{MDF}_{\text{ARM}}$  has the same general down-core trend as  $\chi_{FD\%}$  and values (32 to 40 mT) are  
331 consistent with those of detrital and low coercivity biogenic magnetic minerals (Egli, 2004).  
332 Ultrafine SP particles form commonly in soils and by burning (Evans and Heller, 2003; Herries,  
333 2009), the latter being likely in Portus from cooking and heating.

334 Unmixing analysis of 50 IRM acquisition curves (Heslop, 2015) indicates that there are three end-  
335 members (EM1, 2 and 3; Fig. 4) with the fit yielding an  $R^2$  value of 0.988. The three magnetic  
336 subpopulations EM1, EM2, and EM3 have coercivities ( $B_{1/2}$ ; Fig. 4) of 21, 31, and 11 mT,  
337 respectively. The proportion of the higher coercivity population EM2 is greater in muddy sediments  
338 (units C, D, E), while the intermediate coercivity EM1 component has peak values in sandy deposits  
339 (units A, B, F). Peak concentrations of EM1 also correspond to lower relative contributions from  
340 the other magnetic component (Fig. 4), which hints at a sudden input of EM1. Moreover, EM1  
341 peaks are associated with peak pyroxene contents (Fig. 3), which likely come from basalts in the  
342 Tiber River catchment; peak HIRM values (Fig. 3) may also correspond to inputs of these volcanic  
343 grains, or to pedogenic magnetic minerals such as goethite and hematite brought by runoff.  
344 Altogether, the magnetic properties suggest that EM1 corresponds to detrital magnetic minerals of  
345 fluvial origin and that EM2 corresponds to biogenic magnetic minerals from the harbour. The  
346 lowest coercivity population EM3 is present in lower proportions (generally  $< 30\%$ ; Fig. 4), with  
347 higher absolute values in the sands of units A and B (not shown), and hysteresis loops, coercivity  
348 ratios, and remanence ratios indicative of MD magnetic minerals.

### 349 3.3 Chronology

350 Radiocarbon dating, calibration, and Bayesian depth-age modelling reveals two distinct periods in  
351 core CPS1 (Fig. 6; Table 2). The first period corresponds to sandy unit A in the lower portion of the  
352 core and has one date (calibrated 765-492 BC; modelled 764-434 BC). The second period  
353 corresponds to homogeneous muddy unit D (414-165 cm) and comprises four dates between cal AD  
354 28 and AD 311. Another date from the upper part of the high-energy flow deposit (unit C) is  
355 derived from extracted pollen/charcoal (sample OZS598) and gives a large modelled age  
356 uncertainty (Table 2) because of the presence of charcoal from burnt material with a possible  
357 inherited age. Sub-centimeter-sized ceramic fragments in the central layer of the high-energy flow  
358 deposit (unit C) indicate human activity and provide some chronological insight. The small ceramic

359 fragments have similar color and texture to those in the uppermost debris layer (unit E) that caps the  
360 harbour deposit (unit D).

361 The four dates in homogeneous unit D reveal rapid sediment deposition during the Roman Empire.  
362 The calibrated radiocarbon dates fall within the interval cal AD 28 – AD 311, and Bayesian age  
363 modelling, which takes into account the type of dated material, gives slightly younger probable  
364 dates and a larger age interval AD 128 – AD 616 because the uppermost sample OZS602 consists  
365 of a pollen/charcoal mixture (Table 2; Fig. 6). The radiocarbon-based chronology of unit D assumes  
366 constant sedimentation rates (linear interpolation) between the median modelled ages (Table 2) and  
367 suggests deposition over 165 years during the second and third century AD, at average  
368 sedimentation rates of ~1 cm per year (Fig. 6B). The true harbour deposition period could be shorter  
369 or longer than suggested by age modelling because of the uncertain uppermost date. The harbour  
370 deposit period is short relative to radiocarbon dating and calibration errors, and therefore the  
371 radiocarbon-based chronology cannot be used to date historical events. Instead, the core CPS1  
372 chronology provides a detailed sequence of events. Each cm of muddy unit D in core CPS1  
373 integrates on average less than one year, and each cube sample no more than 2 years.

## 374 4. DISCUSSION

### 375 4.1 High-resolution chronostratigraphy of core CPS1

376 The often complex stratigraphy of ancient Roman harbours (e.g., Hadler et al., 2015; Finkler et al.,  
377 2017) can be simplified to a three-unit sequence: a pre-harbour deposit, a harbour deposit, and a  
378 post-harbour or abandonment deposit (e.g., Goiran and Morhange, 2001; Marriner and Morhange,  
379 2007). The same general stratigraphy is reported here from the first stationary piston core recovered  
380 from Portus, and the new core chronostratigraphy is interpreted based on high-resolution multi-  
381 proxy analysis, age modelling, and historical context.

382 In core CPS1, the *pre-harbour deposit* corresponds to unit A and consists of stratified fluvio-coastal  
383 sediments. The smaller grain sizes and fine laminations in the upper part of unit A indicate calmer  
384 hydrological conditions possibly associated with a fluvial or lagoon environment. The radiocarbon-

385 dated seed from the uppermost laminations of unit A yield an age (Table 2; Fig. 6A) consistent with  
386 previous geomorphological and stratigraphic studies that indicate that the Tiber River mouth was  
387 located near the site in the first millennium BC (Bellotti et al., 2011; Giraudi et al., 2009).

388 The *harbour deposit* consists of unit D and covers a continuous period of about one and a half  
389 centuries (Fig. 2B). The distinctive sedimentological signature of the harbour deposit relative to the  
390 natural pre-harbour deposit is evident from granulometry, physical and magnetic properties, bulk  
391 sediment, and clay mineralogy data (Figs. 2, 3, 4). The harbour deposit in the man-made Canale  
392 Traverso is characterised by greater mud content and finer detrital minerals (Fig. 3) that indicate a  
393 calmer, more isolated depositional environment. Relatively higher pyrite contents, lower  
394 ferrimagnetic mineral concentrations (lower  $\kappa$ ,  $M_s$ , NRM, ARM), and the presence of biogenic  
395 magnetite (EM2; Fig. 4) indicate a more organic-rich deposit and more hypoxic conditions. An age  
396 for the start of the harbour deposit in Canale Traverso core CPS1 is obtained by interpolation and is  
397 modelled at AD 120 – AD 186 (Fig. 6). This age estimate suggests that the onset of harbour  
398 deposition in core CPS1 followed the massive enlargement and reorganization of Portus by Trajan  
399 that was completed by AD 117 (Keay et al., 2005). However, it is believed that Canale Traverso  
400 was built earlier, during the first century AD, as part of the initial Claudius harbour construction.  
401 This is based on the hypothesis of Testaguzza (1970) who suggested that the Darsena, Canale  
402 Traverso, and Fossa Traiana formed a complex built at the same time, and on the radiocarbon  
403 chronology of Canale Traverso core CT1 that contains sediment from the first century (Salomon et  
404 al., 2012). If the Canale Traverso was built during the first century AD, the age of the basal harbour  
405 deposit in core CPS1 does not date construction of Canale Traverso, but rather indicates that  
406 dredging operations removed about one century of previous harbour deposit. This issue is  
407 considered in section 4.3. The modelled age for the end of harbour deposit is obtained by  
408 interpolation at AD 164 - AD 643 (Fig. 6). The large age interval is inherent to the nature of the  
409 dated material (charcoal-rich pollen sample OZS602). Nevertheless, the modelled age interval AD

410 177 - AD 436 is obtained if sample OZS602 is not considered and a constant sedimentation rate is  
411 assumed from sample OZS601 to the top of the harbour sequence (Fig. 6).

412 Units E-F represent the *post-harbour deposit*. Accumulation of debris (unit E) into Canale Traverso  
413 marked the end of its use as a waterway. Uppermost unit F includes floodplain deposits and the  
414 recent soil deposit. What stands out in the stratigraphy of core CPS1 are the two consecutive  
415 reworked deposits (units B and C) located between the pre-harbour and harbour deposits. For the  
416 first time at Portus, reworked layers are identified precisely using high-resolution core scanning and  
417 multi-proxy analyses. Units B and C are discussed in more detail below.

#### 418 4.2 New Chronostratigraphic Marker

419 The estimated end of harbour deposition in Canale Traverso is consistent with dates from cores S13  
420 and TR14 located downstream in the access channel (70 m and 170 m distance from CPS1; Fig 1B),  
421 but is significantly younger than in core CT1, which is located upstream toward Fossa Traiana (130  
422 m distance from CPS1; Fig. 1B). Based on grain size data in core CT1, Salomon et al. (2012)  
423 reported two distinct harbour deposits; one with low Tiber River influence and one with high Tiber  
424 River influence (Fig. 7; Salomon et al., 2012). Our pilot magnetic analysis supports this stratigraphy  
425 with a distinct magnetic mineral assemblage in the core CT1 high Tiber River influence deposit and  
426 similar magnetic assemblages for the CT1 low Tiber influence deposit and core TR20 from the  
427 access channel (supplementary material 1). The sudden change in CT1 sediment type corresponds  
428 to a sharp  $\kappa$  increase, which can be correlated to a similar sharp  $\kappa$  increase at a similar median  
429 modelled age in core CPS1 (Fig. 7). While the sharp  $\kappa$  increase in core CT1 likely corresponds to  
430 increased river transport of ferrimagnetic detrital particles, the sharp  $\kappa$  increase in core CPS1  
431 corresponds to the debris layer (unit E; Fig. 2), which also has large amplitude changes attributed to  
432 abundant ceramic fragments, rocks, and overall heterogeneous debris having highly variable  
433 magnetic properties. At a similar modelled age in core TR14, a sudden onset of marine influence is  
434 evident in the access channel based on geochemistry data (Fig. 7; Delile et al., 2014). Core CPS1 is  
435 located between cores TR14 and CT1 and, thus, provides a link for correlating the Portus cores.

436 Damming of Canale Traverso with debris at the junction of the access channel (site CPS1) would  
437 have blocked river input into the harbour basins, changing predominantly fluvial to marine  
438 influence in the access channel (TR14), and increasing river influence upstream of Canale Traverso  
439 (CT1). Combining age modelling,  $\kappa$  records, and previous grain size and geochemistry studies  
440 (Salomon et al., 2012; Delile et al., 2014) reveals a new chronostratigraphic marker at Portus at the  
441 time Canale Traverso was decommissioned. This change may correspond to Portus becoming a  
442 town in the 4<sup>th</sup> century when the immediate vicinity of Canale Traverso changed its function from  
443 administrative to residential (Keay et al, 2005; Paroli, 2005).

#### 444 4.3 Reworking of Sediments into a Roman Canal

##### 445 4.3.1 The Dredged Deposit

446 CT scan and digital images clearly reveal an oblique erosional contact that cuts natural laminated  
447 deposits and is overlain by 93 cm of sands with angular mud clasts in the upper part (unit B; Fig. 2).  
448 The sands deposited before and after the oblique contact have similar physical, mineralogical, and  
449 magnetic properties (Figs. 2, 3, 4, 5), which suggests that the reworked unit B was sourced locally.  
450 The presence of rectangular-shaped mud clasts with immature texture (Lie et al., 2017) further  
451 indicates a local mass movement event such as a river bank collapse, a landslide or debris flow,  
452 possibly triggered by human activities or an earthquake. However, observed immature mud clasts  
453 embedded in much coarser sands imply a local source of mud deposited in a calm environment and  
454 a local source of sands deposited in a vigorous hydrodynamic environment. Contemporaneous mud  
455 and sand deposits within a small fabricated canal appears unlikely, and a local mass movement is  
456 unlikely to explain the oblique erosive contact in the canal center. Another hypothesis is a weather-  
457 related event such as storm, tsunami or flood. Such high-energy events have been reported in the  
458 nearby Ostia harbour (Goiran et al., 2014; Hadler et al., 2015); whilst event deposits in Portus are  
459 reported here for the first time, the river influence at Portus is well-documented (Goiran et al., 2010;  
460 Delile et al., 2014; Salomon et al., 2012; 2014). A flood appears likely; however, a high-energy  
461 flow is inconsistent with the ungraded sands and immature mud clasts of unit B, which rather point

462 to low velocity flow. A more likely explanation is that unit B corresponds to an admixture of coastal  
463 sands and harbour mud that were shoveled locally and dropped into place, forming a dredged  
464 deposit (Table 3; Fig. 8). Multi-proxy analyses of the mud clast (sample 6-8) indicate contrasting  
465 magnetic and physical properties with the later harbour mud of unit D, which points to different  
466 depositional conditions before and after dredging. The differences include distinctively higher S-  
467 ratio and higher EM2 biogenic proportion in the mud clast (Figs. 3, 4), and the lowest D50 value of  
468 core CPS1, with the grain size of sample 6-8 plotting in the lower energy or decantation end of  
469 Figure 5F. It is, thus, possible that dredging was performed at the time of the major Trajan work,  
470 which drastically changed the harbour infrastructure (Keay et al., 2005) and hydrodynamics (Millet  
471 et al., 2014).

472 Further support for calm hydrodynamic conditions at the time of deposition of unit B comes from  
473 the paleomagnetic inclination. Muddy sands in upper unit B have a strong, stable, and well-defined  
474 magnetic remanence with inclination values near the GAD field value for the latitude of Portus  
475 (Figs. 9D, 10). In unit B, three distinct magnetic particle populations (Fig. 4) form a single  
476 magnetization component, which indicates that a detrital remanent magnetisation (DRM)  
477 acquisition process at about the time of deposition is more likely than a chemical remanent  
478 magnetisation (CRM) that would have happened anytime later. Additional support for a DRM  
479 comes from the predominance of biogenic particles in the supposedly dredged deposit because such  
480 particles are small and readily undergo chemical alteration. Such rapid remanence acquisition  
481 within an event deposit resembles redeposition experiments of slurries with low water content  
482 (Barton et al, 1980) and adobe brick fabrication, where clay and water mixtures rapidly acquire a  
483 shear or shock remanent magnetisation when thrown into a mould (Games, 1977; 1983). The  
484 paleomagnetic results, thus, indicate limited resuspension of slurry material and calm hydrodynamic  
485 conditions at the site, which further supports the suggested dredged deposit interpretation for unit B.

## 486 4.3.2 The Hyperpycnal Deposit

487 Immediately overlying the dredged deposit (unit B) is another reworked deposit of 66 cm thickness  
488 (unit C), which is composed of a coarsening upward basal sequence, a 30 cm-thick central debris  
489 layer, and a fining upward top sequence. Unit C is a typical hyperpycnite that results from high-  
490 energy flow that sorts particles into a coarsening sequence during the waxing stage of the flow, and  
491 a fining upward sequence during the waning stage of the flow (Mulder, 2003; Fig. 8). The  
492 hyperpycnal deposit in core CPS1 is readily identifiable using grain size, bulk mineralogy, and  
493 magnetic data (Fig. 10). In addition to grain size sorting, flattened paleomagnetic inclinations (Figs.  
494 9C and 10) are an indicator of high-energy deposits (St-Onge et al., 2004). The distinct sedimentary  
495 properties of the central debris layer provide a clear river input signature, particularly using the  
496 fluvial magnetic component EM1, the content of high coercivity minerals (HIRM) and pyroxenes,  
497 and a higher smectite/(illite+chlorite) ratio (Fig. 3). Coeval peak values of these parameters in the  
498 debris layer point to a common fluvial source and transport mechanism for different types of detrital  
499 material.

500 Hyperpycnal deposits in natural environments are associated with dense water flowing into less  
501 dense water, the most typical example being sediment-laden river floodwaters plunging into the sea  
502 (Mulder et al., 2003). Catastrophic drainage events such as glacial outburst flood and dam-breaks  
503 can also generate hyperpycnites (e.g., St-Onge et al., 2004; Mulder et al., 2009; Mulder and  
504 Chapron, 2011; Duboc et al., 2017). To our knowledge, this is the first report of a hyperpycnal  
505 deposit in a fabricated structure. How did a hyperpycnite form in Canale Traverso? Did the high-  
506 intensity flow have natural causes, such as in association with a river flood, storm, or tsunami event,  
507 or was it human-induced such as through opening of a canal gate, or from dredging? The hypothesis  
508 of a major river flood is ruled out by the presence of a thick central debris layer in unit C. While a  
509 typically steady and sustained flood flow carries a suspended load over hours to weeks, the debris  
510 layer of unit C instead indicates bedload transport during a short-duration high-energy flow.  
511 Tsunami or extreme storm waves are high-energy, short-duration events that form graded deposit

512 with debris (Robke and Vött, 2017). A single fining upward sequence (waning flow) is typically  
513 reported for such events, but the coarsening upward sequence (waxing flow) observed in Canale  
514 Traverso could have been generated by inflow and backflow events from the sea, travelling in the  
515 harbour basins, and up the canals. However, there is no erosive contact at the base of unit C as may  
516 be expected for a tsunami, the debris layer has a clear fluvial signature (Table 3), and previous  
517 combined geophysical surveys, sedimentological, macro- and microfossil, and geochemical  
518 analyses have not revealed a tsunamite in Portus (Delile et al., 2014; DiBella et al., 2011; Goiran et  
519 al., 2010; Pepe et al., 2013; Sadori et al., 2010; Salomon et al., 2012; 2014).

520 Assuming that the deposit tracked the flow hydrograph, the presence of debris and an erosive  
521 contact are key to identifying the type of flow and the conditions under which the hyperpycnite  
522 formed (Mulder, 2003). According to Mulder (2003), the debris layer located below the fining  
523 upward sequence and the absence of an erosive basal contact in unit C best correspond to a natural  
524 dam-break event. Erosion of a natural sediment dam appears to be a good analogue for opening a  
525 canal gate in a Roman harbour context. Opening Canale Traverso would have suddenly increased  
526 water flow from the river (waxing flow; Fig. 8), eroding and transporting sediment and debris that  
527 would have blocked the canal, or accumulated on the river side of the gate, and then water flow  
528 would have decreased (waning flow; Fig. 8) until steady flow is reached. Another hypothesis is a  
529 dredging-generated hyperpycnal deposit, where the fining upward sediment would be deposited by  
530 suction under the dredging shovel. However, absence of an erosive contact between units B and C  
531 does not support this idea, which also does not account for the coarsening-upward and debris layers  
532 (Table 3). Based on multi-proxy sediment analysis and high-resolution stratigraphy, our preferred  
533 interpretation for unit C is a canal dam-break hyperpycnal deposit (Fig. 8), which is an analogue for  
534 natural dam-breaks in a fabricated structure.

#### 535 4.3.3 Event deposits summary

536 Units B and C are consecutive event deposits. Our multi-proxy data set suggests that the event  
537 deposits were human-induced, with our preferred interpretation of unit B being a dredged deposit,

538 and unit C being a canal dam-break hyperpycnal deposit (Table 3). The flood, tsunami or storm  
539 hypotheses are plausible alternative interpretations for unit C, which was deposited under high-  
540 energy hydrodynamic conditions. The hypothesised dredging event may have occurred during the  
541 Trajanic redevelopment (AD 112-117) (Keay et al., 2005), based on age modelling and the presence  
542 of different muds in the harbour deposit (unit D) and in the dredged deposit (mud clast sample 6-8).  
543 Sediment mix and drop during dredging would have formed the dredged deposit, while waxing and  
544 waning flows would have formed the hyperpycnal deposit (Fig. 8). For the first time, major event  
545 deposits are described in Portus, and identification of a 159-cm-thick reworked sediment has  
546 important implications for age-depth modelling and water depth reconstructions. The  
547 Palaeoenvironmental Age-Depth Model (PADM) chart (Fig. 6A; Salomon et al., 2016) indicates  
548 water depth based on core stratigraphy and reconstructed local relative sea level (Goiran et al.,  
549 2009). The water depth in Canale Traverso after dredging (top of Unit C) was about 3 m, which  
550 provides sufficient draught for ships 150 t and smaller, such as *Bourse de Marseille* and *Fiumicino*  
551 *I* (Gassend, 1982; Pomey and Tchernia, 1978; Pomey and Rieth, 2005; Boetto, 2010).

#### 552 4.4 Canal Gate Usage

553 Were canal gates used in Portus for flood control? Tiber floods in the early first century AD had  
554 silted the fluvial harbour of Ostia, which was then abandoned (Goiran et al., 2014; Hadler et al.,  
555 2015). It is probable that the Romans used canal gates for flood management in Portus considering  
556 that a main motivation for canal construction was to free Rome from flood dangers (AD 46  
557 inscription - Thylander, 1952: B310 = CIL XIV 85 = ILS 207; and AD 102-109 inscription -  
558 Thylander 1952: B312 = CIL XIV 878 = CIL VI 964 = ILS 5797a; Keay et al., 2005). Canal gate  
559 technology was presumably used as early as the 3<sup>rd</sup> century BC by Ptolemy II in the Ancient Suez  
560 Canal, which was re-built by Trajan (Moore, 1950). The major redevelopment of Portus by Trajan  
561 in AD 112-117 (Keay et al., 2005) took place soon after a catastrophic Tiber flood as reported by  
562 Pliny the Younger in the winter of 108-109 AD (Syme, 1985). However, it is unknown if and how  
563 that flood impacted Portus. We now evaluate if multi-proxy analysis of the harbour deposit in core  
564 CPS1 (unit D) is consistent with canal gate use.

565 The hyperpycnal deposit has a clear fluvial input signature. Sub-unit D2 has a similar fluvial  
566 signature, with peak values of EM1, HIRM, pyroxene, and higher smectite/(illite+chlorite) ratio  
567 (Fig. 3). However, there are significant differences, including the absence of an erosive contact or  
568 discontinuity in D2, no grading, unchanged magnetic particle concentrations (Fig. 2), and no  
569 flattened paleomagnetic inclinations (Fig. 9A). These results indicate that D2 is not an event  
570 deposit, but rather a harbour deposit with marked river input over a period of about 60 years. Grain  
571 size, magnetic, and bulk mineralogical data clearly indicate increased sand content, increased  
572 fluvial magnetic particle populations (EM1), and coarser detrital minerals in D2 relative to the  
573 previous harbour deposit D1 (Fig. 3). The D90-D50 diagram also indicates greater hydrodynamic  
574 energy for D2 relative to D1 (Fig. 5F). These results are consistent with ostracod and pollen  
575 analyses, which indicate greater freshwater input in the upper part of the harbour sequence in core  
576 S5 (location on Figure 1; Sadori et al., 2010), as well as increased magnetic susceptibility in core  
577 TR14 (Figs. 1 and 7; Delile et al., 2014). Together, the results can be interpreted as: (1) sub-unit D2  
578 being deposited in an open Canale Traverso configuration that connected the Tiber River and the  
579 harbour basins after a gate-controlled period (sub-unit D1), and/or (2) rapid climate change to  
580 wetter conditions and stronger fluvial impact in the early 3<sup>rd</sup> century AD. River input proxies for  
581 sub-unit D1 (Fig. 3) do not have coeval peak values. Only the clay mineral indicator has peak  
582 values such as in the hyperpycnal deposit and D2. It is possible that if the coarser grains tracked by  
583 EM1, HIRM, and pyroxene content were blocked from Canale Traverso by canal gates, finer clay  
584 particles may have flowed over the gates during flood events or reached the canal by local runoff in  
585 Portus. Overall, reduced river input in D1, sulphidic diagenetic conditions indicated by increased  
586 pyrite formation, and warmer and more saline conditions indicated by higher gypsum content (Fig.  
587 3) are consistent with D1 being deposited under gate controlled conditions. The multi-proxy dataset  
588 provides abundant evidence in support of the idea of canal gate usage for water flow management.  
589 The period of sub-unit D2 corresponds to the onset of Roman Empire contraction which is  
590 characterized by increased climate instability (Büntgen et al., 2011), political turmoil, cultural

591 change, socio-economic instability (Duncan-Jones, 2006), and human health issues such as the  
592 large-scale Plague of Cyprian (AD 251-266; Vuorinen, 1997). Thus, climatic or socio-economic  
593 instabilities could possibly explain a sudden change in harbour management and canal gate  
594 operation, and the later decommissioning of Canale Traverso. The capacity to close the only inland  
595 waterway from the Tiber River to the capital city's harbour basins would undoubtedly have been  
596 advantageous during flood events and at times of war. The sedimentary archives in Portus have an  
597 undeniably strong anthropogenic influence; the depositional basin itself is man-made and the site  
598 has witnessed intense human activity. Core CPS1 provides evidence of over one and a half centuries  
599 of nearly annually-resolved harbour occupation at the height of the Roman Empire, with humans  
600 using technologies in response to natural environmental and societal stresses.

## 601 5. CONCLUSIONS

602 Multi-proxy and high-resolution sediment analysis, including piston coring, core scanning, grain  
603 size, paleo- and rock-magnetism, bulk and clay mineralogy, and radiocarbon dating was applied in  
604 an ancient harbour context to understand its history. This method revealed 159 cm of event deposits  
605 (units B and C) and 249 cm of harbour sediment (Unit D) in core CPS1. Without piston coring, a  
606 fine stratigraphy would not be preserved and it is likely that the suggested dredged deposit (Unit B)  
607 would be indiscernable from the pre-harbour deposit (unit A), and most crucially for water depth  
608 reconstruction, the 66-cm-thick hyperpycnal deposit (Unit C) would have been included in the  
609 harbour deposit (Unit D). Portus core correlation was achieved using magnetic susceptibility and  
610 age modelling, and a chronostratigraphic marker at the time Canale Traverso was decommissioned  
611 is proposed to reconcile previous core interpretations. Piston coring, whole core CT-scanning, and  
612 magnetic analyses have proven useful for event stratigraphy and geoarchaeology. Routine analysis  
613 of the type presented here could greatly improve chronostratigraphic analysis and water depth  
614 reconstruction of ancient harbour deposits.

615 Our works also provides rare insights into Roman harbour technologies. Dredging activity in Canale  
616 Traverso during the Trajanic-Hadrianic period (2<sup>nd</sup> century AD) provides a well-supported

617 hypothesis for explaining unit B. The dredging event has not been reported previously and predates  
618 the known 4<sup>th</sup> century AD dredging depicted in a Roman epigraph. Portus was built in a delta,  
619 without bedrock that may hold direct evidence of canal gates. While future excavation of canal  
620 walls may provide direct evidence, this work provides indirect sedimentological indicators  
621 consistent with hypothesised canal gates at Portus, and supports the ideas of Testaguzza (1970) and  
622 Salomon et al. (2014). Magnetic analysis has been applied here to ancient harbour geoarchaeology  
623 for the first time. It has proven useful at Portus for core correlation (magnetic susceptibility),  
624 identification of event deposit (paleomagnetic inclination,  $\kappa_{ARM}/\kappa$ , IRM acquisition), and for  
625 providing a river input proxy (S-ratio, HIRM, IRM acquisition). Sedimentary magnetism is a  
626 versatile tool that can, therefore, be added to the geoarchaeologist's toolkit (Marriner and  
627 Morhange, 2007).

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915 **FIGURE CAPTIONS**

916 Figure 1. A) Location of Portus near the Tiber River in Italy and B) position of the coring sites,  
917 including the stationary piston cores CPS1 and CPS4, and previous rotary cores discussed in the  
918 text.  $\kappa$  is magnetic susceptibility. C) Modern and D) ancient configuration of Portus (modified from  
919 Goiran et al., 2010). E) Piston coring of core CPS1 in September 2011.

920 Figure 2. The stratigraphy of core CPS1 is primarily based on A) CT number, B) gamma-ray  
921 attenuation, C) magnetic susceptibility, and D) mud content. Core CT-scan images and photographs  
922 are on the left, with the sampling log (cubes and u-channels) and simplified stratigraphic log with  
923 the types of deposit for units A to F. Square symbols are used for data measured on cube samples,  
924 and simple line for u-channel samples. The dashed line indicates the D1/D2 transition within the  
925 harbour deposit. E) Typical grain size distributions for each deposit type.

926 Figure 3. Bulk, clay, and magnetic mineralogy results. A simplified stratigraphic log with deposit  
927 type is shown on the left. The dashed line indicates the D1/D2 transition within the harbour deposit.

928 Figure 4. Magnetic end-member (EM) analysis with three components representing distinct  
929 magnetic particle populations in core CPS1 obtained by unmixing 50 isothermal remanent  
930 magnetization (IRM) acquisition curves (Heslop and Dillon, 2007; Heslop, 2015). A) IRM end-  
931 members and B) down-core variations of the relative proportions of EM1, 2, and 3. Magnetic  
932 properties and multi-proxy investigations indicate that EM1 represents detrital minerals of fluvial  
933 origin, EM2 represents biogenic minerals from the harbour, and EM3 represents lower coercivity  
934 multi-domain (MD) particles.

935 Figure 5. Magnetic grain size indicators. A) Coercivity ratio  $H_{cr}/H_c$ , B) magnetization ratio  $M_r/M_s$ ,  
936 C) median destructive field of the anhysteretic remanent magnetization  $MDF_{ARM}$ , and D)  
937 frequency-dependent magnetic susceptibility  $\chi_{FD\%}$ , compared with bulk grain size indicators E) mud  
938 content, and F) D90 vs D50 biplot (Passega, 1964) for core CPS1.

939 Figure 6. Radiocarbon-based chronology for core CPS1. Bayesian age modelling of A) full CPS1  
940 core in a Palaeoenvironmental Age-Depth Model (PADM) chart (Salomon et al., 2016) and B) the  
941 harbour deposit. Light grey distributions are calibrated ages and dark grey are modelled ages with a  
942 simple sequence Bayesian model performed with the Charcoal Outlier Model using OxCal v4.2.3  
943 (Bronk Ramsey and Lee, 2013). A simplified stratigraphic log and lithology for units A to F are  
944 shown.

945 Figure 7. Comparison of Portus core stratigraphy and magnetic susceptibility data along a NNW-  
946 SSE transect. The cores are aligned at the proposed chronostratigraphic marker at the time of  
947 decommissioning Canale Traverso. Debris dammed Canale Traverso at the access channel entrance  
948 (core CPS1; this study) and resulted in more marine conditions in the access channel (core TR14;  
949 Delile et al., 2014) and more fluvial conditions upstream in Canale Traverso (core CT1; Salomon et  
950 al., 2012). The top of core TR14 is at -0.19 m relative to current sea level, core CPS1 at +0.3 m, and  
951 CT1 at +1.8 m. See core locations in Figure 1.

952 Figure 8. Core CPS1 reconstructed depositional history.

953 Figure 9. Typical demagnetization plots, orthogonal projection diagrams, and stereonet projections  
954 for A-B) the harbour deposit, C) the hyperpycnal deposit, and D) the dredged deposit.  
955 Demagnetization steps considered for the principal component analysis (PCA) calculation are  
956 indicated and highlighted in red. Stable and well defined paleomagnetic directions were obtained  
957 for (A, B) harbour mud samples and most reworked deposits, with inclination values (bold) near the  
958 expected GAD field value of  $60.8^\circ$  (D). In contrast, flattened to reversed polarity inclinations are  
959 obtained within the high-energy flow (C).

960 Figure 10. Magnetic parameters for identified reworked deposits in core CPS1. Magnetic  
961 susceptibility ( $\kappa$ ), saturation magnetization ( $M_s$ ), hard isothermal remanent magnetization (HIRM),  
962 paleomagnetic inclination, the magnetic grain size indicator  $\kappa_{ARM}/\kappa_{LF}$ , and concentration of  
963 magnetic population EM1 are proxies for the hyperpycnal deposit, with peak values in the central  
964 debris layer. Grain size data indicate a typical hyperpycnite sequence, coarsening upward from  
965 predominantly fine silts to coarse sands during the waxing flow, and fining upward to fine sands  
966 during the waning flow. The dredged deposit has relatively homogenous properties within the sand,  
967 with distinct values associated with mud clasts. Note that inclination values near the geocentric  
968 axial dipole (GAD) value are recorded in the muddy sands of the dredged deposit.

969 **SUPPLEMENTARY FIGURES**

970 Sup. Figure 1. Results of rock- and paleo-magnetic pilot studies.

971 Sup. Figure 2. First-order reversal curve (FORC) diagrams and hysteresis loops for samples from  
972 sub-units D1 and D2 of core CPS1.

973 **TABLES**

974 Table 1. Laboratory analyses performed on core CPS1.

975 Table 2. Radiocarbon dating of core CPS1, calibration, and age modelling.

976 Table 3. Summary of data and hypotheses for the event deposits of core CPS1

ACCEPTED MANUSCRIPT

Table 1. Laboratory analyses performed on core CPS1.

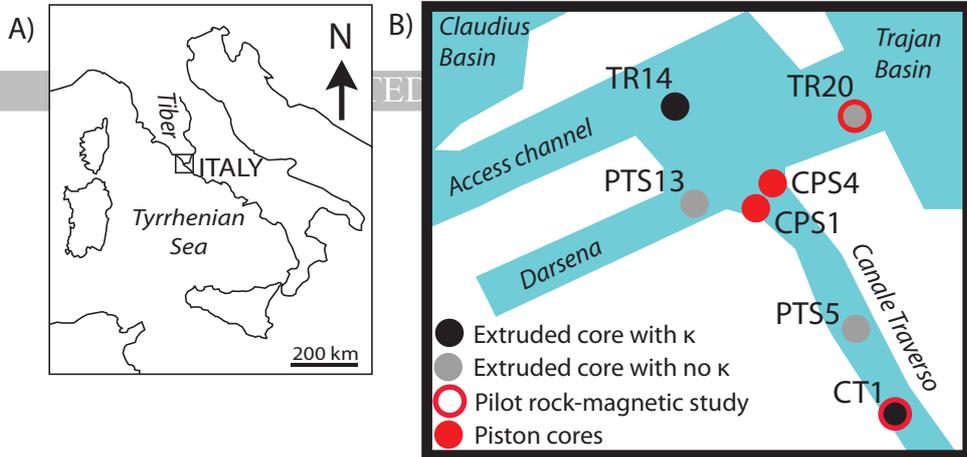
Analysis	Resolution
CT scan (density proxy)	0.06 cm
Gamma ray attenuation (density proxy)	1 cm
Volumetric magnetic susceptibility ( $\kappa$ )	1 cm
Image scan	0.025 cm
Volumetric magnetic susceptibility ( $\kappa$ )	0.5 cm
Natural remanent magnetisation (NRM)	1 cm
Anhysteretic remanent magnetisation (ARM)	1 cm
Frequency dependant magnetic susceptibility ( $\chi_{FD}$ )	10 cm
Laser granulometry (bulk grain size )	10 cm
X-ray diffraction (bulk mineralogy)	10 cm
Hysteresis curves and properties ( $M_r$ , $M_s$ , $H_{cr}$ , $H_c$ )	10 cm
Isothermal remanent magnetisation (IRM) acquisition	10 cm
First-order reversal curves (FORC)	7 samples

Table 1. Radiocarbon dating of core CPS1, calibration, and age modelling.

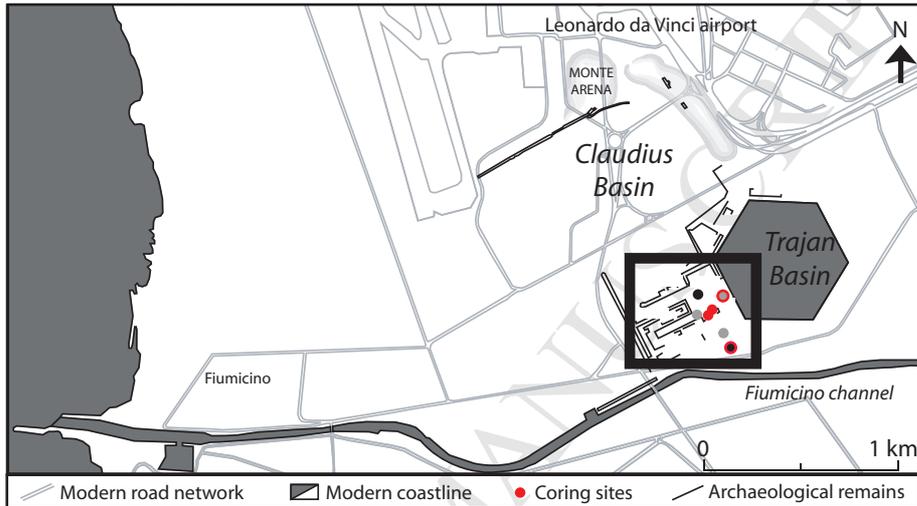
Core depth (cm)	Unit	Sample name	Laboratory code	Material	$\delta^{13}C$	$^{14}C$ age (yr BP)	Calibrated age (AD/BC)		Modelled age (AD/BC)		
							min	max	min	max	median
170-171	D	CPS1-3/25-26cm	OZS602	pollen/charcoal	-25.0 $\pm 0.1$	192 5 $\pm$ 20	A D 28	AD 127	AD 16 3	AD 616	AD 294
227-228	D	CPS1-3/82-83cm	OZS601	coniferous bud	-24.5 $\pm 0.1$	181 0 $\pm$ 25	A D 13 1	AD 311	AD 15 5	AD 311	AD 220
269-270	D	CPS1-4/25-26cm	OZS600	seed	-23.8 $\pm 0.1$	186 5 $\pm$ 25	A D 80	AD 225	AD 14 0	AD 226	AD 186
357-358	D	CPS1-5/15-16cm	OZS599	seed	-25.3 $\pm 0.1$	181 0 $\pm$ 20	A D 13 1	AD 310	AD 12 8	AD 202	AD 155
414-415	C	CPS1-5/72-73cm	OZS598	pollen/charcoal	-25.0 $\pm 0.1$	265 0 $\pm$ 20	83 3 B C	795 BC	73 5 BC	AD 55	432 BC
580-581	A	CPS1-7/46-47cm	OZS597	seed	-26.2 $\pm 0.1$	247 0 $\pm$ 25	76 5 B C	492 BC	76 4 BC	434 BC	625 BC

Table 3. Summary of data and hypotheses for the event deposits of core CPS1.

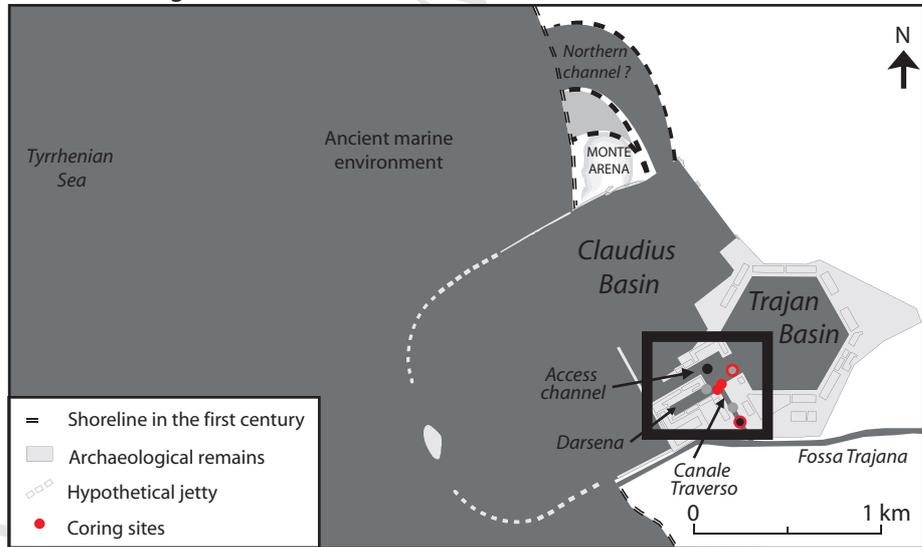
		Climate-driven		Human-driven	
		Tsunami	Flood	Canal gate	Dredging
		HIGH ENERGY FLOW			LOW ENERGY FLOW
Unit C	Fining upward				
	Debris layer (30 cm)				
	Fluvial signature of debris layer				
	Coarsening upward				
Unit B	GAD paleomagnetic inclination				
	Angular mud clasts				
	Massive (ungraded) sands				
	Erosive contact				

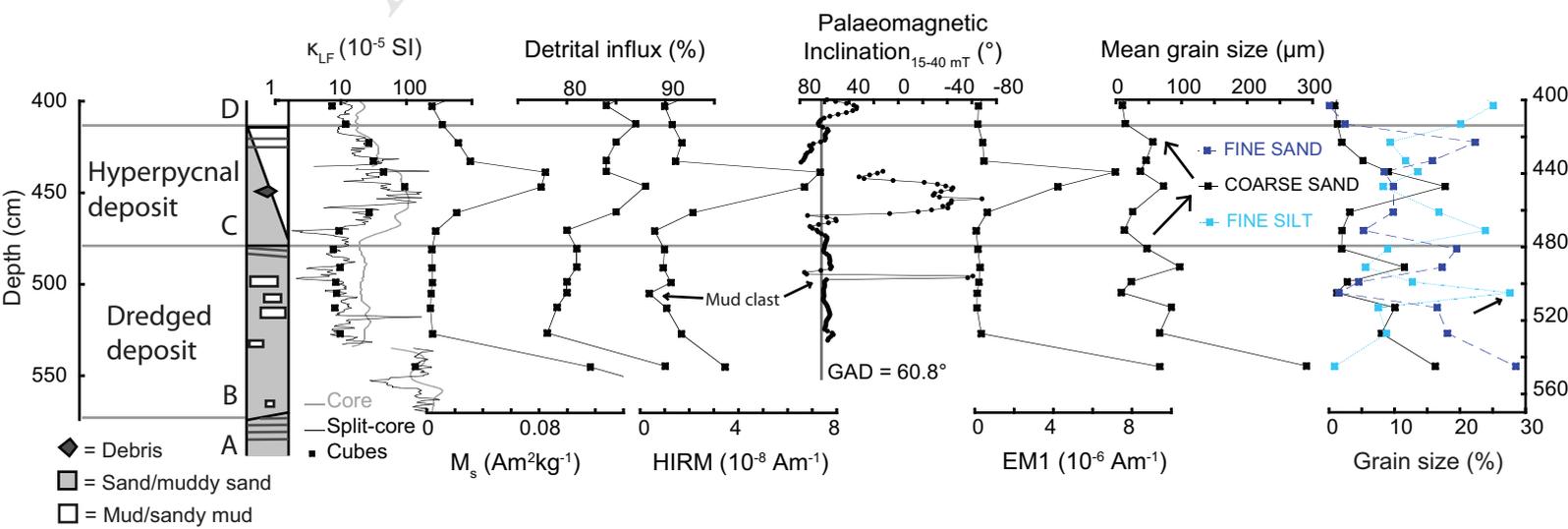


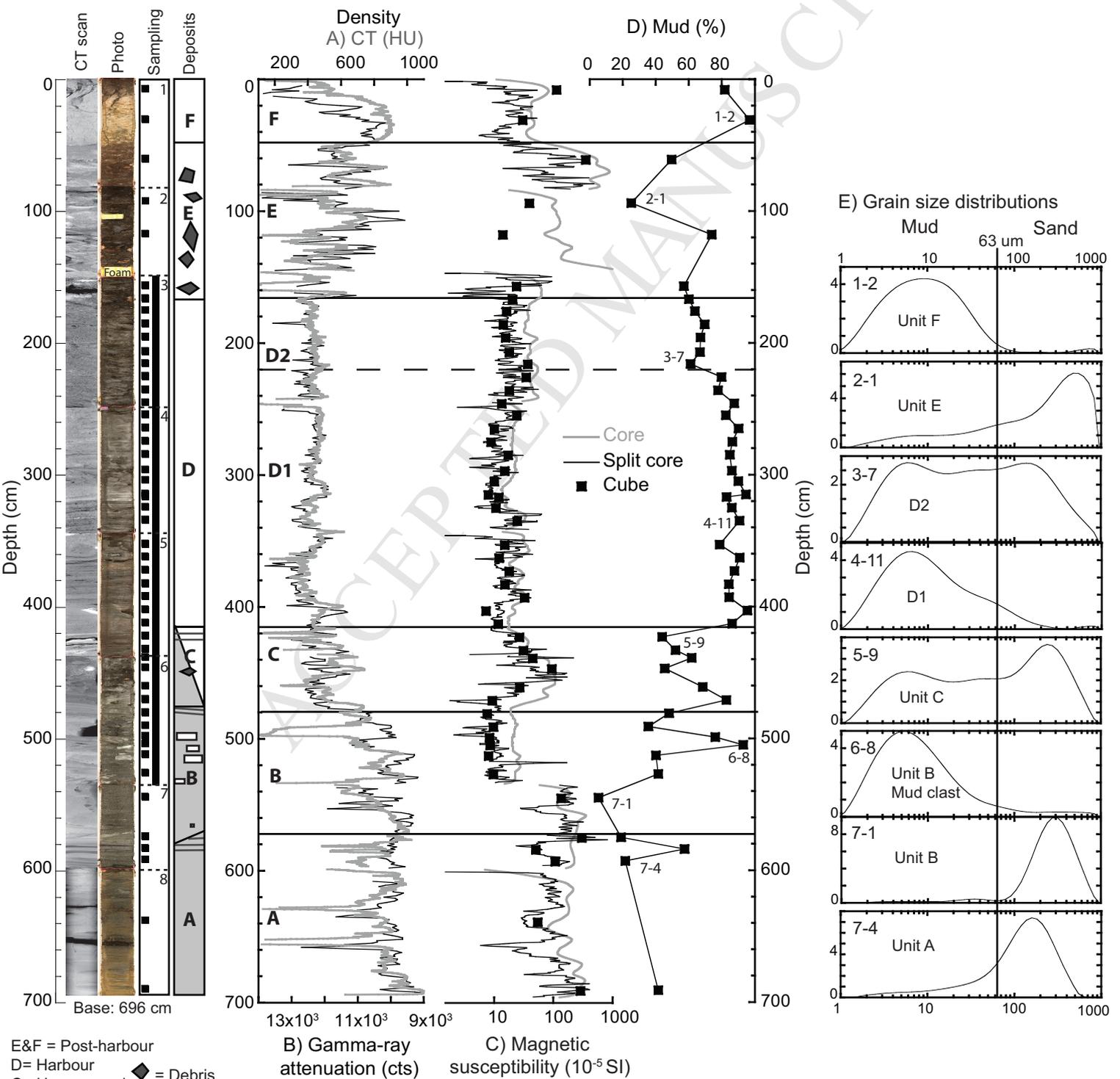
C) Modern configuration of the Portus area

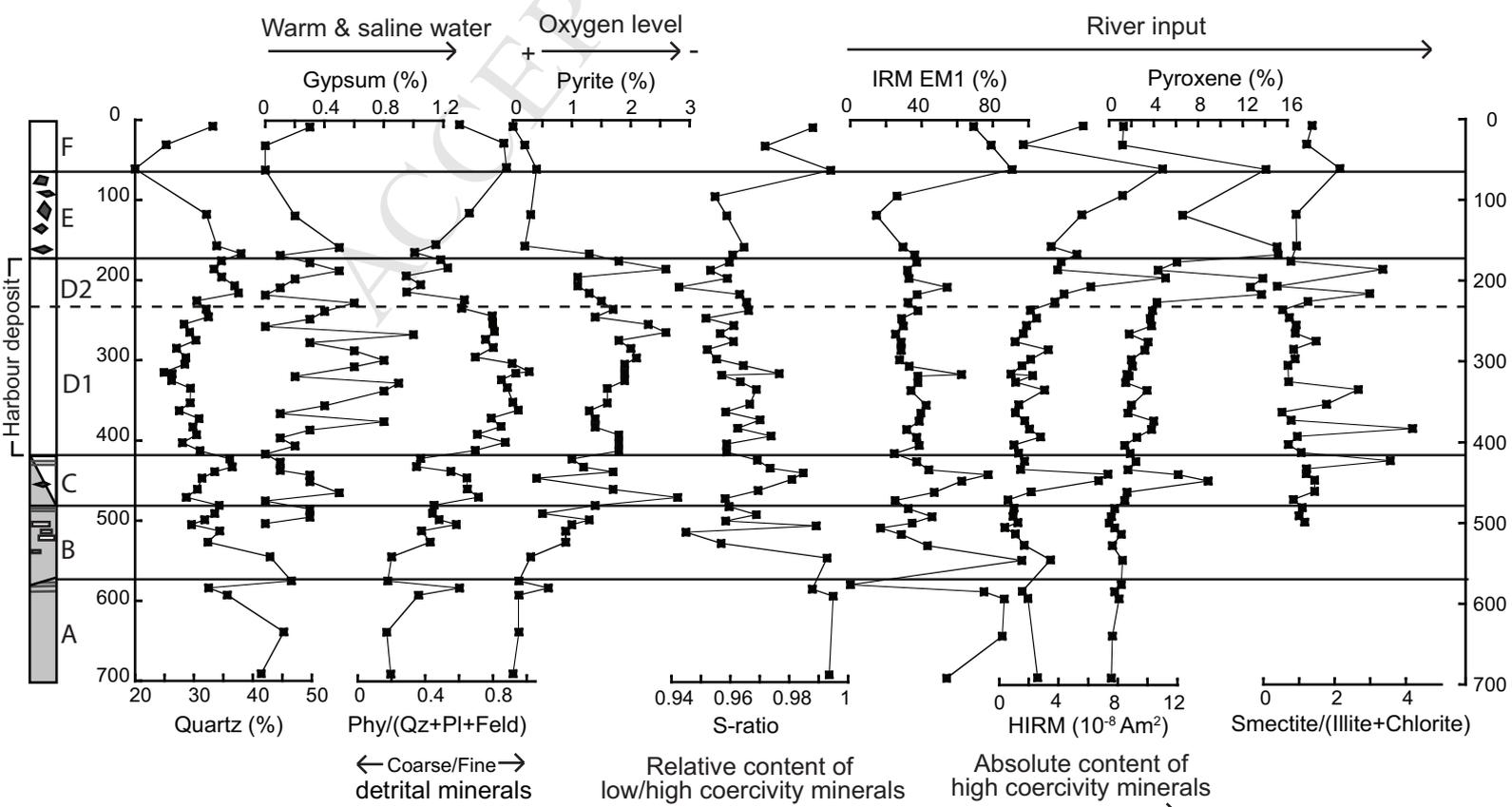


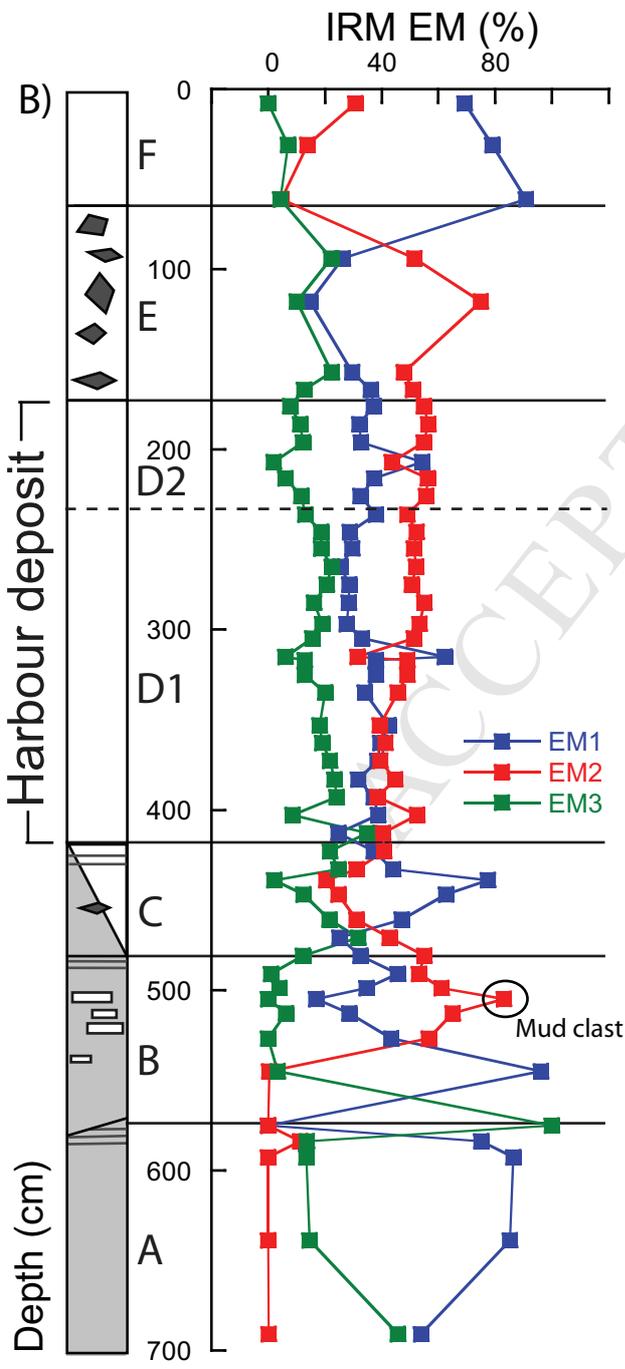
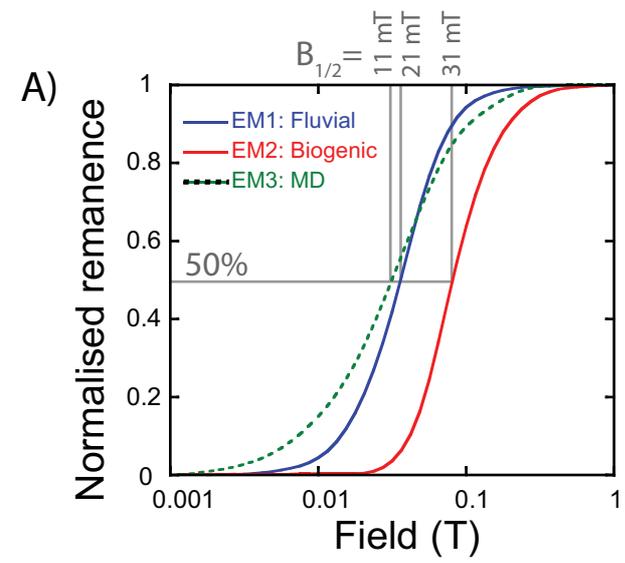
D) Ancient configuration of Portus

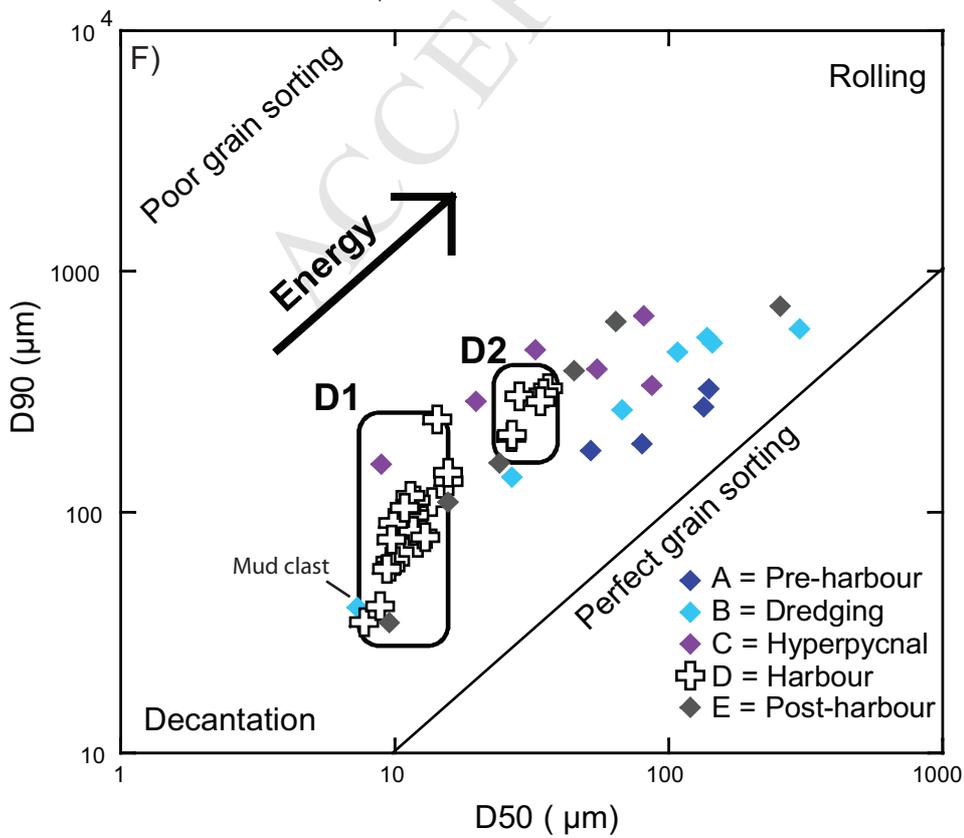
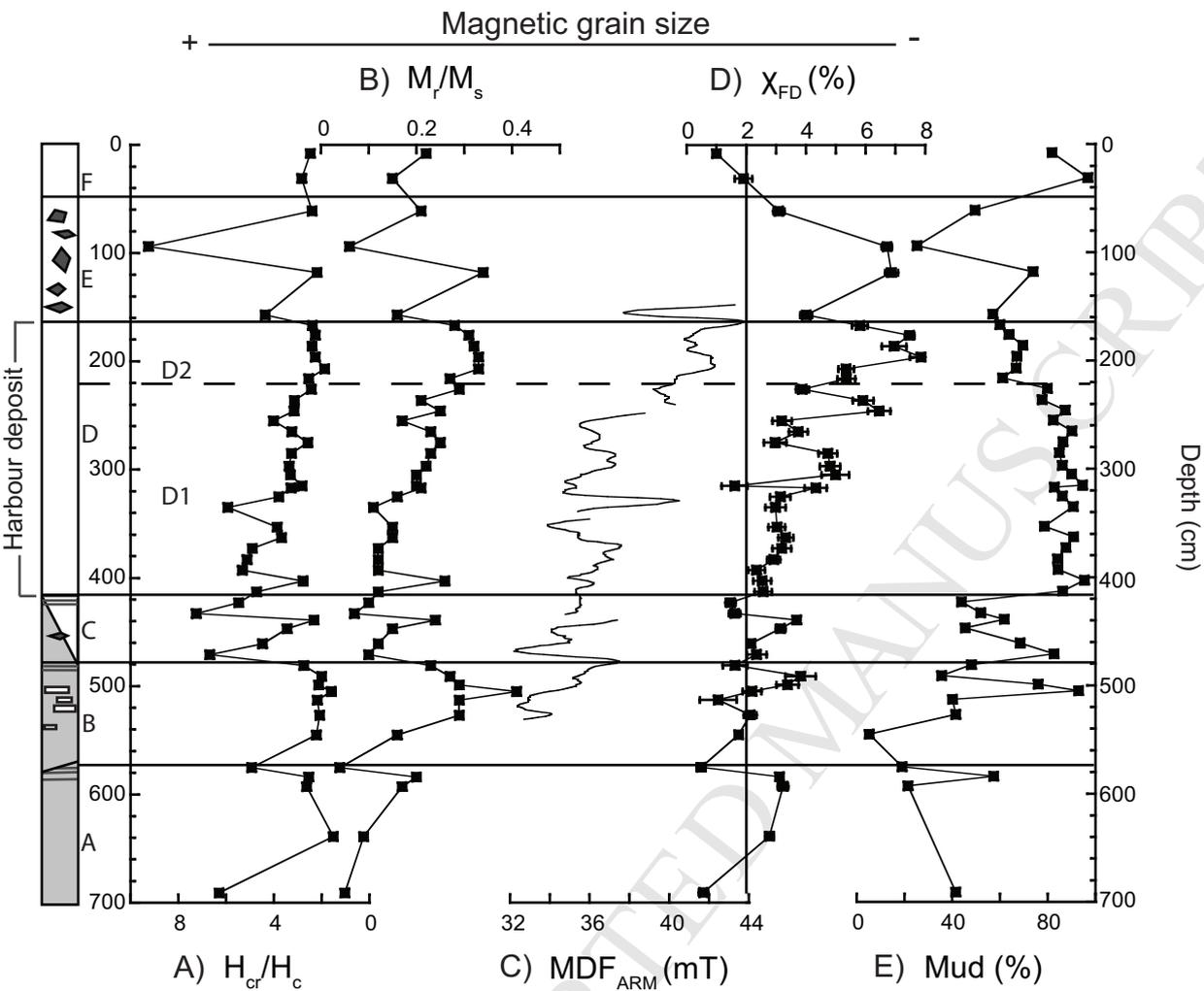




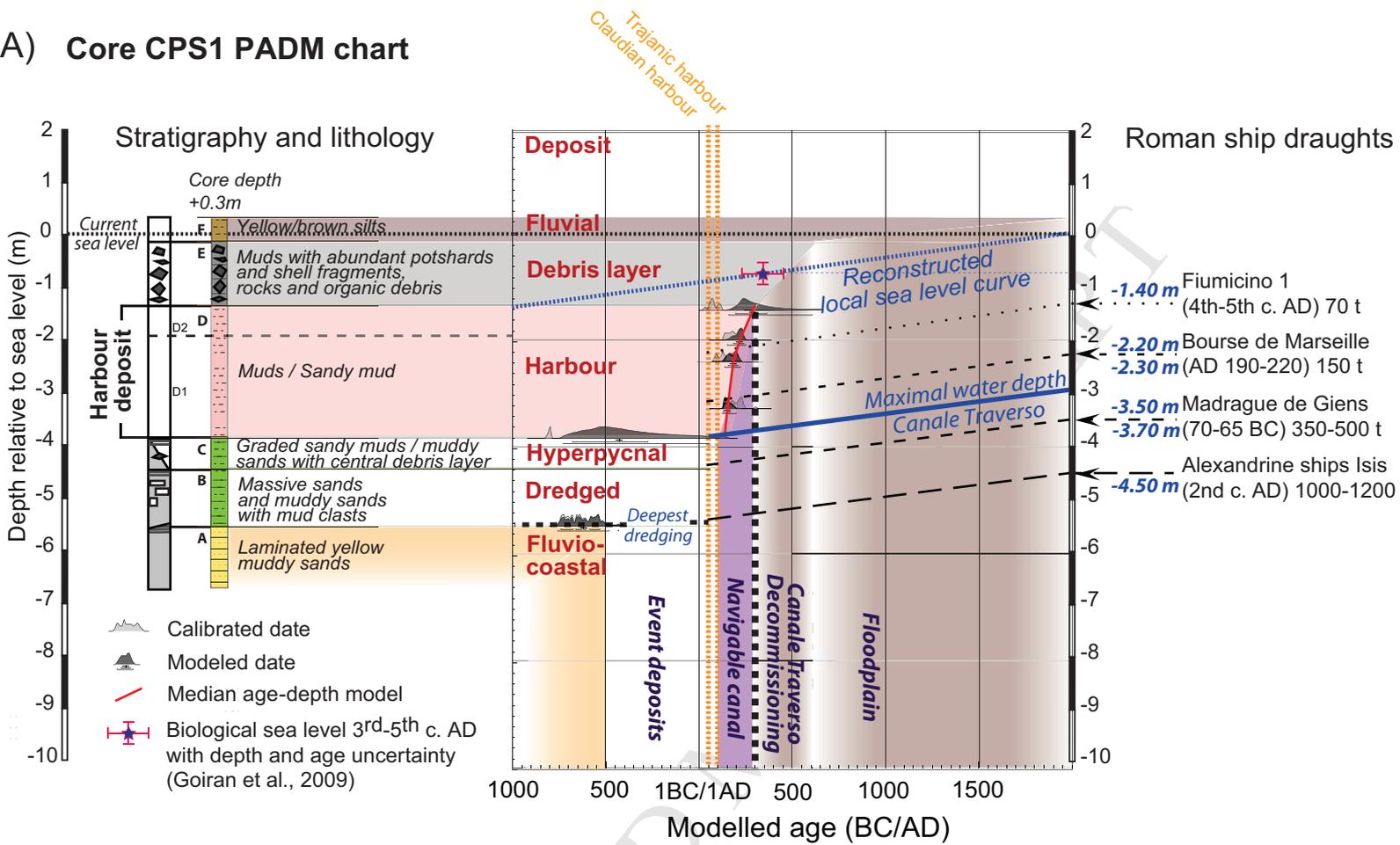




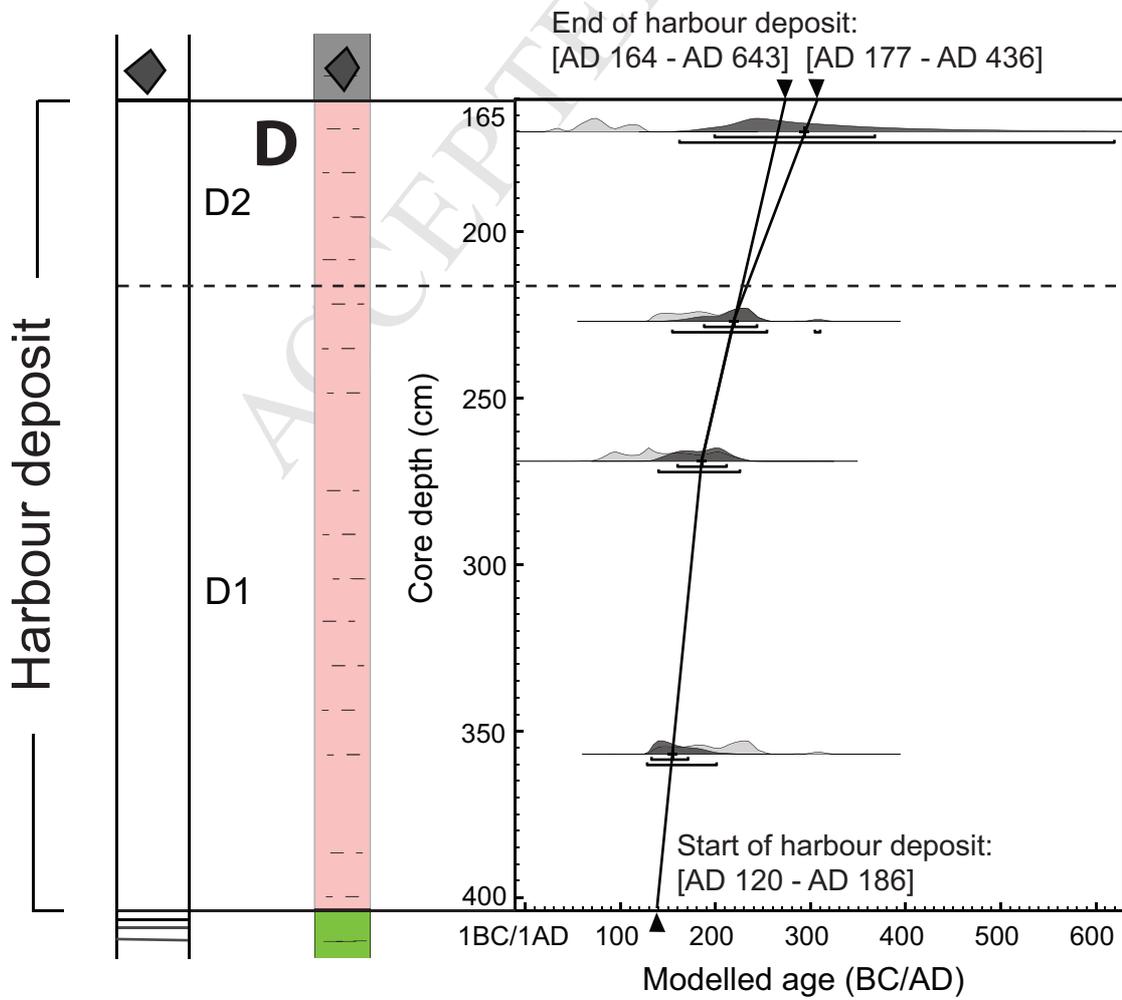


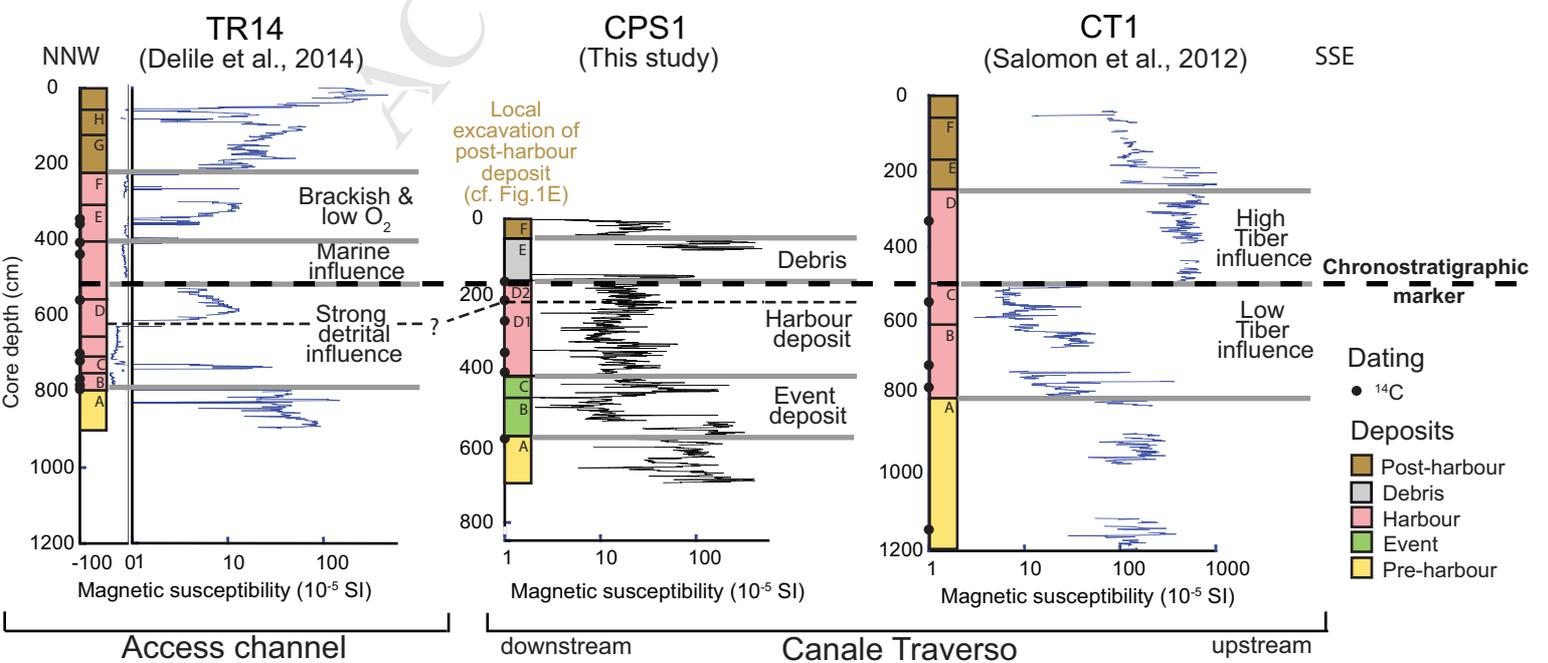


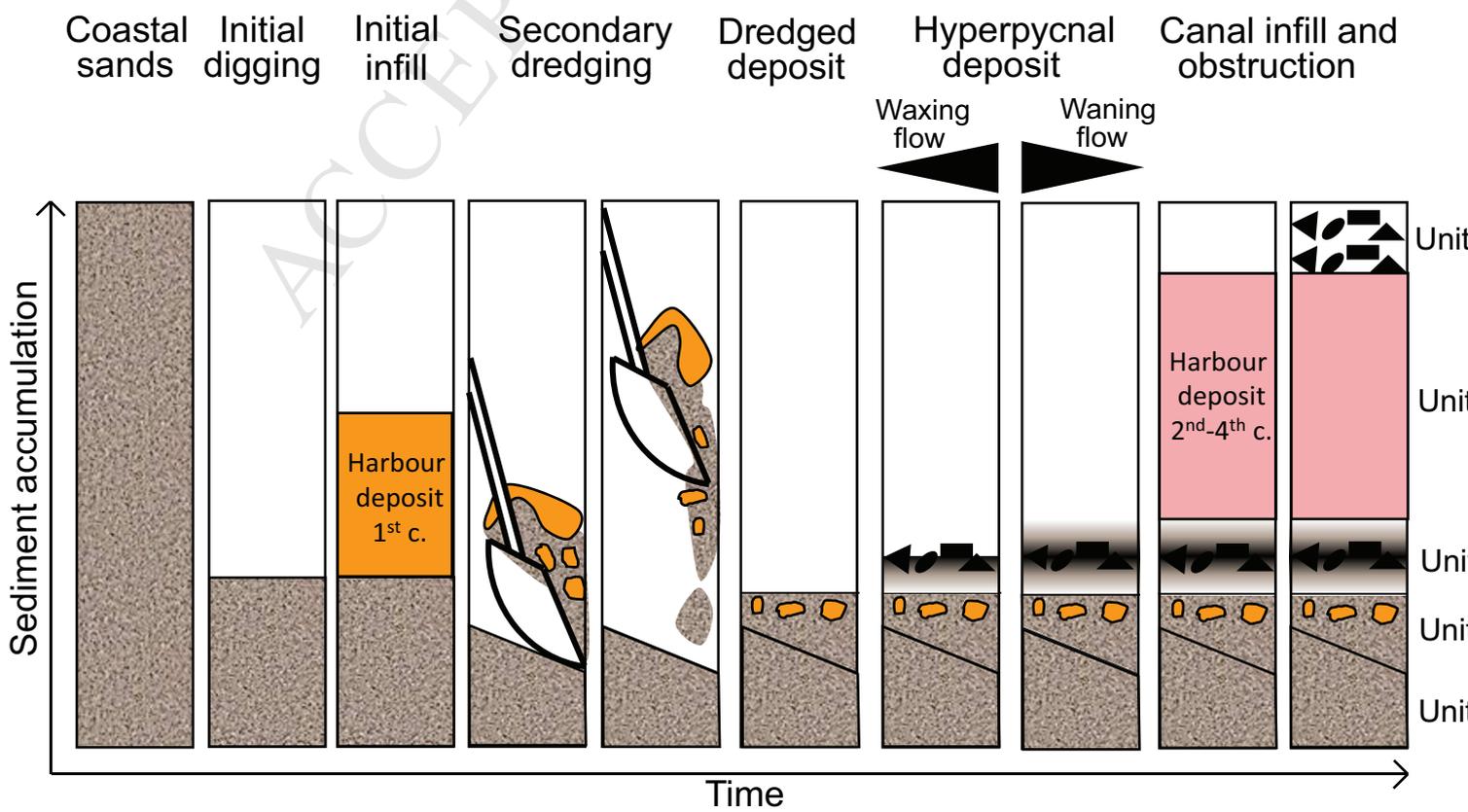
## A) Core CPS1 PADM chart



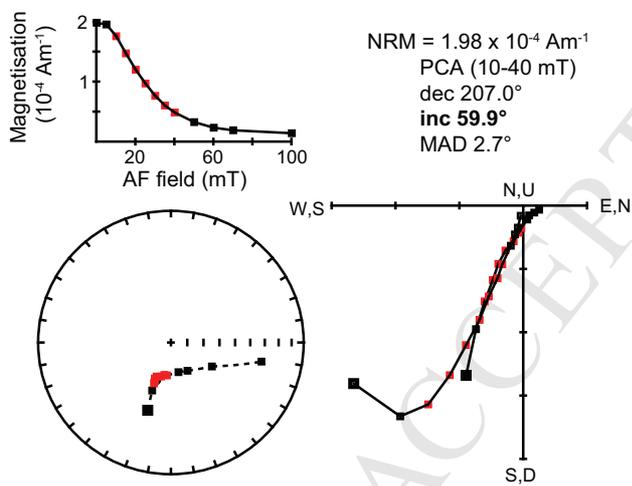
## B)



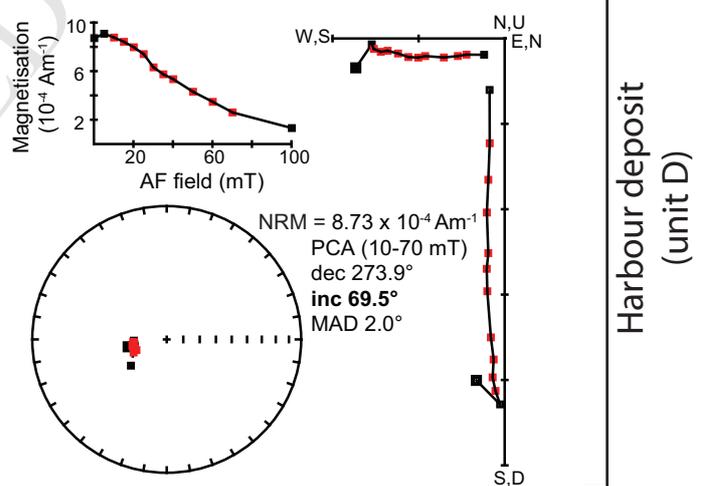




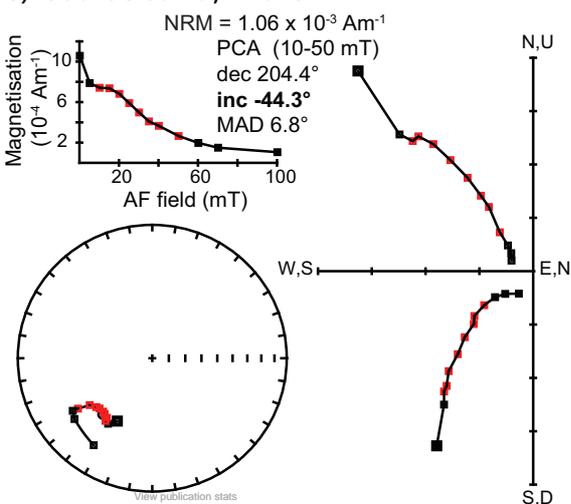
A) Mud sub-unit D2, 190 cm



B) Mud sub-unit D1, 327 cm



C) Coarse sand, 449 cm



D) Sand, 492 cm

