#### UNIVERSITÉ DU QUÉBEC À RIMOUSKI

# LA PHOTORÉACTIVITÉ DE LA MATIÈRE ORGANIQUE DISSOUTE DANS LES EAUX SOUTERRAINES D'UNE ÎLE SUBARCTIQUE

# (PHOTOREACTIVITY OF DISSOLVED ORGANIC MATTER IN GROUNDWATER OF A SUB-ARCTIC ISLAND)

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#### **RÉSUMÉ**

Les eaux souterraines sont des sources potentielles importantes de matière organique dissoute (DOM) pour les eaux côtières. Là, cette matière est soumise à des transformations photochimiques susceptibles d'influencer d'importants processus biogéochimiques marins. Par ailleurs, comme la DOM présente dans les aquifères est généralement peu ou pas exposée à la lumière, elle représente un substrat très intéressant pour étudier la photoréactivité de certaines matières organiques sources, ayant des caractéristiques particulières. Dans cette étude, réalisée sur un aquifère des Îles-de-la-Madeleine dans le Golfe St-Laurent, nous avons recueilli des échantillons DOM à deux stations situées le long d'un transect orienté de la plage vers l'intérieur des terres. Les échantillons ont été caractérisés par spectroscopies d'absorbance et de fluorescence et leur photoréactivité a été comparée. La DOM de l'eau souterraine de la plage (beach groundwater, BGW) était âgée, très colorée et d'origine terrigène. Elle avait un poids moléculaire élevé (MW), une forte humification et un faible contenu en protéines. La DOM de l'eau souterraine plus à l'intérieur des terres (inland groundwater, IGW) contenait du matériel frais, moins coloré et d'origine microbienne. Elle présentait un faible poids moléculaire (MW), un faible degré d'humification et un contenu en protéines élevé. Pour les deux échantillons, BGW et IGW, l'exposition à des radiations solaires simulées a entraîné une augmentation du rapport E<sub>2</sub>/E<sub>3</sub>, de l'indice biologique (BIX) et de la concentration en ammonium (NH<sub>4</sub><sup>+</sup>) et une diminution d'absorbance, du coefficient d'absorption spécifique à 254 nm (SUVα<sub>254</sub>), de l'indice de fluorescence (FI), de l'indice d'humification (HIX), du MW et du carbone organique dissous (DOC). L'irradiation a également réduit les intensités des pics de fluorescence de type humique et des

composants humiques modélisés par le traitement statistique PARAFAC. En se basant perte relative d'absorption, suite à l'irradiation, les changements photochimiquement induits dans BGW étaient comparables à ceux dans IGB pour SUVα<sub>254</sub> et considérablement plus grands pour MW, BIX et DOC, et substantiellement plus petits pour FIX, HIX. Sur la base des photons absorbés, la DOM de BGW était beaucoup plus, modérément plus et moins photoréactive que la DOM de IGW en termes de photoblanchiment d'absorbance, de photominéralisation du DOC et de la photoammonification, respectivement. Les résultats de cette étude, en combinaison avec des travaux antérieurs pour les eaux de surface, conduisent à la formulation d'un paradigme provisoire: la DOM d'origine terrigène est plus sujet au photoblanchiment de l'absorbance, mais moins sujet à la photoammonification que la DOM d'origine microbienne, alors que les deux DOM présentent des réactivités comparables en ce qui concerne la photominéralisation du DOC. Le contenu scientifique de ce memoire a été publié dans Marine Chemistry avec la citation de : Qi L, Xie H, Gagné J-P, Chaillou G, Massicotte P, Yang G-P, 2018. Photoreactivities of two distinct dissolved organic matter Chemistry, pools in groundwater of a subarctic island. Marine https://doi.org/10.1016/j.marchem.2018.03.003.

#### **Abstract**

Groundwater is a potentially significant source of dissolved organic matter (DOM) to coastal oceans where it is subject to photochemical transformation and thus possibly influences major marine biogeochemical processes. Furthermore, groundwater DOM usually receives little prior light exposure, making it suitable for probing the photoreactivity of source organic materials. In this study we collected two DOM pools in beach and inland groundwater of the Îles-de-la-Madeleine in the Gulf of St. Lawrence, characterized them with absorbance and fluorescence spectroscopy, and compared their photoreactivities. Beach groundwater (BGW) primarily comprised old, highly colored terrestrial DOM having high molecular weight (MW), strong humification, and low protein contents, whereas inland groundwater (IGW) largely contained fresh, less colored microbial-derived DOM with low molecular weight (MW), weaker humification, and higher protein contents. For both BGW and IGW, exposure to solar-simulated radiation led to increases in the E<sub>2</sub>/E<sub>3</sub> quotient, biological index (BIX), and ammonium (NH<sub>4</sub><sup>+</sup>) and decreases in absorbance, specific absorption coefficient at 254 nm (SUVα<sub>254</sub>), fluorescence index (FI), humification index (HIX), MW, and dissolved organic carbon (DOC). The irradiation also reduced the intensities of humic-like fluorescence peaks and PARAFAC-modeled humic components. On a fractional-absorption-loss basis, the photochemically induced fractional changes in BGW were comparable to those in IGB for SUVα<sub>245</sub> and E<sub>2</sub>/E<sub>3</sub>, considerably larger for MW, BIX, and DOC, and substantially smaller for FI and HIX. On an absorbed-photon basis, the efficiencies of absorbance photobleaching and DOC photomineralization for DOM in BGW were 8.3 times and 2.0 times those of the respective photoprocesses for DOM in IGW, while the efficiency of photoammonification for DOM in BGW was 41% of that for DOM in IGW. Results from this study, in combination with those reported previously for surface waters, lead to a tentative paradigm: terrigenous DOM is more prone to absorbance photobleaching but less prone to photoammonification than microbial-derived DOM while the two DOM pools are comparably reactive with respect to DOC photomineralization. The scientific content of this thesis has been published in Marine Chemistry with the citation of: Qi L, Xie H, Gagné J-P, Chaillou G, Massicotte P, Yang G-P, 2018. Photoreactivities of two distinct dissolved organic matter pools in groundwater of a subarctic island. Marine Chemistry, https://doi.org/10.1016/j.marchem.2018.03.003.

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**List of Abbreviations and Symbols** 

Abbreviation	Meaning
AQY	Apparent quantum yield
BGW	Beach groundwater
BIX	Biological index
CDOM	Chromophoric dissolved organic matter
DIC	Dissolved inorganic carbon
DOC	Dissolved organic carbon
DOM	Dissolved organic matter
DON	Dissolved organic nitrogen
EEM	Excitation and emission matrix
FI	Fluorescence index
FDOM	Fluorescent DOM
$F_{max}$	Maximum fluorescence intensity
HIX	Humification index
IGW	Inland groundwater
MW	Molecular weight
PARAFAC	Parallel factor
PLA	Percent of light absorbed
R.U.	Raman Units
SGD	Submarine groundwater discharge
UVA	Ultraviolet-A (320-400 nm)
UVB	Ultraviolet-B (290-320 nm)

Symbol	Meaning	Unit
$\overline{A}$	Cross-section area of quartz-windowed	$m^2$
	irradiation cells	
$a_{\text{CDOM}}$	Absorption coefficient of CDOM	$\mathrm{m}^{-1}$
$a_{\mathrm{CDOM,0}}$	$a_{\rm CDOM}$ at irradiation time zero	$\mathrm{m}^{-1}$
$a_{\mathrm{CDOM,t}}$	$a_{\text{CDOM}}$ at irradiation time $t$	$\mathrm{m}^{-1}$
$a_{\mathrm{Fe}}^{ullet}$	Fe(III)-specific absorption coefficient	$\mathrm{L}\mathrm{mg}^{\text{-1}}\mathrm{m}^{\text{-1}}$
	(epierian)	
AFL <sub>250-400</sub>	Average fractional loss of CDOM	dimensionless
	absorption coefficient from 250-400 nm	
$ARF_{250-400}$	Average remaining fraction of CDOM	dimensionless
	absorption coefficient from 250-400 nm	
$a_{\text{tot}}$	Sum of $a_{\text{CDOM}}$ and $a_{\text{w}}$ in irradiation cells	$m^{-1}$
$a_{\text{TOT}}$	Sum of $a_{\text{CDOM}}$ , $a_{\text{w}}$ , and absorption	$m^{-1}$
	coefficient of suspended particles in the	
	groundwater-receiving water body	
$a_{ m w}$	Absorption coefficient of pure water	$\mathrm{m}^{-1}$
$E_2/E_3$	Ratio of absorption coefficient at 250	
	nm to that at 365 nm	
L	Light pathlength of quartz-windowed	m
	irradiation cells	
<u>l</u>	Light pathlength of irradiation tubes	m

$\overline{M_{ m w}}$	Weight-average MW	kDa
$M_{ m n}$	Number-average MW	kDa
n	Number of samples in statistical analysis	dimensionless
N	Dilution factor for DOM discharged	dimensionless
	from groundwater into the receiving	
	seawater body	
$Q_0$	photon flux just below the front window	mol photons m <sup>-2</sup> h <sup>-1</sup> nm <sup>-</sup>
<b>∠</b> ∘	of quartz-windowed irradiation cells	1
$Q_{\mathrm{a}}$	Photons absorbed by CDOM in quartz-	mol photons m <sup>-2</sup> s <sup>-1</sup> nm <sup>-</sup>
£ű	windowed irradiation cells	1
$Q_{ m d0 ext{-}}$	Downwelling solar photon flux just	mol photons m <sup>-2</sup> d <sup>-1</sup> nm <sup>-</sup>
£u0-	below the surface of the groundwater-	1
	receiving body	
$\mathbb{R}^2$	Determination of coefficient	dimensionless
$R_{\rm DOC/CDOM}$	Ratio of DOC loss to AFL <sub>250-400</sub>	dimensionless
$S_{275-295}$	Spectral slope coefficient between 275	$\mu m^{-1}$
5273-293	and 295 nm	MIII
$SUV\alpha_{254}$	Specific UV absorption coefficient	$L mg^{-1} m^{-1}$
DO V 0.254	(decadic) at 254 nm	Ling in
$S_{\lambda}$	Spectral slope coefficient at wavelength	$\mu$ m <sup>-1</sup>
$\mathcal{S}_{\mathcal{K}}$	centered at $\lambda$ within an interval of 21 nm	MIII
Z	Depth of the surface mixed layer or	m
L	photoactive layer in the groundwater-	111
	receiving water body	
λ	Wavelength	nm
$\lambda_{\rm em}$	Fluorescence emission wavelength	nm
$\lambda_{\mathrm{ex}}$	Fluorescence excitation wavelength	nm
$\Lambda$	Action spectra of photobleaching or	m <sup>-1</sup> CDOM m <sup>-2</sup> s <sup>-1</sup> or
71	photoammonification	mol NH <sub>4</sub> m <sup>-2</sup> s <sup>-1</sup>
,,	Mean cosine of photons	dimensionless
μ Φαρανί	AQY of CDOM absorbance	m <sup>-1</sup> CDOM (mol
$\Phi_{ ext{CDOM}}$	photobleaching	photons) <sup>-1</sup>
<del>_</del>	Solar irradiance spectrum-weighted	m <sup>-1</sup> CDOM (mol
$\Phi_{CDOM}$	mean $\Phi_{ ext{CDOM}}$	photons) <sup>-1</sup>
Фрад		mol DOC (mol
$arPhi_{ m DOC}$	AQY of photochemical DOC loss	photons) <sup>-1</sup>
	AOV of NU+ photograduation	· /
$\Phi_{ m NH_4^+}$	AQY of NH <sub>4</sub> <sup>+</sup> photoproduction	mol NH <sub>4</sub> (mol
<del>-</del>	Calan imadianaa anaatmum yyaiahtad	photons) <sup>-1</sup>
$\Phi_{\mathrm{NH_4^+}}$	Solar irradiance spectrum-weighted	mol NH <sub>4</sub> (mol
	mean $\Phi_{ m NH_4^+}$	photons) <sup>-1</sup>
$\rho$	polydispersity	dimensionless
$ au_{CDOM}$	Photochemical turnover time of CDOM	d
	absorbance	
$ au_{\mathrm{DOC}}$	Photochemical turnover time of DOC	d
$ au_{\mathrm{DOM}}$	Photochemical turnover time of DOM	d
$ au_{\mathrm{DON}}$	Photochemical turnover time of DON	d

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#### **Chapter 1. Introduction**

#### 1.1. Dissolved organic matter in submarine groundwater

Submarine groundwater discharge (SGD), which is composed of terrestrial fresh groundwater and recirculated seawater, plays an important role in the ocean's hydrological cycle (Burnett et al., 2001; Moore, 2010). The chemical composition of SGD often differs from that expected from simple mixing of terrestrial and seawater endmembers due to active processes of production, removal, and transformation occurring in coastal aquifers (Beck et al., 2007; Chaillou et al., 2016; Santos et al., 2009). Generally, SGD contains elevated concentrations of carbon (Cai et al., 2003; Goñi and Gardner, 2003; Santos et al., 2009), nutrients (Slomp and Van Cappellen, 2004), and trace metals such as iron (Windom et al., 2006) compared with coastal waters, which has been hypothesized to influence major biogeochemical cycles in marginal seas (Moore, 2010). Published studies, albeit limited in numbers, have demonstrated that SGD is highly enriched with dissolved organic carbon (DOC) and thus could be a potentially important source of dissolved organic matter (DOM) to coastal oceans (Goñi and Gardner, 2003; Santos et al., 2009). The source and/or origin of this DOM is, however, not well known. Marine-derived solutes are considered as the main source of DOM in tidal sands (McLachlan and Brown, 2006; Anschutz et al., 2009; Kim et al., 2012). However, terrestrial input and in situ production can be other important sources of DOM in beach groundwater as recently demonstrated by Couturier et al. (2016).

#### 1.2. Literature review on photochemistry of DOM in groundwater

A large part of the submarine groundwater is delivered directly into shallow water bodies near the shore (Burnett et al., 2003; Johannesson et al., 2011). SGD entering deeper depths may be transported to the surface by buoyancy and physical mixing (e.g. tidal action). DOM, more precisely, its chromophoric component (CDOM), in SGD is thus subject to photooxidation by solar radiation, as is for CDOM in surface water discharged into coastal zones. Exposure of CDOM to sunlight leads to photobleaching, photomineralization, photoammonification, and photoalteration of the bioavailability and chemical structure of DOM including its aromaticity and molecular size (Mopper et al., 2015; Osburn and Morris, 2003; Sharpless and Blough, 2014; Zepp, 2003). CDOM photochemistry thus profoundly impacts the optics and biogeochemical processes in aquatic environments. Despite the recognized ecological and environmental importance of CDOM photochemistry in surface water, little attention has been devoted to understanding the photoreactivity of CDOM in SGD. To our knowledge, there are only two published studies on groundwater DOM photochemistry. Bertilsson et al. (1999) determined photoproduction of dissolved inorganic carbon (DIC), ammonium (NH $_4^+$ ), and certain selected carboxylic acids in two groundwater samples collected from a boreal watershed in northern Sweden. Koopmans and Bronk (2002) measured photoproduction of NH<sub>4</sub> in groundwater taken from coastal Georgia and South Carolina. To date, the effect of solar irradiation on the optical and chemical properties of groundwater DOM remains largely unknown.

Notably, the photoreactivity of CDOM changes with its origin of formation (Aarnos et al., 2012; Brooks et al., 2006) and light exposure history (Powers and Miller, 2015a;

Sharpless et al., 2014; Vähätalo and Wetzel, 2004). In particular, the dose dependence of many CDOM photoprocesses is mostly restricted to interfaces where unexposed CDOM transits from shaded to unshaded environments such as land-water interface and upwelling zones (Vähätalo and Wetzel, 2004; Zhang et al., 2006). Consequently, surface waters downstream of these interfaces may not be suitable for studying photochemistry of the source material. In this regard, groundwater, which usually experiences little exposure before flowing into open water bodies, offers a good opportunity for examining the dose and source dependences of CDOM photoreactivity.

#### 1.3. Objectives

Using two groundwater CDOM pools of different origins, this project aims to elucidate

- 1) the effect of solar radiation on the optical and chemical characteristics of groundwater CDOM,
- 2) the influence of the source of CDOM on its photochemical behavior in terms of photobleaching, photomineralization, and photoammonification, and
- 3) the difference in the photochemical turnover of these two groundwater CDOM pools in the receiving coastal water.

#### Chapter 2. Methods

#### 2.1. Study sites

The sampling site is located on the Martinique Beach of the Îles-de-la-Madeleine in the Gulf of St. Lawrence (Canada) (Figure 1). As part of the Maritimes Permo-Carboniferous Basin in the Northeastern Appalachian Geological Province (Brisebois, 1981), the Îles-de-la-Madeleine are an archipelago encompassing approximately 15 islands, with seven of them joined as tombolos. The core of the main island is composed of interbedded volcanic and volcanoclastic rocks flanked by sandstones of the Cap-aux-Meules formation developed during the Permian Inferior period (Brisebois, 1981). This sandstone formation forms the main aquifer in the archipelago that features high pumping rates due to elevated transmissivities (mean: ~4 × 10<sup>-3</sup> m<sup>2</sup> s<sup>-1</sup>) and recharge rates (mean: ~230 mm yr<sup>-1</sup>, i.e. ~30% of the annual snow and rain precipitation) (Leblanc, 1994; Madelin'Eau, 2004). Groundwater flows from the central volcanic core recharge zone to the unconfined sandstone aquifer and finally to the sea in the form of surficial and submarine discharges.

The archipelago's coasts are mostly fine-grained (~300 μm) sandy barrier systems resulting from the erosion of Holocene deposits and eolian transport. Sediments mainly comprise silicate fragments while the underlying Permian sandstone aquifer contains finer red-orange sands (grain size: ~100 μm) composed of silicate and aluminosilicate with iron-coated silicate grains (Chaillou et al., 2014). At the top of the sandstone aquifer, there is a fragmented organic-rich horizon, with a total organic carbon content of ~20% weight percent. This buried horizon is mainly composed of terrestrial plant detritus that was formed ~900 years B.P., as revealed by <sup>14</sup>C age dating (Juneau, 2012) and is a

geological evidence of local and regional submergences over the last millennia (Grant, 1970; Juneau, 2012; Scott et al., 1995). The sandy Martinique Beach acts as a shallow unconfined aquifer at the shoreline and releases diffuse fresh inland groundwater to the coastal water. Within the beach, fresh and cold groundwater flows towards the seaward discharge region below a narrow intruding saline circulation cell located near the top of intertidal sediments (Chaillou et al., 2016). The discharge zone is not well characterized but likely has a narrow zone of ~100 m from the beach face (Chaillou et al., submitted) where fresh SGD fluxes were estimated around 0.03 m<sup>3</sup> s<sup>-1</sup> based on Darcy's Law (Chaillou et al., 2016). This same study revealed a short groundwater residence time along the subterranean estuary (~32 d), low oxygen conditions (~20% saturation), and a fresh groundwater-born DOC flux of 27 kg d<sup>-1</sup> at the beach face.

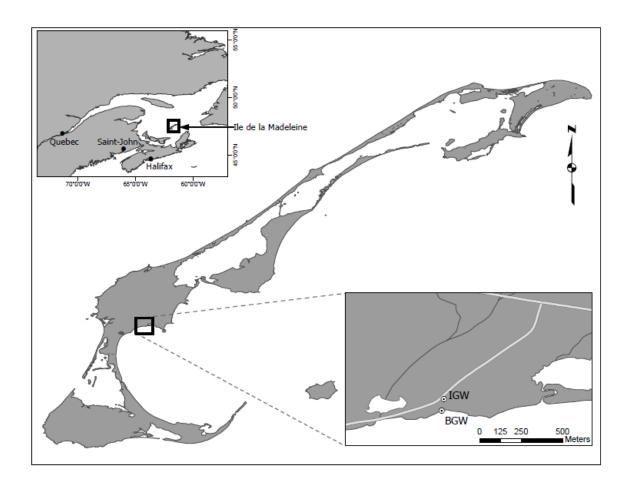


Figure 1. Map of sampling sites.

#### 2.2. Sample collection

Sampling was carried out on the Martinique Beach during the low spring tide period in June 2012, when the unconfined aquifer was supposedly recharged with freshly thawed snow pack. Beach groundwater (BGW) was collected from a site (47.32°N, -65.98°W) ~15 m from the high tide mark (Figure 1) using a multi-level porewater sampler as described by Chaillou et al. (2015). The sampler was installed at least two days before sampling to allow sediments around the sampler to regain equilibrium. Groundwater was pumped from 80 cm below the beach surface into an in-line flow cell fitted with a calibrated YSI 600QS multi-parametric probe to monitor in situ temperature, dissolved oxygen concentration, and salinity. After these parameters had stabilized, groundwater was passed through a Pall AcroPak® 1000 filtration capsule (0.2 µm poresize polyethersulfone membrane) and collected into two 4-L clear-glass bottles wrapped with black plastic film to avoid light exposure. Because of the relatively high pumping flow rate (~1 L min<sup>-1</sup>), the collected water was thought to cover all depths that were shallower than 80 cm. In addition to BGW, inland groundwater (IGW) was taken from a 5-m deep residential water well, ~50 m landward from the BGW sampling site (Figure 1), by pumping water (5.0 m deep) to the surface through silicone tubing. As for BGW, IGW was collected only after physiochemical parameters had stabilized, and was filtered and stored in the same manner as for BGW. The sample bottles, closed with PTFE-lined screw caps, were transported to the home laboratory in Rimouski under dark and cool conditions. Before use, the glass bottles were soaked with 10% HCl overnight, thoroughly rinsed with Nanopure water, air-dried, and then combusted at 450 °C for 5 h.

The filtration capsules were flushed with 15 L Nanopure water and conditioned with 2 L sample water.

#### 2.3. Sample irradiation

Immediately before irradiation, sample water was re-filtered through 0.2- $\mu$ m polyethersulfone membrane filters and transferred into 93-mL cylindrical quartz tubes (i.d.: 2.2 cm; length: 24.5 cm). The tubes were closed with ground-glass stoppers without headspace except where noted. The tubes were horizontally placed in a temperature-controlled water bath (15  $\pm$  1 °C), submerged 0.5 cm below the water surface and irradiated in time series under full-spectrum radiation using a Suntest XLS+ solar simulator equipped with a 1.5 kW xenon lamp. A special UV glass filter was installed to remove UV radiation with wavelengths shorter than 290 nm.

IGW and BGW were irradiated over time courses of 140 h and 474 h, respectively, to examine the temporal trends of photobleaching and DOC photomineralization, and of 96 h and 120 h, respectively, to determine the kinetics of NH<sub>4</sub><sup>+</sup> photoproduction. Duplicate tubes were sacrificed periodically for analyses described in section 2.4. The choice of irradiation time depended on the initial CDOM abundance in the sample (Table 3) and the rate of the photoprocess of interest, with longer irradiations performed for the more colored water and the slower processes. For the 474-h time-course exposure of the highly colored BGW (Table 1), the irradiation tubes were left with 31-mL lab air and gently shaken at each sampling time to mitigate oxygen depletion by photochemical oxygen consumption (Andrews et al, 2000) and to lessen transport limitations on photochemical kinetics due to severe inner filter effects (Hu et al., 2002). The solar simulator's photon fluxes at the irradiation surface were determined with an OL-754 spectroradiometer fitted

with a 2-inch OL IS-270 integrating sphere and calibrated with an OL 752-10E irradiance standard. The photon fluxes in the UVB (280–320 nm), UVA (320–400 nm), and visible (400–600 nm) were 1.5, 1.2, and 1.4 times those of the noontime clear-sky sunlight measured in May at Rimouski (48.453° N, 68.511° W), Canada (Hong et al., 2014). One day of solar-simulated irradiation thus corresponded to 6.0-d UVB, 4.8-d UVA, and 5.7-d visible irradiation with clear-sky sunlight at the latitude of 48.453°N, assuming 1-d clear-sky irradiation to be equivalent to 0.25-d noontime irradiation (Miller and Zepp, 1995).

Additional irradiations were performed to quantify wavelength-resolved apparent quantum yields (AQYs) of photoammonification ( $\Phi_{NH_4^+}$ ), following the method reported by Xie et al. (2012). Briefly, the 0.2- $\mu$ m re-filtered water was dispensed into quartz-windowed cylindrical cells (i.d.: 2.0 cm; length: 14.0 cm) which were irradiated at a temperature of 15 ± 1 °C using a SUNTEST CPS solar simulator equipped with a 1.5-kW xenon lamp. Six light treatments were obtained with successive Schott long band-pass cutoff filters: WG280, WG295, WG305, WG345, WG395, and WG435. Light exposure lasted for 12 h for both IGW and BGW. Spectral photon fluxes under each cutoff filter were measured with the OL-754 spectroradiometer. The absorbance photobleaching data obtained from these same irradiations were used to derive spectrally resolved AQYs of CDOM photobleaching ( $\Phi_{CDOM}$ ). Parallel dark controls accompanied all light treatments and showed negligible changes in the measured parameters relative to the initial samples

**Table 1.** Description of optical indices and EEM peaks used in the text.  $\lambda_{ex}$ : excitation wavelength;  $\lambda_{em}$ : emission wavelength.

Description of optical indices ar	id EEM peaks used in the text. $\lambda_{ex}$ : excita	tion wavelength, $\lambda_{em}$ . emission wavelength	•
Absorbance-based index	Calculation	Function	Reference
$a_{\text{CDOM}}(330)$ : absorption	2.303 times absorbance at 330 nm	Indicator of CDOM content; other	Blough and Del Vecchio,
coefficient at 330 nm (m <sup>-1</sup> )	divided by the pathlength of the	wavelengths such 325 nm, 350 nm, 375	(2002); Nelson et al. (2013);
	spectrophotometer's cuvette	nm, etc. have also been adopted	Osburn et al. (2009a)
$SUV \alpha_{254}$ : specific $UV$	Decadic (i.e. base 10) absorption	Indicator of aromatic carbon content of	Weishaar et al., 2003
absorption coefficient (L mg <sup>-1</sup>	coefficient divided by DOC	CDOM, with higher values indicating	
$m^{-1}$ )	concentration in mg L <sup>-1</sup>	higher aromatic carbon contents	
$S_{275-295}$ : spectral slope between	Exponent of nonlinear fit of an	Proxies of average molecular weight	Fichot and Benner (2012); Helms
275 nm and 295 nm (nm <sup>-1</sup> )	exponential function to the absorption	and lignin content of CDOM, with	et al. (2008)
	spectrum over the wavelength range	higher values indicating lower	
	from 275 nm to 295 nm	molecular weights and lower lignin	
		contents	
$E_2/E_3$ : the E2 to E3 quotient	Ratio of $a_{\text{CDOM}}(250)$ to $a_{\text{CDOM}}(365)$	Proxies of aromaticity and average	Peuravuori and Pihlaja (1997)
		molecular weight of CDOM, with	
		higher values indicating lower	
		molecular weights and lower	
		aromaticity	
Fluorescence-based index	Calculation	Function	Reference
FI: fluorescence index	Ratio of emission intensity at 450 nm	Indicator of relative contribution of	McKnight et al. (2001)
(dimensionless)	to that at 500 nm with excitation at	terrestrial and microbial DOM sources,	
	370 nm	with higher values indicating higher	
		contributions of microbial-derived	
		DOM	
BIX: biological index	Ratio of emission intensity at 380 nm	Indicator of relative contribution of	Huguet et al. (2009)
(dimensionless)	to that at 430 nm with excitation at	autochthonous DOM, with higher	
	310 nm	values indicating high contributions of	
		1 1 1 507	
		recently produced DOM of	
THY 1 10° 1		autochthonous origin	QL (2002)
HIX: humification index	Ratio of emission intensity integrated	autochthonous origin Indicator of humic substance content	Ohno (2002)
HIX: humification index (dimensionless)	over 435-480 nm to that over 300-	autochthonous origin Indicator of humic substance content and degree of humification, with higher	Ohno (2002)
		autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of	Ohno (2002)
(dimensionless)	over 435-480 nm to that over 300- 345 nm with excitation at 254 nm	autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of humification	
(dimensionless)  EEM peak name	over 435-480 nm to that over 300- 345 nm with excitation at 254 nm Wavelength	autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of humification Component	Reference
(dimensionless)  EEM peak name A	over 435-480 nm to that over 300- 345 nm with excitation at 254 nm  Wavelength $\lambda_{ex} = 260 \text{ nm}; \ \lambda_{em} = 400-460 \text{ nm}$	autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of humification Component UVC humic-like	Reference Coble (2007)
(dimensionless)  EEM peak name A C	over 435-480 nm to that over 300- 345 nm with excitation at 254 nm  Wavelength $\lambda_{ex} = 260 \text{ nm}; \ \lambda_{em} = 400-460 \text{ nm}$ $\lambda_{ex} = 320-360 \text{ nm}; \ \lambda_{em} = 420-460 \text{ nm}$	autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of humification  Component  UVC humic-like UVA humic-like	Reference Coble (2007) Coble (2007)
(dimensionless)  EEM peak name A C M	over 435-480 nm to that over 300-345 nm with excitation at 254 nm  Wavelength $\lambda_{ex} = 260 \text{ nm}; \ \lambda_{em} = 400-460 \text{ nm}$ $\lambda_{ex} = 320-360 \text{ nm}; \ \lambda_{em} = 420-460 \text{ nm}$ $\lambda_{ex} = 290-310 \text{ nm}; \ \lambda_{em} = 370-410 \text{ nm}$	autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of humification  Component  UVC humic-like  UVA humic-like  UVA humic-like	Reference Coble (2007) Coble (2007) Coble (2007)
(dimensionless)  EEM peak name A C	over 435-480 nm to that over 300- 345 nm with excitation at 254 nm  Wavelength $\lambda_{ex} = 260 \text{ nm}; \ \lambda_{em} = 400-460 \text{ nm}$ $\lambda_{ex} = 320-360 \text{ nm}; \ \lambda_{em} = 420-460 \text{ nm}$	autochthonous origin Indicator of humic substance content and degree of humification, with higher values indicating higher degrees of humification  Component  UVC humic-like UVA humic-like	Reference Coble (2007) Coble (2007)

#### 2.4. Analyses

#### 2.4.1. Optical analyses

CDOM absorbance was scanned at intervals of 1 nm from 200 nm to 800 nm using a Perkin-Elmer lambda-35 dual beam UV-visible spectrometer fitted with a 1- or 5-cm quartz cuvette and referenced to Nanopure water. The Napierian absorption coefficient,  $a_{\text{CDOM}}(\lambda)$  (m<sup>-1</sup>), where  $\lambda$  is wavelength in nanometers, was calculated as 2.303 times the absorbance divided by the pathlength of the cuvette in meters. A baseline correction was made by subtracting the average absorbance between 683 nm and 687 nm from all spectral values (Babin et al., 2003). The decadic specific absorption coefficient at 254 nm (SUV  $\alpha_{254}$ ), the spectral slope coefficient between 275 nm and 295 nm ( $S_{275-295}$ ), and the  $E_2/E_3$  quotient are calculated from the measured absorption coefficients following the procedures described in Table 1. Furthermore, the wavelength distribution of spectral slope, also termed the spectral slope curve ( $S_{\lambda}$ ), between 250 nm and 550 nm was analyzed at 21 nm intervals with a 1 nm step according to the method of Loiselle et al. (2009).  $S_{\lambda}$  was computed using non-linear exponential fitting and represents the spectral slope at the center wavelength of a given 21-nm wavelength band.

Fluorescence excitation-emission matrices (EEMs) were obtained using a FLUOROMAX-4 spectrofluorometer to characterize the composition of CDOM (Coble, 1996; Boehme et al., 2004). The bandwidths were set to 5 nm for excitation and 2 nm for emission. A series of emission scans between 280 nm and 560 nm were collected over excitation wavelengths between 240 nm and 500 nm at 5-nm increments. Excitation wavelengths below 240 nm and emission wavelengths below 300 nm were removed due to low signal to noise ratios. Before optical analysis, the samples were allowed to warm

to room temperature. Nanopure water was used as blank to remove most of the first- and second-order Raman scattering. The spectra of all measurements were normalized to Raman Units (R.U.), employing the Raman Scatter Peak correction reported by Lawaetz and Stedmon (2009). BGW samples with high absorbance were diluted with Nanopure water to  $a_{\rm CDOM}(280) < 11.3~{\rm m}^{-1}$  to eliminate the inner-filtering effect (Ohno, 2002). Major peaks in the EEMs were identified using the wavelength ranges reported by Coble (2007), and fluorescence (FI), biological (BIX), and humification (HIX) indices are computed from the corrected EEMs per protocols published in the literature (Table 1). Note that Table 1 also lists the biogeochemical functions of the optical indices employed in this study.

Parallel factor analysis (PARAFAC) was performed to decompose the EEMs into a set of underlying fluorescent components (Stedmon et al., 2003; Stedmon and Bro, 2008). To reduce the dominance of high fluorescence intensity signals, the EEMs were first scaled to a unit of variance within the sample mode to construct the calibration model (Bro, 1997). PARAFAC models from 2 to 7 components with constraints of nonnegativity in all modes were successively conducted with MATLAB (version 2008b; MathWorks 2008) using DOM Fluorescence Toolbox (DOM Fluor version 1.6) and validated as described by Stedmon and Bro (2008). The parameters obtained from the PARAFAC model were used to calculate an approximate abundance of each component, expressed as F<sub>max</sub> in R.U., which corresponds to the maximum fluorescence intensity for a particular sample.

#### 2.4.2. Chemical analyses

The molecular weight (MW) of CDOM was determined using high-pressure size exclusion chromatography (HPSEC) according to the protocol method reported by Chin et al. (1994) and modified by Lou and Xie (2006). Briefly, samples (100 µL) were injected at controlled temperature (20°C) into a Waters Protein-Pak 125 column carrying a mobile phase of 0.1 mol L<sup>-1</sup> sodium chloride aqueous solution buffered with phosphate to pH 6.8. The absorbance of eluents was detected at 254 nm with a Waters 2487 dual- $\lambda$ absorbance detector. Peak integration was performed using Waters Breeze GPC software. The void volume and total permeation volume of the column were determined using blue dextran and acetone, respectively. The system was calibrated with acetone (MW: 58) and sodium polystyrene sulfonate standards (PPS, Polysciences, MA) with molecular weights of 1.8, 4.6, 8, 18, and 35 kDa. Data from the PPS 1.8 kDa were excluded because it displayed wide multiple peaks, as also reported by Zhou et al. (2000). Calculation of number- and weight-average MW followed the formulae of Chin et al. (1994). The ratio of weight-average MW  $(M_{\rm w})$  to number-average MW  $(M_{\rm n})$  gives polydispersity  $(\rho)$ , a measure of the heterogeneity of CDOM (Chin et al., 1994).

DOC samples were acidified to pH ~2 with 2 N HCl to remove the dissolved inorganic carbon and analyzed in triplicate using a Shimadzu TOC-Vcpn carbon analyzer calibrated with potassium biphthalate. The system was checked, at intervals of seven consecutive sample analyses, against Hansell's low-carbon and deep Florida Strait (700 m) reference waters with DOC concentrations ([DOC]s) of 1  $\mu$ mol L<sup>-1</sup> and 41–44  $\mu$ mol L<sup>-1</sup>, respectively. The coefficient of variation on five replicate injections was <1.5%.

Determination of NH<sub>4</sub><sup>+</sup> concentrations followed the fluorometric method of Holmes et al. (1999). Samples were mixed with a working regent containing *o*-phthaldialdehyde

(OPA), sodium sulfite, and sodium borate. The addition of sodium sulfide minimizes potential interference from amino acids (Holmes et al., 1999). The sample and working reagent mixture was allowed to stay in the dark for 2.5 h before its fluorescence was measured with the FLUOROMAX-4 spectrofluorometer. The procedure was calibrated using freshly prepared ammonium chloride aqueous solutions covering concentrations from 0 to 5.0  $\mu$ mol L<sup>-1</sup>. Corrections were made for the sample's background fluorescence and matrix effect caused by substances such as sea salts and DOM (Holmes et al., 1999). The precision of NH<sub>4</sub><sup>+</sup> determination was  $\pm$  0.01  $\mu$ mol L<sup>-1</sup>.

Fe(II) and Fe(III) concentrations in the unirradiated samples were determined using the ferrozine method originally reported by Stookey (1970) and modified by Viollier et al. (2000). Briefly, a water sample was added with ferrozine and the absorbance at 562 nm was recorded with the Perkin-Elmer lambda-35 spectrometer. The mixture was then amended with hydroxylamine hydrochloride and allowed to react for 10 min to complete the reduction of Fe(III) to Fe(II). An ammonium acetate buffer was added to the mixture and the absorbance at 562 nm was determined again. The concentrations of Fe(II) and Fe(III) were calculated from equations 3 and 4 in Viollier et al. (2000) using the absorbance values recorded before and after the reduction of Fe(III). The procedure was calibrated with Fe(III) standards prepared with FeCl<sub>3</sub> and diluted in a NaCl solution to match the sample salinity.

#### 2.4.3. Statistical analyses

Statistical analyses were performed using the linear regression software of SigmaPlot for Windows (version 10.0, Systat Software Inc., 2006).

#### 2.5. Retrieval of spectral AQYs

The spectral AQY of NH<sub>4</sub><sup>+</sup> photoproduction ( $\Phi_{NH_4^+}(\lambda)$ ) is defined as the number of moles of NH<sub>4</sub><sup>+</sup> produced per mole of photons absorbed by CDOM at wavelength  $\lambda$ . The photon flux absorbed by CDOM,  $Q_a(\lambda)$  (mol photons m<sup>-2</sup> s<sup>-1</sup> nm<sup>-1</sup>), in an irradiation cell was calculated according to Hu et al. (2002):

$$Q_a(\lambda) = AQ_0(\lambda)a_{\text{CDOM}}(\lambda)(a_{\text{tot}}(\lambda))^{-1}(1 - \exp(-a_{\text{tot}}(\lambda)L))$$
 (1)

where A (m<sup>-2</sup>) and L (m) denote the cross-sectional area and pathlength of the irradiation cells, respectively,  $Q_0(\lambda)$  (mol photons m<sup>-2</sup> s<sup>-1</sup> nm<sup>-1</sup>) signifies the photon flux just below the front window of the cell, and  $a_{tot}(\lambda)$  (m<sup>-1</sup>) is the sum of  $a_{CDOM}(\lambda)$  and the absorption coefficient of pure water at wavelength  $\lambda$  taken from Pope and Fry (1997) and Buiteveld et al. (1994).  $\Phi_{NH_4^+}(\lambda)$  is assumed to decrease with wavelength quasi-exponentially (Xie et al., 2012):

$$\Phi_{NH_4^+}(\lambda) = m_1 \exp(m_2 (\lambda + m_3)^{-1})$$
 (2)

where  $m_1$ ,  $m_2$ , and  $m_3$  are fitting parameters. The amount of produced in an irradiation cell over the exposure period can then be predicted as the product of the assumed function and the number of photons absorbed by CDOM integrated from 290 nm to 600 nm. The optimum values of the unknown parameters in Eq. 2 are obtained using an Matlab-coded iterative curve-fit method (Johannessen and Miller, 2001; Zhang et al., 2006) by varying these parameters from initial estimates until the minimum difference between the measured and predicted production is achieved.

The spectral AQY of CDOM absorbance photobleaching ( $\Phi_{\text{CDOM}}(\lambda)$ ) was calculated as the loss of CDOM absorption coefficient at 330 nm ( $a_{\text{CDOM}}(330)$ ) per mole of photons absorbed by CDOM at wavelength  $\lambda$  (Osburn et al., 2009). In the calculation of  $\Phi_{\text{CDOM}}(\lambda)$ ,  $a_{\text{CDOM}}(330)$  is assigned the units of m<sup>-1</sup> CDOM per m<sup>3</sup> water (Vähätalo and Wetzel, 2004; Osburn et al., 2009a).  $\Phi_{\text{CDOM}}(\lambda)$  was modeled using Eq. 2 in the same manner as  $\Phi_{\text{NH}_4^+}(\lambda)$ . Based on  $R^2$  and the normalized root-mean-square error for the regression of the photochemical rates calculated from the retrieved  $\Phi(\lambda)$  against the measured rates (Koehler et al., 2016), the quasi-exponential function of Eq. 2 performed better than the simple exponential form (e.g. Vähätalo and Wetzel, 2004) for both BGW and IGW, though the difference for BGW photobleaching was marginal (Table 2). The quasi-exponential model was thus chosen for this study. Note that AQYs have units of mol NH<sub>4</sub>+ (mol photons)<sup>-1</sup> for photoammonification and m<sup>-1</sup> CDOM (mol photons)<sup>-1</sup> for photobleaching. For brevity, these units are omitted herein.

Table 2.

Comparison of the simple exponential and quasi-exponential models for deriving apparent quantum yield spectra of absorbance photobleaching and photoammonification. nRMSE and R<sup>2</sup> are normalized root-mean-square error and coefficient of determination, respectively, for the regression of modeled against measured rates. Simple exponential model is signified by  $\Phi(\lambda) = m_1 \exp(-m_2(\lambda - 290))$  and quasi-exponential model by  $\Phi(\lambda) = m_1 \exp(m_2(\lambda + m_3)^{-1})$ , where  $\Phi(\lambda)$  denotes apparent quantum yield at wavelength  $\lambda$  (nm), 290 is the reference wavelength in nanometers, and  $m_1$ ,  $m_2$ , and  $m_3$  are fitting parameters.

are money parameters.										
	Photoblea	ching			Photoammonification					
	BGW		IGW		BGW		IGW			
Model	$\mathbb{R}^2$	nRMSE	$\mathbb{R}^2$	nRMSE	$\mathbb{R}^2$	nRMSE	$\mathbb{R}^2$	nRMSE		
		(%)		(%)		(%)		(%)		
Simple exponential	0.986	9.2	0.843	17.1	0.983	13.3	0.982	11.2		
Quasi- exponential	0.988	8.2	0.978	6.2	0.997	5.4	0.993	6.7		

To evaluate the relative contributions of different spectral regimes to photobleaching and photoammonification under natural solar radiation, action spectra ( $\Lambda(\lambda)$ ) of these two photoprocesses are calculated according to Song et al. (2013):

$$\Lambda(\lambda) = Q(\lambda)\Phi(\lambda) \tag{3}$$

where  $\Phi(\lambda)$  denotes  $\Phi_{\mathrm{NH}_4^+}(\lambda)$  or  $\Phi_{\mathrm{CDOM}}(\lambda)$  and  $Q(\lambda)$  (mol photons m<sup>-2</sup> s<sup>-1</sup> nm<sup>-1</sup>) is the noontime clear-sky solar photon flux recorded at Rimouski (48.453° N, 68.511° W), Canada, on 24 May 2005.  $\Lambda(\lambda)$  signifies the spectral distribution of the depth-integrated photobleaching (m<sup>-1</sup> CDOM m<sup>-2</sup> s<sup>-1</sup>) or photoammonification (mol NH<sub>4</sub><sup>+</sup> m<sup>-2</sup> s<sup>-1</sup>) rate in an idealized water column in which CDOM is the sole light absorber. Moreover, to facilitate comparison of AQY spectra between different samples, a solar spectrum-weighted mean AQY( $\overline{\Phi}$ ) is defined following Zhang et al. (2006):

$$\overline{\Phi} = \int_{290}^{600} Q(\lambda)\Phi(\lambda)d\lambda \left(\int_{290}^{600} Q(\lambda)d\lambda\right)^{-1} \tag{4}$$

where  $\overline{\Phi}$  denotes  $\overline{\Phi}_{\text{CDOM}}$  or  $\overline{\Phi}_{\text{NH}_4^+}$  with  $\Phi(\lambda)$  corresponding to  $\Phi_{\text{NH}_4^+}(\lambda)$  or  $\Phi_{\text{CDOM}}(\lambda)$ . This normalization reduces the full AQY spectrum to a single value equivalent to the loss of  $a_{\text{CDOM}}$  or the production of NH<sub>4</sub><sup>+</sup> divided by the amount of photons absorbed by CDOM in the idealized water column delineated above.  $\overline{\Phi}$  accounts for both the magnitudes and shapes of the AQY and solar irradiance spectra, thereby giving more weights to the wavelengths at which the values of  $\Lambda(\lambda)$  are elevated.

#### **Chapter 3. Results and Discussion**

For brevity, sample names for a property will be added, whereby needed, as a superscript to the right side of the symbol or abbreviation representing that property. For example,  $SUV\alpha_{254}^{BGW}$  denotes  $SUV\alpha_{254}$  for BGW, while  $SUV\alpha_{254}^{IGW}$  stands for  $SUV\alpha_{254}$  for IGW. It is further stipulated that each symbol and abbreviation, where applicable, can be used as both a singular and a plural form.

#### 3.1. Initial sample properties and sources of DOM

#### 3.1.1. Chemical properties

Table 3 summarizes results from chemical and optical analyses before irradiation. IGW was essentially freshwater (salinity: 0.31) but BGW contained significant amounts of salts (salinity: 2.82) due to entrainment of recirculated seawater (Chaillou et al., 2014; Couturier et al., 2016). In situ temperatures of the two samples (~11-12 °C) were similar but BGW contained much lower dissolved oxygen (~15% saturation) than did IGB (~64% saturation). The pH<sup>IGW</sup> (6.86) was marginally below the neutral level, whilst pH<sup>BGW</sup> (6.29) was more acidic, consistent with the reducing condition of BGW as reflected by its lower dissolved oxygen level. Total dissolved Fe in BGW (332.7 μmol L<sup>-1</sup>) was 261 times higher than in IGW (1.3 μmol L<sup>-1</sup>), with Fe (III) accounting for 79% and 95% of the total Fe in BGW and IGW, respectively. The far higher total Fe in BGW was due apparently to its reducing condition combined with a highly elevated DOC load (see below), since the solubility of iron is higher under reducing conditions and Fe complexes with DOM (Deflandre et al., 2002).

Table 3.

Physical, chemical, and optical properties of beach groundwater (BGW) and inland groundwater (IGW) before irradiation. Dissolved oxygen (O<sub>2</sub>), temperature (T), pH, and salinity were measured in situ. Numbers in parentheses are relative deviations of duplicate samples (%).

Sample	$O_2 \atop (\mu mol \ L^{\text{-}1})$	T (°C)	pН	Salinity	Total Fe (µmol L <sup>-1</sup> )	DOC (µmol L <sup>-1</sup> )	$a_{\text{CDOM}}(330)$ (m <sup>-1</sup> )	$S_{275-295}$ ( $\mu m^{-1}$ )	$E_2/E_3$	SUV $\alpha_{254}$ (L mg <sup>-1</sup> m <sup>-1</sup> )	M <sub>w</sub> (kDa)	M <sub>n</sub> (kDa)	ρ	FIX	BIX	HIX
BGW	50.0	12.4	6.29	2.82	332.7	6094	389.2	11.8	3.94	5.39	2.2	0.56	5.67	1.20	0.38	29.9
DGW 30.0 12.4 0.29	0.29	3.29 2.02	(0.2)	(0.8)	(0.07)	(2.6)	(0.2)	3.39	3.2	0.50	3.07	1.20	0.36	29.9		
IGW	220.0	11.1	6.86	0.21	1.3	196	4.36	15.4	7.93	2.85	1.0	0.24	2.97	1.61	0.70	6.65
IGW	220.0	11.1	0.80	0.31	(1.5)	(1.6)	(4.0)	(0.6)	(0.4)	2.83	1.0	0.34	2.97	(0.8)	(0.2)	(0.1)

BGW (6094  $\mu$ mol L<sup>-1</sup>) exceeded IGW (196  $\mu$ mol L<sup>-1</sup>) in DOC concentration by a factor of 31. Similarly, DOC<sup>BGW</sup> was 43 times higher than that in the seawater endmember (140  $\mu$ mol L<sup>-1</sup>; Courtier et al., 2016), implying a localized source of DOM in BGW. The extremely high DOC stock in BGW has been ascribed to DOC leached from the paleosol horizon overlying the sandstone aquifer in the beach area (Chaillou et al., 2014, 2015; Courtier et al., 2016). This old-age, forest-soil horizon is carbon-rich (20% total organic carbon by weight; Couturier et al., 2016) and was <sup>14</sup>C-dated to ~900 BP (Juneau, 2012).  $M_W^{BGW}$  (3.2 kDa) was more than triple  $M_W^{IGW}$  (1.0 kDa).  $M_W^{BGW}$  falls in the  $M_W$  range between Suwanee River fulvic (1.39 kDa) and humic (3.94 kDa) acids (Lou and Xie, 2006) but is closer to the latter;  $M_W^{IGW}$  is, nonetheless, below the value of Suwanee River fulvic acid and at the boundary of low- and high-MW DOM. Hence, humic acids were likely the main component of DOM<sup>BGW</sup>, while light fulvic acids were the dominant fraction of DOM<sup>1GW</sup>. Moreover,  $\rho^{BGW}$  (5.67) was nearly two times higher than  $\rho^{IGW}$  (2.97), suggesting a higher heterogeneity of DOM<sup>BGW</sup> relative to DOM<sup>1GW</sup>.

## 3.1.2. Absorption properties

Compared with DOC, the difference in  $a_{\text{CDOM}}$  between the two samples was even larger, reaching a factor of ~89 at 330 nm (Table 3). As a result, SUV $\alpha_{254\_BGW}$  (5.39 L mg<sup>-1</sup> m<sup>-1</sup>) was nearly twice SUV $\alpha_{254}^{\text{IGW}}$  (2.85 L mg<sup>-1</sup> m<sup>-1</sup>), suggesting a much richer aromaticity of CDOM<sup>BGW</sup> (Ghernaout, 2014; Weishaar et al., 2003). Although  $a_{\text{CDOM}}$  for both samples decreased about exponentially with wavelength, the absorption spectrum of IGW was markedly steeper than that of BGW (Figure 2A) and  $S_{\lambda}^{\text{IGW}}$  was higher than  $S_{\lambda}^{\text{BGW}}$  (Figure 2B). Both spectral slope curves possessed a broad elevation from ~300 nm

to ~440 nm superimposed with smaller peaks but the elevation was less pronounced for BGW (Figure 2B), again implying a greater proportion of humic acids in BGW than in IGW (Loiselle et al., 2009). The percent difference between the two curves varied greatly with wavelength, exhibiting a deep trough centered at ~287 nm, a double-peak elevation from 320 nm to 367 nm, and lower values in the visible (Figure 2B). S<sub>275-295</sub> (centered at 285 nm in Figure 2B) for IGW (15.4  $\mu$ m<sup>-1</sup>) surpassed that for BGW (11.8  $\mu$ m<sup>-1</sup>) by 30.5% but the difference was among the smallest in the UV regime. The commonly adopted wavelength band of 275-295 nm is thus not optimal for describing the spectral difference of our samples. The steeper slope of IGW also led to E<sub>2</sub>/E<sub>3</sub><sup>IGW</sup> (7.93) being twofold higher than  $E_2/E_3^{BGW}$  (3.94). The lower values of  $S_{275-295}$  and  $E_2/E_3$  for BGW are consistent with the higher MwBGW described above. The E2/E3 values here were used to estimate  $M_{\rm w}$  based on the empirical relationship between  $M_{\rm w}$  and  $E_2/E_3$ ,  $M_{\rm w}$  =  $0.315*\exp(4.96/(-1.72+E_2/E_3))$ , established by Lou and Xie (2006), giving  $M_W^{BGW}$  of 2.93 kDa and  $M_{\rm W}^{\rm IGW}$  of 0.70 kDa. These values are close to the measured ones (3.2 kDa for BGW and 1.0 kDa for IGW), validating E<sub>2</sub>/E<sub>3</sub> as an indicator of MW in this study.

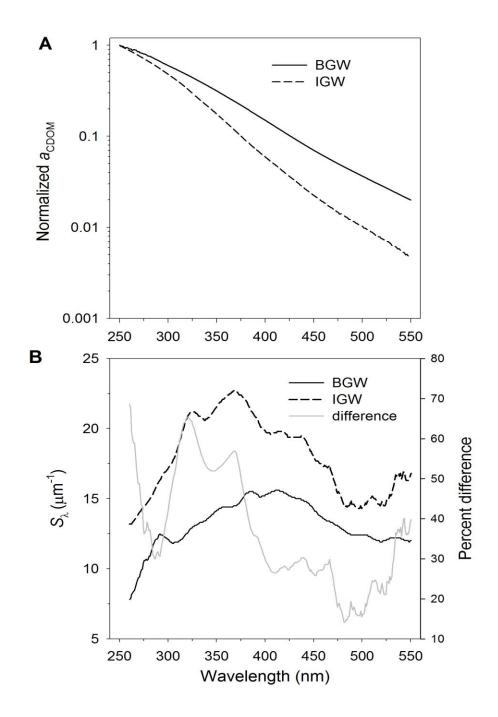


Figure 2. Normalized CDOM absorption coefficient spectra (A) and spectral slope curves (B) for original (unirradiated) samples. Percent difference in panel B is defined as (IGW - BGW) / IGW \*100.

## **3.1.3.** Fluorescence properties

FI<sup>BGW</sup> (1.20) and BIX<sup>BGW</sup> (0.38) were considerably lower than FI<sup>IGW</sup> (1.78) and BIX<sup>IGW</sup> (0.70), while HIX<sup>BGW</sup> (29.9) exceeded HIX<sup>IGW</sup> (6.65) by 4.5 times. In general, FI values of less than 1.4 point to DOM of terrestrial origin and values of 1.9 or higher represent DOM of microbial origin (Birdwell and Engel, 2010; McKnight et al., 2001). BIX values ranging from 0.8 to 1.0 indicate fresh, microbially derived DOM, while values below 0.6 signify little autochthonous material (Huguet et al., 2009). Fresh DOM derived from plant biomass usually displays HIX values of <5 (Sazawa et al., 2011), whereas soil-derived DOM has values between 10 and 30 (Wickland et al., 2007). Clearly, the fluorescence indices of BGW are in line with the values for old, soil-derived DOM, while those of IGW are at the borderline between CDOM of terrestrial and microbial origins but closer to the latter.

The EMMs of BGW and IGW both displayed two marked fluorescence peaks (Figure 3A,B), corresponding to peaks A and C defined by Coble (2007) (Table 1). The two peaks were well separated for IGW but largely connected together for BGW. Peak A for IGW blue-shifted by 10 nm in excitation and 16 nm in emission relative to peak A for BGW. These blue shifts imply lower degrees of substitution and conjugation of aromatic moieties in IGW (Stedmon and Nelson 2015), in accordance with the lower MW and SUVα<sub>254</sub> values for IGW. The intensities of peaks A and C for BGW (33.27 R.U. and 26.20 R.U.) were 59 and 71 times those for IGW (0.56 R.U. and 0.37 R.U.). The ratio of peak C to A for BGW (0.79) was 20% higher than that for IGW (0.66), indicating a higher proportion of humic acid-like *vs.* fulvic acid-like material in BGW compared with IGW (Chen et al., 2003; Hansen et al., 2016). In addition to the evident peaks of A and C,

IGW also exhibited relatively weak fluorescence signals at  $\lambda_{ex}/\lambda_{em}$  240-270 nm/330-380 nm that were partially overlapped by the periphery of peak A. This EEM region, albeit outside peak B and T defined by Coble (2007) (Table 1, Figure 3), falls into EEM Region IV corresponding to soluble microbial by-product-like material (Chen et al., 2003). Residual and split-half PARAFAC analyses (Stedmon and Bro 2008) demonstrated that five distinct components, C1–C5 (Figure 4), explained 99.8% of the variance and thus adequately modeled different fluorescent DOM (FDOM) profiles in the dataset combining BGW and IGW both before and after irradiation. The maximum fluorescence intensities (F<sub>max</sub>) of C1-C4 in the initial BGW were 21.9 R.U. (C1), 17.2 R.U. (C2), 17.8 R.U. (C3), and 9.6 R.U. (C4), while C5 was undetectable ( $F_{max} = 0$  R.U.). All five components were detected in the initial IGW, with F<sub>max</sub> of 0.23 R.U. (C1), 0.13 R.U. (C2), 0.29 R.U. (C3), 0.30 R.U. (C4), and 0.25 R.U. (C5). The abundances of C1, C2, C3 and C4 were thus 96, 130, 61, and 31 times higher in BGW than in IGW. Based on the percent contribution of each individual component (excluding C5) to the summation of C1-C4 (Figure 15), BGW was enriched with C1 and C2 (particularly C2) but depleted with C3 and C4 (particularly C4) relative to IGW. According to previous studies (Cobble, 1996; Stedmon et al., 2003; Massicotte and Frenette, 2011), C1 and C2 represent terrestrial humic-like FDOM, C3 and C4 autochthonous humic-like FDOM and C5 autochthonous protein-like FDOM. PARAFAC modeling thus suggests that DOM<sup>BGW</sup> originated primarily from terrestrial plants while DOMIGW was derived principally from microbial processes and thus enriched with proteins, although both DOM pools possessed mixed terrestrial and autochthonous signatures.

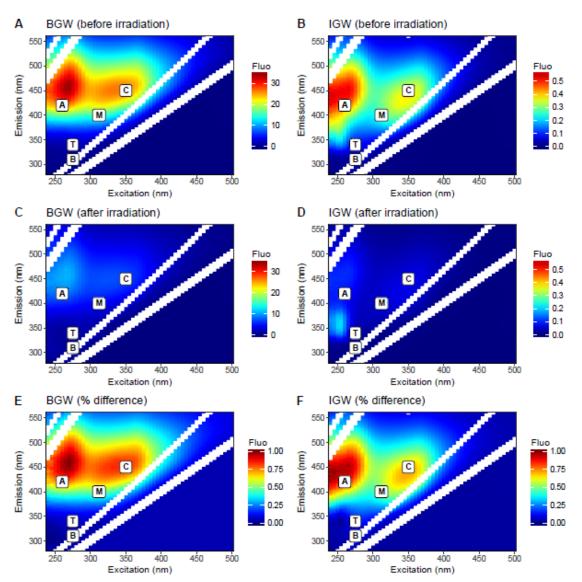


Figure 3. Excitation-emission matrices (EMMs) before and after irradiation for BGW (A and C) and IGW (B and D). Before-after differences are shown in panels E (BGW) and F (IGW). The positions of the EMM peaks defined by Coble (2007) (Table 1) are also displayed. Irradiation lasted for 474 h for BGW and 140 h for IGW.

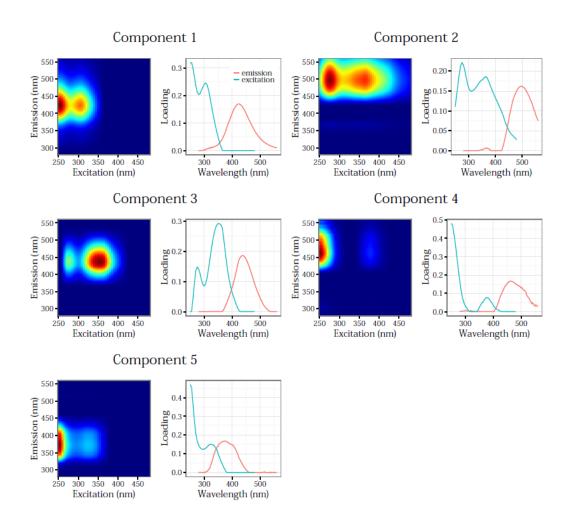


Figure 4. Excitation-emission contours of five components identified by PARAFAC modeling (left panels) and split-half validations of excitation and emission loadings (right panels).

### 3.1.4. Sources of DOM

Taken together the absorption indicators (SUV  $\alpha_{254}$ ,  $S_{\lambda}$ ,  $E_2/E_3$ ), fluorescence indices (FI, BIX, HIX), EEM peaks, PARAFAC components, and MW values, it can be inferred that BGW predominantly contained old DOM of terrestrial origin characterized by large molecular size, high aromaticity, and strong humification, whereas IGW mostly comprised fresh, microbially derived DOM typically having smaller molecular size, lower aromaticity, and weaker humification. A dominant terrestrial origin for DOM<sup>BGW</sup> is consistent with the proposition of a paleosol source noted above, though microbial production occurring along the groundwater flow path could also have contributed to some extent to the DOM pool in the intertidal area (Couturier et al., 2016). For IGW, its DOM signature might also be influenced by the hydrological, physical, and biogeochemical processes taking place in the overlying soil column. Generally, surface plant litter and soil are significant sources of DOM in groundwater (Baker et al., 2000; Shen et al., 2015). Precipitation drives the transport of DOM down to the saturated zone of the soil column. Selective adsorption of high-MW DOM such as lignin-derived phenols onto soil minerals in the unsaturated zone diminishes the aromaticity of DOM reaching the saturated zone (Shen et al., 2015). The chemical composition and optical properties of this diagenetically altered soil DOM tend to resemble those of microbially derived DOM. Aquifers in the Martinique Beach at the time of sampling (June) were recently recharged with snowmelt (Chaillou et al., 2014). Hence, in addition to DOM microbially produced in the aquifer and/or from the overlying soil column, the DOM<sup>IGW</sup> could partly originate from the diagenetically modified DOM from the surface soil.

#### 3.1.5. Potential interference of iron

The presence of significant Fe(III) can interfere with the optical properties of CDOM, since Fe(III) absorbs both UV and visible radiation (Weishaar et al., 2003; Xiao et al., 2013; Poulin et al., 2014). Xiao et al. (2013) reported Fe(III)-specific absorption coefficients ( $a_{\text{Fe}}^*$ ) over a broad wavelength range from 320 nm to 650 nm at 1-nm intervals, while Poulin et al. (2014) published  $a_{\rm Fe}^*$  values at individual wavelengths of 254 nm, 280 nm, 350 nm, and 400 nm. Both groups found little dependence of  $a_{\rm Fe}^*$  on the source and characteristics of DOM with which Fe(III) is associated. The  $a_{\rm Fe}^*$  values obtained by Poulin et al. (2014) agreed well with those by Xiao et al. (2013) at 350 nm and 400 nm but are somewhat higher (~15%) than the values at 254 nm and 280 nm extrapolated from the results of Xiao et al. (2013)(错误!未找到引用源。). Based on the Fe(III) concentrations reported in Section 3.1.1, we estimated the effects of Fe(III) on the absorbance-based indicators in our samples using  $a_{\rm Fe}^*$  values determined by Poulin et al. (2014) (254 nm for SUV  $\alpha_{254}$ ) and by or extrapolated from Xiao et al. (2013) (330 nm for  $a_{\text{CDOM}}(330)$ , 250 nm and 365 nm for E<sub>2</sub>/E<sub>3</sub>, and 275-295 nm for  $S_{275-295}$ ). The results are summarized in Table 4. For the Fe-rich BGW, the correction lowers  $a_{CDOM}(330)$  and  $SUV_{\alpha_{254}}$  by 30.8% and 24.4% and raises  $S_{275-295}$  and  $E_2/E_3$  by 16.1% and 23.0%. For the Fe-poor IGW, the effect is considerably weaker, reducing  $a_{CDOM}(330)$  and SUV  $\alpha_{254}$  by 12.6% and 6.7% and increasing  $S_{275-295}$  and  $E_2/E_3$  by 5.2% and 16.6%. Despite the differing extents of the iron effect, the relative differences in these indicators between the two samples are not affected to a large extent by the Fe-corrections (Table 4). In terms of absolute values, the corrected SUV  $\alpha_{254}^{BGW}$  (4.07 L mg<sup>-1</sup> m<sup>-1</sup>) is still comparable to those of humic substances extracted from highly colored fluvial waters (Weishaar et al., 2003; Poulin et al., 2014). The after-correction  $M_{\rm w}$  estimated from

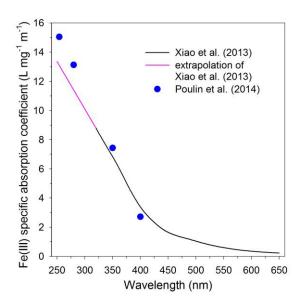


Figure 5. Fe(III) specific absorption coefficient (e-based) published by Xiao et al. (2013) and Poulin et al. (2014). Specific absorption coefficients at <320 nm are linearly extrapolated from the data of Xiao et al. (2013) over 320-350 nm.

the  $M_{\rm w}$ -E<sub>2</sub>/E<sub>3</sub> relationship of Lou and Xie (2006) is 1.53 kDa for BGW and 0.61 kDa for IGW.  $M_{\rm w}^{\rm BGW}$  decreases by almost 50% compared with the before-correction estimate (2.93 kDa) but remains within the high-molecular-weight class. The change in  $M_{\rm w}^{\rm IGW}$  (0.70 $\rightarrow$ 0.61 kDa) is much milder but pushes it towards the lower territory within the low-molecular-weight category.

The interference of Fe with fluorescence depends on various factors, including the ratio of Fe to DOC concentration, CDOM composition, pH, etc., rendering it difficult to quantitatively correct this artifact (Aiken, 2014; Poulin et al., 2014). Given the very low total Fe:DOC ratio in IGW (0.030 mg Fe (mg C)<sup>-1</sup>), the Fe interference in this sample should be minor (Cory, 2005; Poulin et al., 2014). The highly elevated total Fe:DOC ratio in BGW (0.25 mg Fe (mg C)<sup>-1</sup>) made the Fe effect potentially important. However, this effect usually decreases the total fluorescence intensity, increases FI, and preferentially quenches the humic regions of the EEM spectra (i.e. peaks A and C) and the PARAFACderived terrestrial humic components (here C1 and C2) (Aiken, 2014; Poulin et al., 2014). Hence, excluding the Fe effect in BGW would further widen the differences between BGW and IGW in the fluorescence properties determined by this study and their implications for characterizing the source of DOM. To sum up, the potential Fe interferences with absorbance and fluorescence should not change the conclusions reached in section 3.1.4 regarding the sources and characteristics of DOM in BGW and IGW.

Aside from Fe, nitrate and nitrite are also UV-absorbers (Mack and Bolton, 1999) and thus may interfere with CDOM absorbance at elevated concentrations. Yet the aggregate nitrate and nitrite concentration is <100 µmol L<sup>-1</sup> in IGW and <10 µmol L<sup>-1</sup> in

BGW (Couturier et al., 2017), making them unlikely to significantly affect the absorption spectra and the absorbance-based indicators in this study (Weishaar et al., 2003).

**Table 4.**Effects of iron on the optical properties of the groundwater samples.

		асром(330) (m <sup>-1</sup> )			S <sub>275-295</sub> (μm <sup>-1</sup> )			E <sub>2</sub> /E <sub>3</sub>			SUVα <sub>254</sub> (L mg <sup>-1</sup> m <sup>-1</sup> )		
		BGW	IGW	BGW/IGW	BGW	IGW	BGW/IGW	BGW	IGW	BGW/IGW	BGW	IGW	BGW/IGW
Before (BC)	correction	389.2	4.36	89.2	11.8	15.4	0.77	3.94	7.93	0.50	5.39	2.85	1.89
After (AC)	correction	269.5	3.82	70.6	13.7	16.2	0.84	4.85	9.24	0.52	4.07	2.66	1.53
(AC-BC)	/BC*100	-30.8	-12.6		16.1	5.2		23.0	16.6		-24.4	-6.6	

## 3.2. Photobleaching of absorbance

# 3.1.6. Absorption coefficients

Irradiation resulted in loss of  $a_{\rm CDOM}$  throughout the UV and visible regimes for both BGW and IGW, but their spectral patterns of percent loss of  $a_{\rm CDOM}$  differed (Figure 6). While IGW displayed a marked elevation over 280-350 nm, BGW illustrated a much gentler wavelength dependence characterized by a mild convex at halfway exposure and a slightly tilted line after the entire exposure period. At the end of irradiation,  $a_{\rm CDOM}^{\rm BGW}$  lost 97.3% and 99.4% in the UV (250-400 nm) and visible (400-550 nm) ranges, respectively. As the contributions of Fe(III) to  $a_{\rm CDOM}^{\rm BGW}$  are estimated to be 29.6% and 40.9% in corresponding wavelength bands, either most Fe(III) was photochemically reduced to non-absorbing Fe(II) (Blough and Zepp, 1995; Doane and Horwáth, 2010) or the contributions of Fe(III) to absorption are overestimated. The percent losses of  $a_{\rm g}^{\rm IGW}$  were considerably smaller, being 64.7% in the UV and 50.7% in the visible. The lower losses in IGW were partly due to its shorter exposure duration (see Chapter 2.3). The very low initial CDOM absorption in IGW (Table 3) made it difficult to precisely determine  $a_{\rm CDOM}$  after longer exposures with the conventional spectrophotometer used in this study.

Absorption coefficients at single UV wavelengths, such as 325 nm (Nelson et al., 2013), 330 nm (Osburn et al., 2009a), 350 nm (Miller and Zepp, 1995), and 375 nm (Blough and Del Vecchio, 2002), have been frequently used as a measure of CDOM content and photobleaching. As photobleaching alters the spectral slope of CDOM absorption spectra (e.g. Helms et al., 2008, 2013; Sharpless and Blough, 2014), single wavelengths may not be ideal for quantifying photobleaching. Alternatively, some studies reported photobleaching at several wavelengths (e.g. Osburn et al., 2009b) or using

absorption coefficients integrated or averaged over broad wavelength ranges (Vähätalo and Wetzel, 2004; Osburn et al., 2009b; Porcal et al., 2014). In this study we use the average remaining fraction of  $a_{\rm CDOM}$  over 250-400 nm, ARF<sub>250-400</sub>, or its corresponding fractional loss AFL<sub>250-400</sub>, to quantify photobleaching of absorbance. ARF<sub>250-400</sub> and AFL<sub>250-400</sub> are defined as follows:

$$ARF_{250-400} = 151^{-1} \sum_{\lambda=250}^{400} \left( a_{CDOM,t}(\lambda) \left( a_{CDOM,0}(\lambda) \right)^{-1} \right)$$
 (4)

$$AFL_{250-400} = 1 - ARF_{250-400}$$
 (5)

where  $a_{\text{CDOM,t}}$  stands for  $a_{\text{CDOM}}$  at irradiation time t and  $a_{\text{CDOM,0}}$  for  $a_{\text{CDOM}}$  at time zero (dark controls showed negligible changes in  $a_{\text{CDOM}}$ ). The 250-400 nm interval was chosen because  $a_{\text{CDOM}}$  in this band could be more accurately determined than in the visible range and because the percent loss of  $a_{\text{CDOM}}$  for IGW in the UV was higher than that in the visible (Figure 6). Note that results and conclusions based on ARF<sub>250-400</sub> (and AFL<sub>250-400</sub>) should be similar to those based on  $a_{\text{CDOM}}$  at individual wavelengths or averaged over broad wavelength bands, since ARF<sub>250-400</sub> is well correlated with the latter parameters (Figure 7 and Figure 8).

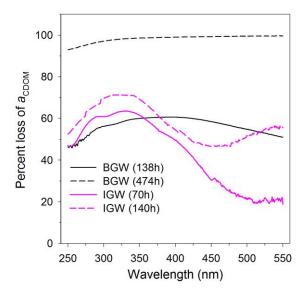


Figure 6. Spectral distribution of percent loss of CDOM absorption coefficient at selected irradiation times.

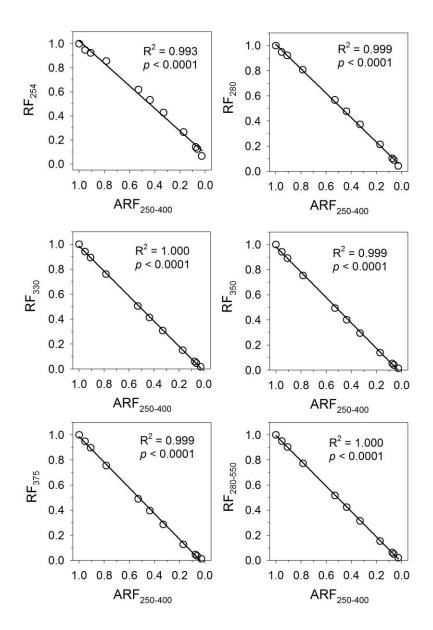


Figure 7. Linear correlation between the remaining fraction of CDOM absorption coefficient at selected individual wavelengths (RF $_{\lambda}$ ) and the average remaining fraction of CDOM absorption coefficient over 250–400 nm (ARF $_{250\text{-}400}$ ) for beach groundwater (BGW). See Eq. 4 in the main text for the definition of ARF $_{250\text{-}400}$ . Y-axis RF $_{280\text{-}550}$  of the right bottom panel represents the remaining fraction of the mean CDOM absorption coefficient over 280–550 nm.

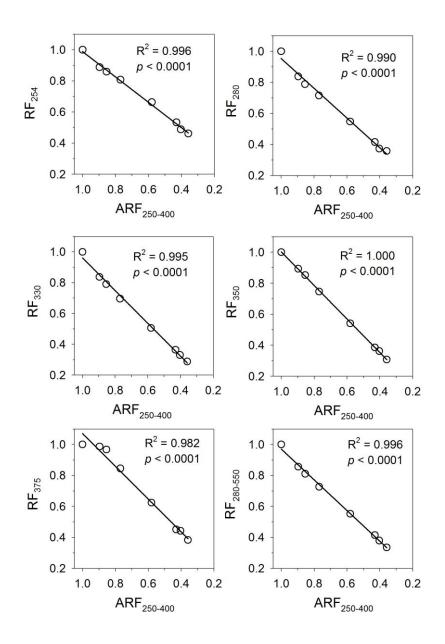


Figure 8. Same as in Figure 7 except for inland groundwater (IGW)

ARF<sub>250-400</sub><sup>BGW</sup> decreased with irradiation time linearly during the first 210 h and exponentially after that (Figure 10). The extremely high absorption of BGW caused strong self-shading (Figure 9) which could have led to the zero-order kinetics observed during the first half of the irradiation (Miller, 1998). ARF<sub>250-400</sub><sup>IGW</sup> declined exponentially throughout the exposure period but the rate constant almost halved during the last 70 h (Figure 5). The IGW data can also be fitted to two different linear relationships with slightly weaker statistic performance compared with the exponential fit (Table 2). The biphasic kinetic behavior of photobleaching seen in the present study has also been observed previously and attributed to the presence of two CDOM components with differing bleaching turnover times (Del Vecchio and Blough, 2002; Grebel et al., 2009). The photobleaching of CDOM<sup>IGW</sup> was apparently faster relative to that of CDOM<sup>BGW</sup> (Figure 5), which could be due partly to the much stronger self-shading in BGW (Figure 9).

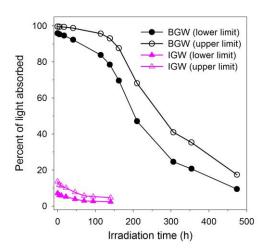


Figure 9. Percent of light absorbed by CDOM averaged over the wavelength range from 280 nm to 400 nm for time-course irradiations described in section 2.3 of the main text. The percent of light absorbed (PLA) is calculated as:

$$PLA(\lambda) = a_{CDOM}(\lambda)(a_{tot}(\lambda))^{-1}(1 - \exp(-a_{tot}(\lambda)l))$$

where l (m) denotes the pathlength of the irradiation tube,  $a_{\text{CDOM}}(\lambda)$  (m<sup>-1</sup>) the absorption coefficient of CDOM, and  $a_{tot}(\lambda)$  (m<sup>-1</sup>) the sum of  $a_{CDOM}(\lambda)$  and the absorption coefficient of pure water at taken from Pope and Fry (1997) and Buiteveld et al. (1994). l is estimated using the equation of Vähätalo and Zepp (2005):

$$l = z(\mu(\lambda))^{-1}$$

where  $z = 0.5\pi r$  with r being the inner radius of the irradiation tube (0.011 m) (Vähätalo and Zepp, 2005) and  $\mu(\lambda)$  denotes the mean cosine of photons (dimensionless). z is estimated to be 0.0173 m for IGW (tube fully filled with sample water) and 0.0121 m for BGW (tube partially filled with sample water, see section 2.3 in the main text). The lower and upper limits of  $\mu(\lambda)$  are 0.5 and 1 (Vähätalo and Zepp, 2005), giving the upper and lower limits of l and PLA, respectively. The lower and upper limits of l are 0.0121 m and 0.0242 m for BGW and 0.0173 m and 0.0346 m for IGW.

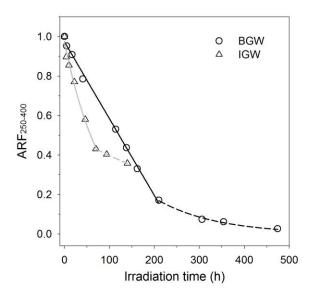


Figure 10. Average remaining fraction of CDOM absorption coefficient from 250 nm to 400 nm (ARF $_{250\text{--}400}$ ) vs. irradiation time. Lines represent best fits of the data. See Appendix for fitted equations and statistics.

### 3.2.3 Absorption-based indicators

Because of different extents of self-shading experienced by BGW and IGW, we chose plots of property vs. ARF<sub>250-400</sub> (except NH<sub>4</sub><sup>+</sup>, see section 3.5) instead of exposure time or cumulative irradiance, to discuss the time-course data and compare the photochemical behavior of the two samples. Besides, in light of the much larger difference in E<sub>2</sub>/E<sub>3</sub> than in  $S_{275-295}$  between the two samples prior to irradiation (Table 3), we opted for E<sub>2</sub>/E<sub>3</sub> to track the change of the spectral shape of  $a_{CDOM}$ , though results from the two parameters are similar (data not shown).

 $E_2/E_3$  increased while SUV $\alpha_{254}$ ,  $M_W$ , and  $\rho$  decreased with photobleaching (Figure 11), in line with previous studies (Lou and Xie, 2006; Brooks et al., 2007). Owing partly to the more extensive bleaching of BGW, the overall changes of these parameters were far larger in BGW than in IGW, being 442% vs. 37.5% for E<sub>2</sub>/E<sub>3</sub>, 87.7% vs. 45.0% for  $SUV_{\alpha_{254}}$ , 38.2% vs. 4.5% for  $M_W$ , and 27.0% vs. 2.7% for  $\rho$ . The changes of  $E_2/E_3$  and  $SUV_{\alpha_{254}}$  for the two samples were, however, similar at comparable fractional absorption loss values (Figure 6A,B), suggesting that aromatic moieties were the dominant bleachable chromophores in both BGW and IGW. Unlike  $E_2/E_3$  and  $SUV\alpha_{254}$ , the fractional losses of  $M_W$  and  $\rho$ , following linear patterns, underwent six times faster in BGW than in IGW on a relative absorption loss basis (Figure 6C). Hence, a much larger proportion of chromophores on CDOMBGW than on CDOMIGW was photochemically converted into less or non-absorbing lower-MW compounds (e.g., low-MW organic acids and inorganic carbon). This led to a substantial reduction of both the average MW and heterogeneity (i.e. smaller  $\rho$ ) of CDOM<sup>BGW</sup>, while the changes of these two quantities were marginal for CDOM<sup>IGW</sup>.

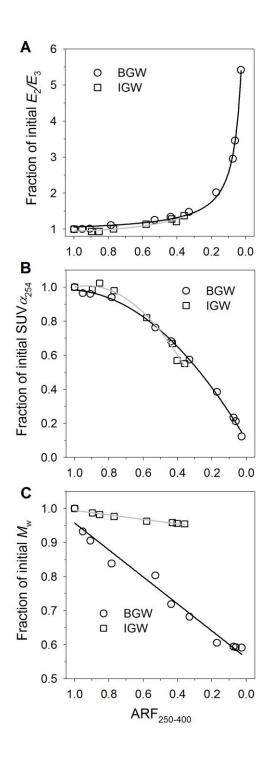


Figure 11. Fractions of initial absorption ratio  $E_2/E_3$  (A), specific UV absorption coefficient at 254 nm (B), and weight-averaged molecular weight (C)  $\nu s$ . average remaining fraction of CDOM absorption coefficient from 250 nm to 400 nm (ARF<sub>250-400</sub>). Black lines are best fits to BGW data and grey lines to IGW data. See Appendix for fitted equations and statistics.

## 3.2.3 AQYs of absorbance photobleaching

Fitted parameters for modeling the AQY spectra of absorbance photobleaching are shown in Table 5. The AQY spectra of BGW and IGW differed markedly in magnitude and shape, with  $\Phi_{\text{CDOM}}(\lambda)^{\text{BGW}}$  superior over  $\Phi_{\text{CDOM}}(\lambda)^{\text{IGW}}$  at  $\lambda < 462$  nm but inferior to the latter at  $\lambda > 462$  nm (Figure 12A). The spectrum of IGW was much flatter than that of BGW in the UV-A and visible regimes (Figure 12A).  $\Phi(330)_{\text{CDOM}}$  and  $\overline{\Phi}_{\text{CDOM}}$  for BGW were nearly 29 and eight times larger, respectively, than those of IGW (Table 5). The action spectra of the two samples both reached maxima at ~330 nm but IGW exhibited a far stronger tailing over the longer wavelengths (Figure 12B), consistent with its flatter AQY spectrum. The visible (400-600 nm) contributed 63% of the total area under the action spectrum for IGW but only 8% for BGW, demonstrating that photobleaching of IGW was dominated by the visible, whereas that of BGW by the UV.

The AQY and action spectra for IGW resemble those for the Mackenzie River water (Osburn et al. 2009) and  $\overline{\Phi}_{CDOM}$  for IGW is close to that for the Mackenzie River water  $(1.12 \times 10^{-2} \, vs. \, 1.36 \times 10^{-2})$  as well (Figure 12), despite the use of different models for describing the shape of the AQY spectrum (i.e. quasi-exponential in this study vs. simple exponential by Osburn et al. (2009)). Interestingly, SUV $\alpha_{254}$  for IGW (2.85 L<sup>-1</sup> mg<sup>-1</sup> m<sup>-1</sup>) is also comparable to that for the Mackenzie River water (2.97 L<sup>-1</sup> mg<sup>-1</sup> m<sup>-1</sup>) which was estimated from  $a_{CDOM}(330)$  (10.22 m<sup>-1</sup>), S (18.62  $\mu$ m<sup>-1</sup>), and DOC (512  $\mu$ mol L<sup>-1</sup>) reported by Osburn et al. (2009). Likewise, the higher  $\overline{\Phi}_{CDOM}$  and SUV $\alpha_{254}$  conforms to previous observations showing a high susceptibility of phenols and other aromatic

moieties to photooxidation and similar loss rates of the aromatic components and absorbance (Sharpless and Blough, 2016 and references therein). It also complies with the positive relationships between AQY of carbon monoxide photoproduction and SUV $\alpha_{254}$  determined for CDOM in natural waters (Zhang et al., 2006; Stubbins et al., 2008) and for model aromatics (Stubbins et al., 2008). Apart from the rich aromaticity, the extraordinarily elevated Fe concentration in BGW might also have contributed to its higher efficiency of photobleaching, since Fe-mediated photo-Fenton reactions, which produce the highly oxidizing hydroxyl radical (Zepp et al., 1992), accelerate CDOM photooxidation and photobleaching (Gao and Zepp, 1998; White et al., 2003; Molot et al., 2005).

**Table 5**. Fitted parameters for Eq. 2 in the text, the apparent quantum yield at 330 nm ( $\Phi(330)$ ), and the solar spectrum-weighted mean apparent quantum yield ( $\overline{\Phi}$ ) (Eq. 4 in the text) for photobleaching (m<sup>-1</sup> CDOM (mol photons)<sup>-1</sup>) and photoammonification (mol NH<sub>4</sub><sup>+</sup> (mol photons)<sup>-1</sup>). nRMSE and R<sup>2</sup> are normalized root-mean-square error and coefficient of determination, respectively, for the regression of modeled against measured rates.

	Sample	$m_1$	$m_2$	$m_3$	Ф(330)	$ar{\Phi}$	nRMSE a	$\mathbb{R}^2$
							(%)	
Photobleaching	BGW	$1.56 \times 10^{-8}$	6154.8	1.39	1.81	$9.33 \times 10^{-2}$	8.2	0.988
	<b>IGW</b>	$2.53 \times 10^{-3}$	280.1	-243.0	0.063	$1.12 \times 10^{-2}$	6.2	0.978
Photoammonification	BGW	$5.63 \times 10^{-10}$	2031.6	-165.5	$1.30 \times 10^{-4}$	$6.97 \times 10^{-6}$	5.4	0.997
	<b>IGW</b>	$2.67 \times 10^{-7}$	882.6	-197.9	$2.13 \times 10^{-4}$	$1.69 \times 10^{-5}$	6.7	0.993

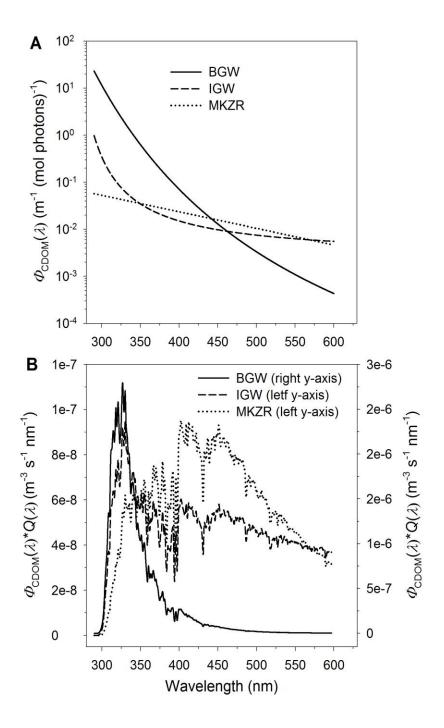


Figure 12. Apparent quantum yield (AQY) spectra of CDOM absorbance photobleaching (A) and their corresponding action spectra (B). The AQY data for the Mackenzie River (MKZR) refers to station R2 (salinity 0.1) in the study of Osburn et al. (2009).

# 3.3. Photobleaching of fluorescence

## 3.1.7. Fluorescence indices

For IGW, FI and HIX decreased exponentially and BIX increased linearly with photobleaching by 12.5%, 80.4%, and 30.8%, respectively, over the entire exposure period (Figure 13). BGW exhibited an overall decrease in FI (1.2%) and HIX (75.6%) and increase in BIX (30.1%) but the trends were less regular (Figure 13). For instance, the upsurge of FI<sup>BGW</sup> (3.7%) over the final 1.3% absorption loss deviated from the earlier linear declining tendency (Figure 13A); BIX<sup>BGW</sup> slightly fluctuated nearby the initial value before rapidly rising at ARF<sub>250-400</sub> < 0.17 (Figure 13B); the general exponential decay of HIX<sup>BGW</sup> was superimposed with a plateau over ARF<sub>250-400</sub> from 0.44 to 0.17. The causes underlying these irregularities are unclear. Notably, the fractional changes of all three indices *vs.* AFL<sub>250-400</sub> were usually much larger in IGW than in BGW, indicating that the indices for IGW were more sensitive to photobleaching than those for BGW.

The reliance of the fluorescence indices on irradiation history seen in the present study corroborates earlier studies showing light-induced decreases in FI of riverine and wetland DOM (Brooks et al., 2006) and decreases in HIX and increases in BIX of DOM isolated from deep-ocean water (Helms et al., 2013). Hence, use of these indices for DOM source identification (FIX and BIX) and age assessment (HIX) can be imprecise or even misleading, if the targeted samples have been subjected to extensive light exposure. Under this circumstance, a larger number of compositionally based optical properties, alongside more sophisticated statistic tools, are needed to infer the source material of DOM (Hansen et al., 2016).

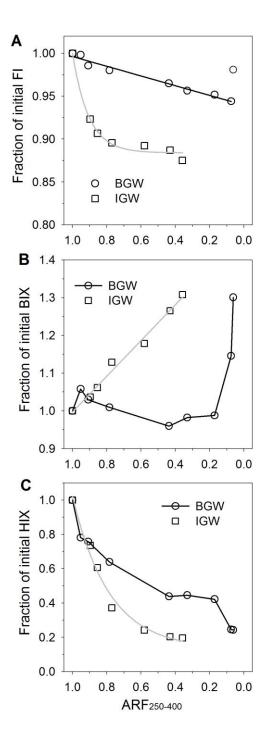


Figure 13. Fractions of initial FIX (A), BIX (B), and HIX (C) vs. average remaining fraction of CDOM absorption coefficient from 250 nm to 400 nm. Black line in panel A is best fit to BGW data (excluding last point) and black lines in panels B and C are data point connecting lines for BGW. Grey lines in panel A-C are best fits to IGW data. See Appendix for fitted equations and statistics.

## 3.1.8. EMM peaks and PARAFAC components

After full-length irradiations, the intensities of peak A dropped by 79% for BGW and 72% for IGW, and peak C by 73% and 68%, respectively. The large decreases in the intensities of these peaks demonstrate a rapid photodegradation of humic-like material, which is consistent with the results of earlier studies (Helms et al., 2013; Timko et al., 2015) and with the photochemical losses of SUV $\alpha_{254}$  and MW (section 3.2). The ratio of peak C to A, an indication of the amount of humic acid-like vs. fulvic acid-like fluorophores (Chen et al., 2003; Hansen et al., 2016), increased from 0.79 to 1.01 for BGW and from 0.66 to 0.78 for IGW, suggesting a preferential photodegradation of fulvic acids like FDOM. Notably, the originally vague fluorescence signals adjoining the lower edge of peak  $A^{IGW}$  (Figure 3B) became a distinct peak centered at  $\lambda_{ex}/\lambda_{em}$  255 nm/356 nm (peak X herein) after irradiation (Figure 3D). The appearance of peak X was due more likely to the decay of peak A than to an increase of its intensity or the formation of a new peak, since the fluorescence signal at  $\lambda_{ex}/\lambda_{em}$  255 nm/356 nm diminished during the first 22 h irradiation  $(0.240 \rightarrow 0.209 \text{ R.U.})$  and thereafter remained rather constant  $(0.207 \pm 0.0025 \text{ R.U.})$ . The initial decline of this fluorescence signal apparently arose from the dwindling of peak A's periphery which likely masked peak X at the beginning (Figure 3B). Peak X falls into the broad protein-like category defined by Chen et al. (2003), suggesting that this proteinaceous material is photochemically recalcitrant. Previous studies also reported photochemical resistance of protein-like fluorescence peaks (Cory et al., 2007; Clark et al., 2008). In terms of the EEM location, peak X is blue-shifted by 20 nm in excitation and red-shifted by 16 nm in emission relative to peak

T (protein- and tryptophan-like) defined by Coble (2007) (Table 1 and Figure 3B). The blue shift in excitation implies that the protein-like material signified by peak X may possess a higher aromaticity than the protein- or tryptophan-like material denoted by peak T.

Figure 14 shows  $F_{max}$  of PARAFAC components as a function of ARF<sub>250-400</sub>. For BGW, C1 decayed gently up to ARF<sub>250-400</sub> ~0.2 and then abruptly with further bleaching (Figure 14A). The overall decay trends of C2, C3, and C4 encountered minor to moderate rebounds during the late stage of photobleaching (Figure 14B-D); the causes underlying these bulges are unclear. C5 remained undetectable up to ARF<sub>250-400</sub> 0.44 when it started rising rapidly (Figure 14E), contrasting with the behavior of C1-C4. C1 to C4 declined by 62%, 89%, 87%, and 64%, respectively, at the end of irradiation, thereby enhancing the relative abundances of C1 and C4 but diminishing those of C2 and C3 (Figure 15). For IGW, all components decreased with photobleaching, exponentially for C1 and linearly for C2-C5. C1-C5 dropped by 100%, 80%, 84%, 73%, and 38%, respectively, after the entire exposure period, which greatly increased the relative abundance of C4 among the humic components (C1-C4) (Figure 15). The fractional loss of C1 in IGW underwent much faster than that in BGW (Figure 14A), while the differences in the loss of C2-C4 between the two samples were considerably smaller (Figure 14B-D). After full-length irradiations, BGW was further enriched with C1 relative to IGW due to the complete disappearance of C1 in IGW (Figure 15).

The loss of the PARAFAC humic components (C1-C4) is consistent with the reduced intensities of the humic-like EEM peaks A and C. The behavior of C1-C4 further reveals that 1) both the terrestrial (C1-C2) and the autochthonous (C3-C4) humic-like

components are photochemically reactive; 2) the photolabilities of C2-C4 are comparable on both intra- and inter-sample bases; 3) the photolability of C1 in IGW is much higher than that of C1 in BGW (and also higher than any other components in both samples), pointing to C1 being the dominant UV absorber in IGW, the presence of different subclasses of C1 in BGW and IGW and/or different matrix effects on photodegradation of C1. The loss of the protein-like C5 in IGW is inconsistent with the constancy of the intensity of its protein-like EEM peak X, suggesting that C5 and peak X may represent different classes of proteinaceous compounds that vary in photolability. Similarly, Bittar et al. (2015) suspected that peak T may comprise multiple protein constituents that differ in composition and hence photolability. The formation of C5 in BGW implies the photochemical synthesis of protein-like or related nitrogen-containing structures. Cory et al. (2015) also observed that solar irradiation increased the fluorescence intensity of protein-like FDOM in stream water. Nitrogen-containing organic compounds or moieties (e.g. amine, amide, amino acids) can be produced from photochemical transformation of CDOM (Bushaw et al., 1996; Tarr et al., 2001) or photochemical incorporation of inorganic nitrogen (nitrate, nitrite, and ammonium) into DOM (Kieber et al., 1997; Thorn and Cox, 2012).

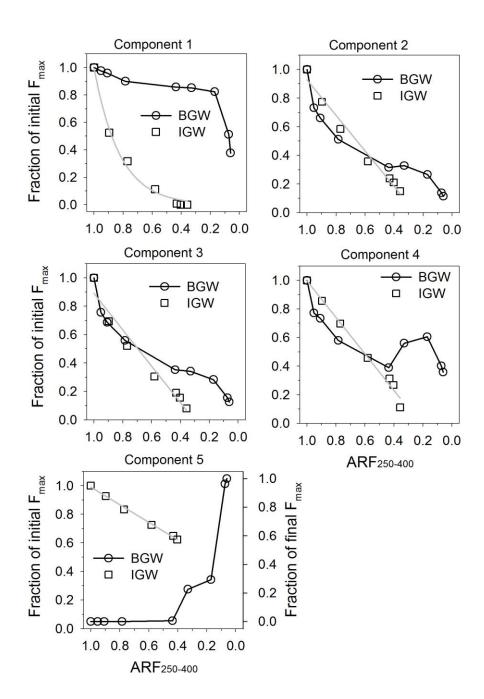


Figure 14. Fractions of initial or final maximum fluorescence intensities of five PARAFAC-modeled components vs. average remaining fraction of CDOM absorption coefficient from 250 nm to 400 nm. In panel E, the right y-axis is for BGW and the left for IGW. The final absolute  $F_{max}$  value for BGW is 1.72 R.U. Black lines are data point connecting lines for BGW. Grey lines are best fits to IGW data. See Appendix for fitted equations and statistics.

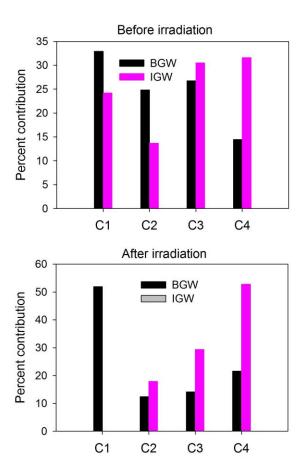


Figure 15. Percent contribution of each individual component among C1-C4 to the summation of C1-C4 before and after full-length irradiations (i.e. 474 h for BGW and 140 h for IGW). The percent contribution of component i is defined as  $F_{\max(i)}/\sum_{i=1}^4 F_{\max(i)} \times 100$ , where  $F_{\max(i)}$  is the maximum fluorescence intensity of component i.

## 3.4. DOC photomineralization

Throughout the exposure periods, the fractional losses of DOC were consistently lower than those of absorption coefficient for both BGW and IGW (Figure 16A). The ratio of DOC loss to AFL<sub>250-400</sub> ( $R_{\text{DOC/CDOM}}$ ) remained nearly constant, 0.33 for BGW and 0.19 for IGW over much of the photobleaching regimes (AFL<sub>250-400</sub>  $^{\text{BGW}} \leq 0.83$  and AFL<sub>250-400</sub>  $^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{BGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{BGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{BGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{BGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{BGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).  $R_{\text{DOC/CDOM}} ^{\text{IGW}} \leq 0.57$ ) (Figure 16B) but increased with further bleaching (Figure 16A).

AQYs of photochemical DOC losses ( $\Phi_{DOC}$ ) for BGW could not be derived from the time-series irradiation, because the irradiation tubes containing BGW were partially filled with air (section 2.3), which made it difficult to quantify the photons absorbed in these tubes. On the other hand, as DOC was not measured due to technical constraints during the multispectral irradiations using cutoff filters,  $\Phi_{DOC}$  could not be directly obtained from these experiments, either. However, given the similarity in the spectral compositions of the radiation under the WG280 cutoff filter and the full-spectrum radiation used in the time-series exposures (Figure 17), the DOC losses under WG280 can be estimated by the corresponding AFL<sub>250-400</sub> multiplied by  $R_{DOC/CDOM}$  obtained from the time-series irradiations. Note that the AFL<sub>250-400</sub> under WG280 (BGW: 0.063; IGW:

0.10) were within the AFL<sub>250-400</sub> ranges showing linear relationships between the fractional DOC loss and AFL<sub>250-400</sub> (Figure 16B). The number of photons absorbed by CDOM was calculated using Eq. 1. and integrated from 290 nm to 450 nm. The choice of this wavelength range for integration is based on two factors: 1) photomineralization is driven predominantly by the UV radiation (Powers and Miller, 2015) and 2) the light source used in the present study exhibited multiple irradiance spikes between 450 nm and 500 nm (Figure 17) which would underestimate  $\Phi_{DOC}^{BGW}$  relative to  $\Phi_{DOC}^{IGW}$ , since the absorption spectrum of BGW was flatter and far more elevated across the visible regime than that of IGW (Figure 2). Dividing the moles of DOC loss by the moles of photons absorbed gives a broadband  $\Phi_{DOC}$  over the 280-450 nm wavelength range. The  $\Phi_{DOC}$  (mol DOC (mol photons)<sup>-1</sup>) thus obtained was  $1.42 \times 10^{-3}$  for BGW and  $6.98 \times 10^{-4}$  for IGW, the former being a factor of two of the latter.

The higher photomineralization efficiency of BGW could be linked to its stronger aromaticity, as inferred from the positive correlation between DIC photoproduction AQY and SUV $\alpha_{254}$  in freshwater lakes (Koehler et al., 2016). However, the twofold difference in  $\Phi_{DOC}$  between BGW and IGW is much smaller than their eightfold disparity in  $\overline{\Phi}_{CDOM}$  (section 3.2.3), suggesting that aromaticity is less important in controlling photomineralization than photobleaching. Photodecarboxylation has been identified as a major mechanism for DOM photomineralization (Ward and Cory, 2016). Interestingly, preferential photodecarboxylation occurs with the highly oxidized, weaker UV-absorbing aliphatic carboxylic acids rather than the less oxidized and strong UV-absorbing aromatic acids (Ward and Cory, 2016). Our result is thus consistent with the finding of Ward and Cory (2016). Similar to photobleaching, the very rich Fe content in BGW would also

enhance its photomineralization efficiency through the reaction of DOM with hydroxyl radicals generated from the photo-Fenton reactions (Gao and Zepp, 1998; Goldstone et al., 2002), even though this effect could be lessened due to the relatively weak acidity (pH: 6.29, Table 3) of the sample (Gu et al., 2017).

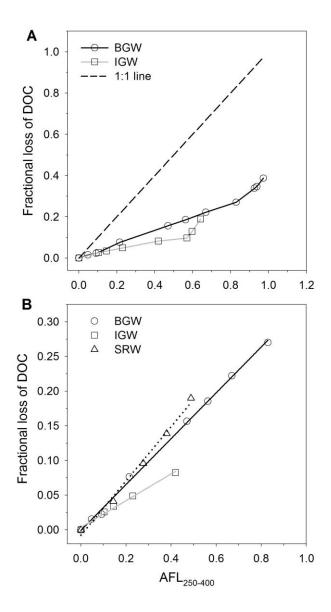


Figure 16. Fractional loss of DOC *vs.* average fractional loss of CDOM absorption coefficient from 250 nm to 400 nm for all data points (A) and excluding the last three data points (B). Lines in panel B are best fits to data (see Appendix for fitted equations and statistics). For comparison, the Saguenay River water (SRW) data is included (Zhang and Xie, 2015).

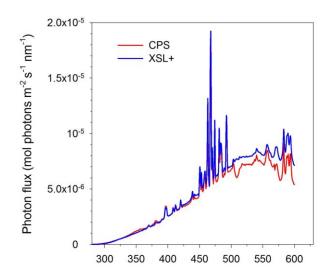


Figure 17. Photon flux spectra of the SUNTES CPS (under cutoff filter WG280) and XLS+ (without cutoff filters) solar simulators. The CPS was used for AQY experiments and the XLS+ for time-course irradiations described in Chapter 2.

#### 3.4.1. Photoammonification

Results from the time-series irradiations are shown with plots of NH<sub>4</sub> concentration vs. absolute loss of absorption coefficient averaged over 250-400 nm ( $\Delta a_{\rm CDOM}(250-400)$ ) (Figure 18). NH<sub>4</sub><sup>+</sup> concentration increased linearly with photobleaching at  $\Delta a_{\text{CDOM}}(250 400)^{\text{BGW}} < 87.56 \text{ m}^{-1} \text{ (R}^2 = 0.985) \text{ and } \Delta a_{\text{CDOM}} (250-400)^{\text{IGW}} < 1.73 \text{ m}^{-1} \text{ (R}^2 = 0.975) \text{ and}$ leveled off with further bleaching. Over the linear ranges, the amount of NH<sub>4</sub> produced per unit of  $\Delta a_{\rm CDOM}(250\text{-}400)$  in IGW (1.12 µmol L<sup>-1</sup> m) was 80 times higher than in BGW (0.014 µmol L<sup>-1</sup> m), demonstrating that DOM<sup>IGW</sup> was highly enriched relative to DOM<sup>BGW</sup>  $NH_4^+$ precursors with reactive and/or sensitizers mediating photoammonification, if secondary photoprocesses were involved. An enrichment of NH<sub>4</sub> precursors on DOM<sup>IGW</sup> is in accordance with the detection of the protein-like EEM peak X and PARAFAC C5 in IGW and their absence in BGW (section 3.1.3). The decreases in the NH<sub>4</sub> production rate suggest a depletion of the precursors and/or the photosensitizers. Alternatively, it could arise from an increased incorporation of NH<sub>4</sub><sup>+</sup> into DOM at higher NH<sub>4</sub> concentrations (Thorn and Mikita, 1992; Kieber et al., 1997; Gardner et al., 1998; Koopmans and Bronk, 2002; Kitidis et al., 2008; Thorn and Cox, 2012). Declining photoammonification rates with increasing photobleaching have also been reported for humic extract solutions and coastal and oligotrophic waters (Bushaw et al., 1996; Vähätalo and Zepp, 2005; (Kitidis et al. 2006).

Fitted parameters for modeling the AQY spectra of photoammonification are shown in Table 5. The two samples displayed similar  $\Phi_{NH_4^+}$  values in the UVB regime but  $\Phi_{NH_4^+}$  ig became increasingly larger than  $\Phi_{NH_4^+}$  with increasing wavelength over the

UVA and visible bands (Figure 19A). The  $\Phi_{NH_4^+}$  values obtained by the present study are comparable to those reported for a low-salinity Baltic Sea water sample (Stedmon et al., 2007) but substantially higher than those determined with the Mackenzie River water (Xie et al., 2012) (Figure 19A). In terms of  $\Phi_{NH_4^+}(330)$  and  $\overline{\Phi}_{NH_4^+}$ , IGW exceeded BGW by 63.8% and 140%, respectively (Table 5), starkly contrasting with the much lower  $\overline{\Phi}_{CDOM}^{IGW}$  vs.  $\overline{\Phi}_{CDOM}^{BGW}$ .  $\overline{\Phi}_{NH_4^+}$  for IGW is two orders of magnitude higher than that for the Mackenzie River water (1.4 × 10<sup>-7</sup>) and 54% higher than that for the coastal Baltic Sea sample (1.1 × 10<sup>-5</sup>), revealing large variability of the photoammonification efficiency in natural waters. UV (mainly UV-A) dominated the action spectra of both IGW and BGW while the contribution of the visible band was considerably higher in IGW (32%) than in BGW (6%) (Figure 19B) due to the AQY spectrum of IGW being flatter at longer wavelengths (Figure 19A). The visible contribution in IGW is close to that in the Baltic Sea sample (25%) but significantly lower than that for the Mackenzie River water (51%).

The higher  $\Phi_{\mathrm{NH}_{4}^{+}}^{\mathrm{IGW}}$  over  $\Phi_{\mathrm{NH}_{4}^{+}}^{\mathrm{BGW}}$  indicates that neither aromaticity nor Fe concentration is the dominant determinant on the photoammonification efficiency. Xie et al. (2012) observed a seaward increase in  $\Phi_{\mathrm{NH}_{4}^{+}}$  from the inshore Mackenzie River estuary towards the offshore Canada Basin and Arnos et al. (2012) reported increasing  $\Phi_{\mathrm{NH}_{4}^{+}}$  with increasing salinity in the Baltic Sea. As low-salinity waters contain more aromatic carbon-rich DOM and often more dissolved Fe as well, the results from these two studies are consistent with those from the present study. Xie et al. (2012) further revealed that  $\Phi_{\mathrm{NH}_{4}^{+}}$  co-varies with the ratio of dissolved organic nitrogen (DON) to DOC (the data of Arnos et al. (2012) manifested the same trend, though these authors did not

explicitly discuss this issue), implying that the concentration of NH<sub>4</sub><sup>+</sup> precursors relative to that of DOC is an important control on  $\Phi_{NH_4^+}$ . Since DOM<sup>IGW</sup> was likely enriched with NH<sub>4</sub><sup>+</sup> precursors compared with DOM<sup>BGW</sup> (see discussion above in this section), the higher  $\Phi_{NH_4^+}$  IGW is in keeping with the earlier finding.

#### 3.4.2. Photochemical turnover of groundwater DOM in the receiving water body

To a first approximation, the photochemical turnover time of a constituent of groundwater DOM ( $\tau_{DOM}$ , d) in the receiving water body can be estimated as:

$$\tau_{\rm DOM} = [{\rm DOM}] N^{-1} \left( \Phi_{\rm DOM} \int\limits_{290}^{600} a_{\rm CDOM}(\lambda) (a_{TOT}(\lambda))^{-1} Z^{-1} N^{-1} Q_{d0-}(\lambda) d\lambda \right)^{-1}$$

$$= Z[DOM] \Phi_{DOM}^{-1} \left( \int_{290}^{600} a_{CDOM}(\lambda) (a_{TOT}(\lambda))^{-1} Q_{d0-}(\lambda) d\lambda \right)^{-1}$$
 (6)

where  $\tau_{\text{DOM}}$  stands for the photochemical turnover time of CDOM absorbance ( $\tau_{\text{CDOM}}$ ), DOC ( $\tau_{\text{DOC}}$ ) or DON ( $\tau_{\text{DON}}$ ); [DOM] for the concentration of CDOM (represented by  $a_{\text{CDOM}}(330)$ ), DOC (mol m<sup>-3</sup>) or DON (mol m<sup>-3</sup>) in groundwater;  $\Phi_{\text{DOM}}$  for  $\overline{\Phi}_{\text{CDOM}}$ , the broadband  $\Phi_{\text{DOC}}$  (section 3.4), or  $\overline{\Phi}_{\text{NH}_4^+}$ ; N for the dilution factor due to mixing with the receiving water; Z for the surface mixed layer or photoactive layer depth (m), whichever is deeper;  $a_{\text{TOT}}(\lambda)$  for the total absorption coefficient (m<sup>-1</sup>) in the receiving water, i.e. the sum of the absorption coefficients of CDOM, suspended particles, and water;  $Q_{\text{d0-}}(\lambda)$  for the downwelling solar photon flux just below the water surface (mol photons m<sup>-2</sup> d<sup>-1</sup> nm<sup>-1</sup>). Z,  $a_{\text{TOT}}(\lambda)$ , and  $Q_{\text{d0-}}(\lambda)$  are common terms for both BGW and IGW.

The ratio of  $a_{\rm CDOM}(330)^{\rm BGW}/a_{\rm CDOM}(330)^{\rm IGW}$  is ~89,  $\bar{\Phi}_{\rm CDOM}^{\rm BGW}/\bar{\Phi}_{\rm CDOM}^{\rm IGW}$ ~8, and  $a_{\rm CDOM}(\lambda)^{\rm BGW}/a_{\rm CDOM}(\lambda)^{\rm IGW}$  ranges from 69 at 290 nm to 330 at 600 nm and averages 173. Eq. 6 thus indicates that  $\tau_{CDOM}^{BGW}$  is at least six times shorter than  $\tau_{CDOM}^{IGW}$  (69\*8/89 = 6.2). The ratio of DOC<sup>BGW</sup>/DOC<sup>IGW</sup> is 31 and  $\Phi_{DOC}^{BGW}/\Phi_{DOC}^{IGW} \sim 2$ , thereby leading to  $\tau_{\rm DOC}^{\rm BGW}$  being at least four times shorter than  $\tau_{\rm DOC}^{\rm IGW}$  (69\*2/31 = 4.4). Estimating the ratio of  $\tau_{DON}^{BGW}/\tau_{DON}^{IGW}$  poses potentially higher uncertainties, since the exact ratio of  $DON^{BGW}/DON^{IGW}$  is unknown. However, an upper limit of  $DON^{BGW}/DON^{IGW}$  can be derived by equating it to DOCBGW/DOCIGW (i.e. 31), considering that DOMBGW is depleted with protein-like components relative to DOM<sup>IGW</sup> (section 3.1.3). Assuming NH<sub>4</sub> to be the dominant inorganic nitrogen photoproduct from DON (Mopper et al., 2015) and using the known values of  $\bar{\Phi}_{\mathrm{NH_4^+}}^{\mathrm{BGW}}/\bar{\Phi}_{\mathrm{NH_4^+}}^{\mathrm{IGW}}(0.71)$  and  $a_{\mathrm{CDOM}}(\lambda)^{\mathrm{BGW}}/a_{\mathrm{CDOM}}(\lambda)^{\mathrm{IGW}}$  $(\ge 69)$ , the upper limit of  $\tau_{DON}^{BGW}/\tau_{DON}^{IGW}$  is estimated to be 0.32 (i.e. 31\*0.71/69). The photochemical turnover of DON<sup>BGW</sup> with respect to photoammonification is thus more than three times faster than that of DONIGW. Likewise, BGW is a bigger photochemical source of NH<sub>4</sub><sup>+</sup> than IGW, despite  $\Phi_{\text{NH}_4^+}^{\text{IGW}}$  being higher than  $\Phi_{\text{NH}_4^+}^{\text{BGW}}$ .

This preliminary assessment demonstrates that the photochemical turnovers of CDOM absorbance, DOC, and DON for DOM<sup>BGW</sup> are all substantially quicker than those for DOM<sup>BGW</sup> due partly (photobleaching and photomineralization) or entirely (photoammonification) to the very large ratio of  $a_{\rm CDOM}(\lambda)^{\rm BGW}/a_{\rm CDOM}(\lambda)^{\rm IGW}$  making DOM<sup>BGW</sup> the dominant photon absorber. A key assumption underlying this assessment is that the change in DOM photoreactivity during seaward transport is small compared with the difference in the original photoreactivity between the two DOM pools or that their

photoreactivities are affected to similar extents. Transport-derived photoreactivity alterations may arise from biogeochemical processes occurring in the aquifers and from mixing of groundwater with seawater. Given the narrow discharge zone and the short groundwater residence time in the sampling area (section 2.1.), the biogeochemical effect may be limited. Increases in salinity and pH during physical mixing may also alter the DOM photoreactivity but few relevant data are available in the literature. Reported effects of salinity on CDOM photobleaching in surface freshwater range from negligible (Song et al., 2016) to ~10% (Minor et al., 2006) and are ~30-40% on DOC photomineralization (Minor et al., 2006). These values are small relative to the differences in CDOM absorption and AQYs of photobleaching and photomineralization between BGW and IGW. Effects of salinity and pH on photoammonification are unknown.

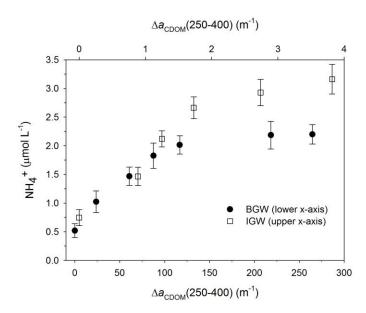


Figure 18. Ammonium concentration vs. loss of absorption coefficient averaged over 250-400 nm during time-course irradiations. Error bars represent absolute differences of duplicate samples.

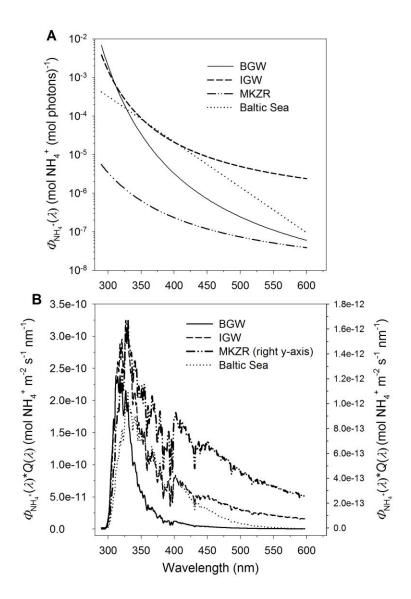


Figure 19. Apparent quantum yield spectra of ammonium photoproduction (A) and their corresponding action spectra (B). For comparison, data from Xie et al. (2012, station 697, salinity 0.15) and Stedmon et al. (2007, station Arkona, salinity 7.7), are included.

## **Chapter 4. Conclusions**

This study assessed the photoreactivities of two groundwater DOM pools with contrasting chemical and optical properties. Absorption and fluorescence spectroscopy revealed that BGW predominantly comprised strongly colored, old, high-MW DOM of terrestrial origin while IGW was largely composed of less colored, fresh, low-MW DOM of microbial origin. For both BGW and IGW, exposure to solar-simulated radiation led to increases in E<sub>2</sub>/E<sub>3</sub>, BIX, and NH<sub>4</sub><sup>+</sup> and decreases in DOC, MW, absorbance, SUVα<sub>254</sub>, FI, HIX, humic-like EEM peaks A and C, and humic-like PARAFAC C1-C4. On a fractional-absorption-loss basis, the photochemically induced fractional changes in BGW were comparable to those in IGB for SUVα<sub>245</sub>, E<sub>2</sub>/E<sub>3</sub>, and C2-C4, considerably larger for MW, BIX, and DOC, and substantially smaller for FIX, HIX, and C1. On an absorbedphoton basis, DOM<sup>BGW</sup> was far more, moderately more, and less photoreactive than DOMIGW in terms of absorbance photobleaching, DOC photomineralization, and photoammonification, respectively. The differing AQYs of these three major photoprocesses conform to the distinct chemical compositions of the two DOM pools elucidated by the spectroscopy. Photochemical turnovers of CDOM absorption, DOC, and DON in the receiving water body is likely much faster for DOMBGW than for DOM<sup>IGW</sup> due to the much higher UV and visible absorption of DOM<sup>BGW</sup>.

As DOM in groundwater experiences little prior light exposure, our results are instrumental for characterizing photoreactivities of different organic source materials in natural waters, particularly those of terrigenous and microbial origins. Generalization of the findings from the present study and those reported previously for surface waters leads to a tentative paradigm: terrigenous DOM is more prone to absorbance photobleaching

but less prone to photoammonification than microbial-derived DOM while the two DOM pools are comparably reactive with respect to DOC photomineralization. The mechanisms behind this pattern are poorly understood and warrant further investigations. Note that this paradigm is based on AQYs not the absolute rates of the photoprocesses in question. In freshwater systems (except large lakes), estuaries, and coastal zones strongly impacted by river runoff, terrestrial CDOM is usually the dominant absorber of the photochemistry-driving UV and visible radiation. In these aquatic environments, the rate of NH<sub>4</sub> photoproduction from terrigenous DOM can be higher than that from autochthonous DOM of microbial origin despite the lower AQY for the former.

## **Conclusions**

Cette étude a évalué les photoréactivités de deux échantillons DOM d'un aquifère présentant des propriétés chimiques et optiques contrastées. Les spectroscopies d'absorption et de fluorescence ont révélé que la DOM BGW comprenait principalement des composés d'origine terrestre fortement colorés, anciens et de haut poids MW tandis que la DOM IGW était en grande partie composé de matière d'origine microbienne moins colorée, plus fraîche et de faible MW. Pour BGW et IGW, l'exposition au rayonnement solaire simulé a entraîné des augmentations de E2/E3, de l'indice BIX et de la concentration en NH<sub>4</sub><sup>+</sup> et des diminutions de la teneur en DOC, des MW, de l'absorbance, du SUVα254, des indices FI et HIX, des intensités des pics A et C des matrices d'excitation-émission de fluorescence et des constituants PARAFAC C1-C4, de type humique. Sur la base de la perte relative d'absorption, suite à l'irradiation, les changements photochimiques induits dans BGW étaient comparables à ceux dans IGB pour SUVα245, E2/E3 et les constituants PARAFAC C2-C4. De plus, ils étaient considérablement plus grands pour MW, BIX et DOC et sensiblement plus petits pour FIX, HIX et le constituant PARAFAC C1. Sur la base de photons absorbés, DOM BGW était beaucoup plus, modérément plus et moins photoréactif que DOM IGW en termes de photoblanchiment de l'absorbance, de photominéralisation du DOC et de la photoammonification, respectivement. Les différents rendements quantiques apparents, AQY, de ces trois photoprocédés majeurs sont conformes aux compositions chimiques distinctes des deux pools DOM élucidés par la spectroscopie. Les temps de transformation photochimique de l'absorption de la matière organique colorée, CDOM, du DOC et de l'azote organique dissous, DON, dans les eaux côtières sont

vraisemblablement beaucoup plus rapides pour DOM BGW que pour DOM IGW. Cela est en grande partie ou entièrement en raison de l'absorption, dans l'UV et le visible, beaucoup plus importante pour DOM BGW. La présence d'un paléosol (la principale source présumée de DOM BGW dans cette étude) sur la côte est de l'Amérique du Nord (Grant, 1970; Gehrels, 1994) et l'enfouissement prévu des aquifères côtiers actuels en raison de l'élévation du niveau de la mer (Post al., 2013) appellent à davantage d'études sur les impacts biogéochimiques potentiels de la décharge des eaux souterraines (submarine groundwater discharge, SGD), y compris ceux qui résultent de la photochimie des DOM, dans les eaux côtières. Comme les eaux souterraines contenant les DOM sont généralement peu exposées à la lumière, nos résultats sont essentiels pour caractériser les photoréactivités de différentes matières organiques dans les eaux naturelles, en particulier celles d'origines terrigène et microbienne. La généralisation des résultats de la présente étude, combiné à ceux rapportés précédemment pour les eaux de surface, conduit à un paradigme provisoire: le DOM terrigène est plus sujet au photoblanchiment de l'absorbance mais moins sujet à la photoammonification que la DOM d'origine microbienne tandis que les deux DOM présentent une réactivité comparable pour la photominéralisation du DOC. Les mécanismes sous-jacents à ce modèle sont mal compris et justifient des recherches plus approfondies. Notez que ce paradigme est basé sur les AQY et non sur les taux absolus des photoprocédés en question. Dans les systèmes d'eau douce (sauf les grands lacs), les estuaires et les zones côtières fortement touchés par le ruissellement des rivières, le CDOM terrestre est habituellement l'absorbeur dominant du rayonnement ultraviolet menant à des transformations photochimiques. Dans ces milieux aquatiques, le taux de photoproduction de NH<sub>4</sub><sup>+</sup> des DOM terrigènes peut être supérieur à celui des DOM autochtones d'origine microbienne, malgré l'AQY plus faible pour les premiers.

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# Appendix

Statistics of regression analysis for figures 10, 11, 13, 14, and 16 in the main text

Figure	Sample label	Equation	a±SE	b±SE	c±SE	$\mathbb{R}^2$	p	n
5	BGW (upper part)	Y=a*X+b	-0.0039±0.00008	$0.97\pm0.01$	N/A	0.997	< 0.0001	8
	BGW (lower part)	Y=a*exp(-b*X)	$0.82\pm0.13$	$0.0075 \pm 0.0006$	N/A	0.991	0.004	4
	IGW (upper part)	Y=a*exp(-b*X)	$0.97 \pm 0.02$	$0.011\pm0.001$	N/A	0.990	< 0.0001	6
	IGW (upper part)	Y=a*X+b	-0.0077±0.0005	$0.95\pm0.02$	N/A	0.982	0.0001	6
	IGW (lower part)	Y=a*exp(-b*X)	$0.52\pm0.002$	$0.0026 \pm 0.00005$	N/A	1.000	0.01	3
	IGW (lower part)	Y=a*X+b	-0.0010±0.00004	$0.50\pm0.0045$	N/A	0.998	0.03	3
	BGW (panel A)	Y=a*exp(b/(X+c))	$0.891 \pm 0.037$	$0.188 \pm 0.016$	$0.077 \pm 0.008$	0.997	< 0.0001	11
6	IGW (panel A)	Y=a*exp(-b*X)+c	$1.66\pm1.07$	3.69±2.16	$0.91\pm0.09$	0.929	0.001	8
	BGW (panel B)	$Y=a*X+b*X^2+c$	$1.69\pm0.06$	$-0.81\pm0.05$	$0.10\pm0.01$	0.998	< 0.0001	11
	IGW (panel B)	$Y=a*X+b*X^2+c$	$2.72\pm0.35$	-1.47±0.26	$-0.25\pm0.10$	0.989	< 0.0001	8
	BGW (panel C)	Y=a*X+b	$0.40\pm0.02$	$0.56\pm0.01$	N/A	0.976	< 0.0001	11
	IGW (panel C)	Y=a*X+b	$0.066 \pm 0.005$	$0.93\pm0.004$	N/A	0.962	< 0.0001	8
	BGW (panel D)	Y=a*X+b	$0.265 \pm 0.013$	$0.714 \pm 0.008$	N/A	0.979	< 0.0001	11
	IGW (panel D)	Y=a*X+b	$0.043 \pm 0.002$	$0.956 \pm 0.002$	N/A	0.983	< 0.0001	8
8	BGW (panel A)	Y=a*X+b	0.057±0.004	0.94±0.003	N/A	0.972	< 0.0001	8
	IGW (panel A)	Y=a*exp(b*X)+c	$(3.0\pm4.2)*10^{-6}$	10.6±1.4	$0.88 \pm 0.004$	0.986	0.0002	7
	IGW (pane B)	Y=a*X+b	$-0.47\pm0.02$	$1.47\pm0.02$	N/A	0.988	< 0.0001	7
	IGW (panel C)	Y=a*exp(b*X)+c	$0.0082 \pm 0.0069$	$4.68\pm0.83$	$0.13\pm0.05$	0.985	0.0002	7
9	IGW (panel A)	Y=a*exp(b*X)	0.0035±0.0013	5.66±0.40	N/A	0.993	< 0.0001	7
	IGW (panel B)	Y=a*X+b	-0.31±0.055	$1.24\pm0.081$	N/A	0.979	< 0.0001	7
	IGW (panel C)	Y=a*X+b	$-0.39\pm0.078$	$1.28\pm0.11$	N/A	0.962	0.0001	7
	IGW (panel D)	Y=a*X+b	$-0.28\pm0.042$	$1.28\pm0.062$	N/A	0.988	< 0.0001	7
	IGW (panel E)	Y=a*X+b	$0.37\pm0.014$	$0.62\pm0.019$	N/A	0.996	< 0.0001	7
10	BGW (panel B)	Y=a*X+b	0.33±0.005	-0.0007±0.0023	N/A	0.999	< 0.0001	8
	IGW (panel B)	Y=a*X+b	$0.19\pm0.01$	$0.0036 \pm 0.0022$	N/A	0.992	0.0003	5
	SRW (panel B)	Y=a*X+b	$0.39\pm0.02$	-0.0076±0.0062	N/A	0.992	0.0003	5